



# A New Reference Instrument for Reflective Solar Inter-Calibration: from MODIS to VIIRS

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*(1) NASA/GSFC; (2) NOAA NESDIS; (3) NASA/LaRC*

*Acknowledgements: S. Wagner, and many other GSICS members*

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- **Motivations (why?)**
  - From MODIS to VIIRS
- **Calibration Approaches for MODIS and VIIRS (make sense?)**
- **Inter-Calibration and Reference Transfer Approaches (how?)**
  - For most spectral bands, VIIRS is more suitable as a reference sensor for inter-calibration
- **Future Considerations (what else?)**

# Motivations

- **Aqua MODIS has been (was) used as the reflective solar (RS) calibration reference for many years by many users, including GSICS community**
  - “Best” characterized and validated earth-observing sensor [at the time](#)
    - ✓ Extensive pre-launch calibration/characterization with lessons from Terra MODIS
    - ✓ Traceable to NIST reflectance standard
    - ✓ Stringent calibration requirements ( $\pm 2\%$  in reflectance)
    - ✓ A set of on-board calibrators, including SD, SDSM, and SRCA
    - ✓ Scheduled lunar observations through the entire mission
    - ✓ Better performance than Terra MODIS in the RS spectral region
  - Successfully operated [since 2002](#) with mission likely to be extended beyond 2023 (NASA HQ Senior Review, 2017)
  - Dedicated efforts for instrument operation and calibration (MCST)
  - Consistently produced long-term data records enabling a broad range of applications
- **Many new missions/sensors have been launched and operated by different countries/agencies in recently years, including VIIRS, OLI, AHI, ABI, and sensors on Sentinels => need for a new reference instrument**

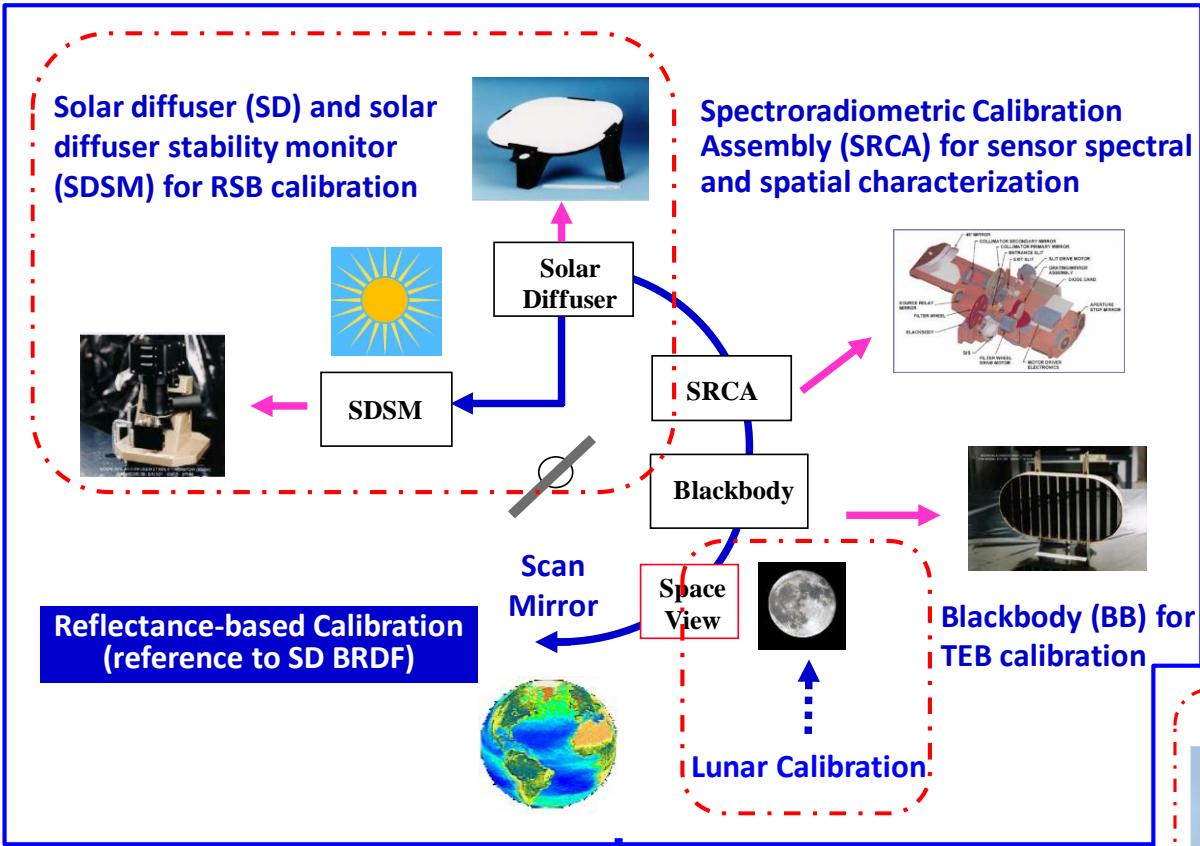
# Motivations

- **VIIRS can and should be used as a new RS calibration reference**
  - Designed with strong MODIS heritage by the same instrument vendor
  - Operated and calibrated based on lessons and strategies from MODIS
  - Applied similar solar and lunar calibration approaches and methodologies
  - S-NPP VIIRS, [launched in 2011](#), provides critical linkage between data records derived from EOS Terra/Aqua MODIS and future JPSS VIIRS sensors
    - JPSS-1 launch: Nov 15, 2017
    - J2/J3/J4 launch: 2021/2026/2031
  - Dedicated calibration effort by NOAA and NASA, including data reprocessing
    - NASA VIIRS SDR and L1B reprocessing; NOAA VIIRS SDR reprocessing
- **Issues to be addressed**
  - For previous sensors already used Aqua MODIS as inter-calibration reference
  - For future sensors likely to use J1 or even J2/J3/J4 VIIRS for inter-calibration
  - **Which VIIRS and what VIIRS data (from NOAA IDPS, NASA SIPS, ...)?**
  - **What about the Moon?**

# Calibration Approaches for MODIS and VIIRS

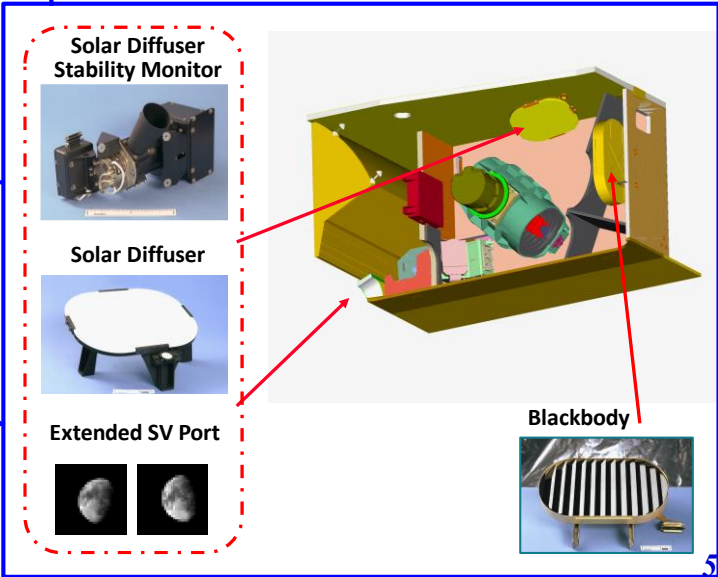
## Similar Calibration Methodologies

← MODIS  
 ↓ VIIRS



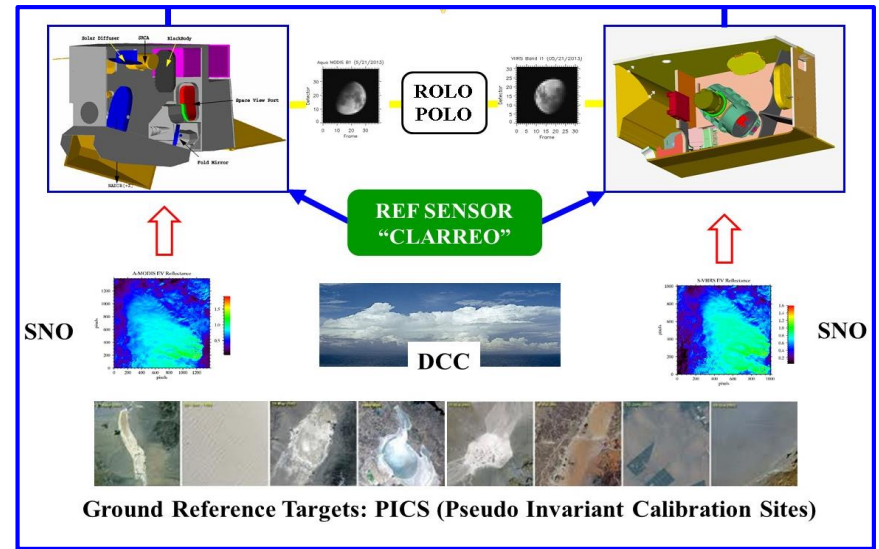
**VIIRS: strong MODIS heritage**

- Instrument Design (OBC)
- Operation
- Calibration



# Inter-Calibration and Reference Transfer Approaches

- Two most popular (if not the best) approaches to transfer calibration reference:
- DCC Method
  - MODIS: works for B1, B3-7, B18, and B26 (not for B2, B8-16, B17, and B19 due to saturation)
  - VIIRS: works for all bands except for M6 (roll over saturation)
- Lunar Calibration
  - MODIS: works for B1-B4, B8-B12, and B17-19\* (\*no matching VIIRS bands); relative approach for B13-16 due to saturation; xtalk correction needed for SWIR bands
  - VIIRS: works for all bands



VIIRS Band	Spectral Range (um)	Nadir HSR (m)	MODIS Band(s)	Range	HSR
DNB	0.500 - 0.900				
M1	0.402 - 0.422	750	8	0.405 - 0.420	1000
M2	0.436 - 0.454	750	9	0.438 - 0.448	1000
M3	0.478 - 0.498	750	3 10	0.459 - 0.479 0.483 - 0.493	500 1000
M4	0.545 - 0.565	750	4 or 12	0.545 - 0.565 0.546 - 0.556	500 1000
I1	0.600 - 0.680	375	1	0.620 - 0.670	250
M5	0.662 - 0.682	750	13 or 14	0.662 - 0.672 0.673 - 0.683	1000 1000
M6	0.739 - 0.754	750	15	0.743 - 0.753	1000
I2	0.846 - 0.885	375	2	0.841 - 0.876	250
M7	0.846 - 0.885	750	16 or 2	0.862 - 0.877 0.841 - 0.876	1000 250
M8	1.230 - 1.250	750	5	SAME	500
M9	1.371 - 1.386	750	26	1.360 - 1.390	1000
I3	1.580 - 1.640	375	6	1.628 - 1.652	500
M10	1.580 - 1.640	750	6	1.628 - 1.652	500
M11	2.225 - 2.275	750	7	2.105 - 2.155	500

**VIIRS is more suitable as the new GSICS reference sensor**

# Future Considerations

## What should be done?

- For previous sensors already used Aqua MODIS as reference: calibration consistency between Aqua MODIS and S-NPP VIIRS
- For future sensors likely to use J1 or J2/J3/J4 VIIRS as reference: calibration consistency between S-NPP and future JPSS VIIRS
- For current sensors: directly use S-NPP VIIRS
- Ideally, a set of inter-calibration coefficients from Aqua MODIS to S-NPP VIIRS and from S-NPP VIIRS to future JPSS VIIRS can be made available to all users (independently validated and reviewed)

## Alternative approaches

- Transfer reference sensor (either MODIS or VIIRS) calibration to DCC and Moon (pros and cons)
- Future CLARREO or TRUTH type of instruments with significantly improved calibration traceability and accuracy, thus much reduced inter-calibration uncertainty

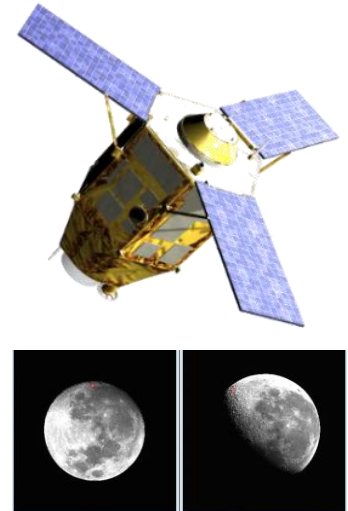
## Coordinated efforts



# What About the Moon?

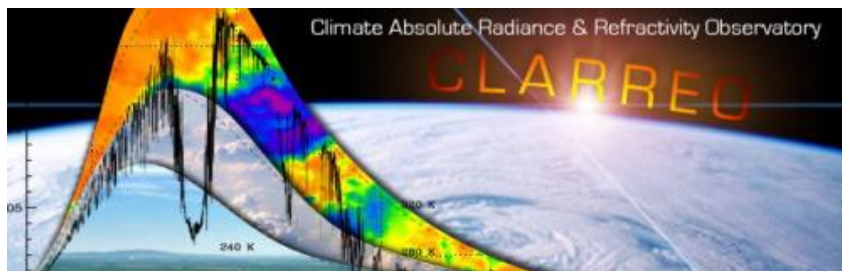
- **Presentations in this workshop**

- Lunar observations made and used by more and more sensors
- Easy access to the lunar models
  - ROLO
  - GIRO
  - SELENE/SP
  - POLO: Pleiades Orbital Lunar Observations (CNES)



- **Potential improvements of lunar model (irradiance) absolute accuracy**

- NASA effort (e.g. CLARREO)
- NIST effort
- USGS effort
- Other efforts



**NIST Lunar Irradiance Program**  
at the Whipple Observatory, Mt. Hopkins AZ  
Claire Cramer, Keith Lykke and John Woodward

The schematic shows a diagram of the measurement setup. It includes an "Integrating sphere source acts as an artificial moon", a "calibrated spectrometer", and an "Integrating sphere spectrometer". Arrows indicate the flow of light and data between these components.

Goal: a lunar irradiance measurement with  $k=2$  uncertainties of 1% or less over the spectral range from 400 nm to 1000 nm

Three photographs are shown at the bottom. The left photo shows the "MMT" (Mount Mitchell Telescope) on a mountain peak with a "calibration source location" marked. The middle photo shows the "NIST dome". The right photo is a topographic map showing the "NIST dome" location, with a "800 m" distance marked between the dome and the MMT.



**Questions?**