



IASI non-linearity correction : impact of the change on IASI-B

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GSICS web meeting

- I IASI-B tables change
- 2 IASI-B inter-comparison results
- 3 On-going studies
- 4 IASI-A status
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- New IASI-B non linearity correction tables were uploaded on 2nd August 2017 (PTSI 13).
- Expected differences wrt. the previous correction (new-old):

For a mean scene temperature of 263 K



Error higher at the beginning of B1 and scene temperature dependent



For different scene temperatures

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2 – Recall of the methodology for direct IASI-A / IASI-B

"Similar" scenes:

IASI-A and B are on the same orbit with a 180° shift

- Numerous common observations (CO) between 2 consecutive tracks, but:
 - » never simultaneous: ~50min temporal shift
 - » off-nadir: from 0° to 39°, opposite angles

Selection on the most relevant scenes

- Use of geoloc., geom., IIS, AVHRR, ECMWF data
- Focus on stable and homogeneous scenes
 - = Night, mostly oceans, 0% or 100% clouds
- Balance "A before B" and "A after B"

For each common observation

- Focus on the central area (same atmospheric thickness)
- Regional averaging of the soundings (300*300km)
- ΔT calculated at elementary channel level

$$\Delta T = \frac{\left(L_{IASI-B} - L_{IASI-A}\right)}{\sqrt{\frac{\partial L_{\sigma}}{\partial T}}(\sigma, 280K)}$$

Mean and stdev computed over the dataset



2 – IASI-A/IASI-B inter-comparison results

IASI-B – IASI-A
 Before the change of non
 linearity correction tables



After the change
Difference null ! Except in CO₂ and O₃ bands

2 – IASI-A/IASI-B inter-comparison results



2 – IASI-A/IASI-B - Scene temperature dependence

Classification of the bias with the scene temperature
 Plot of NedT vs BT (spectrally integrated in B1) + Sliding means
 Amplitude -0.1K (cold scenes) to 0 (warm scenes, close to the black body target temperature
 (293 K))

 Before the change of non linearity correction tables

- After the change
- Dependence wrt scene temperature reduced (curve flattened)
- Residuals maybe due to IASI-A non linearity



Mean

StdDev

cnes

2 – Recall of the methodology for IASI / AIRS, IASI / CRIS



Similar scenes: SNOs (Simultaneous Nadir Overpasses)

- Tolerance in simultaneity : 20 min
- ~30 scenes every 3 days for IASI / AIRS (12000 in 5 years)
- Always at high latitudes

• Spatial match:

→ Regional averaging of the soundings pixels over a 300km*300km area around the orbit crossing point

• Spectral match:

Construction of 33 broad pseudo-bands

- Each PB = intelligent averaging of ~100 elementary channels to get the similarity of the PB spectral functions
- The AIRS missing channels and varying spectral resolution are considered when calculating the IASI coefficients

NB: the convolution of IASI by the CRIS or AIRS SRFs has been performed but is still under exploitation

For each pseudo-band, $\Delta T = (L_{IASI} - L_{AIRS})$

 $\frac{\partial L_{\sigma}}{\partial T}(\sigma, 280K)$

2 – IASI-B/CrIS inter-calibration results

 Before the change of non linearity correction tables



- After the change
- Offset of ~ + 0.15 K
- Comparison is done on cold scenes (gaussian distribution around 245 K)
- Question: did IASI was used as a reference for CrIS non linearity correction ?

2 – IASI-B/CrIS inter-calibration results



2 – IASI-B/AIRS inter-calibration results

 Before the change of non linearity correction tables



- After the change
- Offset of ~ + 0.15 K
- Spectral signature close to IASI-A / IASI-B comparison
- Window channels closer to zero
- Comparison is done on cold scenes (gaussian distribution around 245 K)

2 – IASI-B/AIRS inter-calibration results



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 In order to document and understand all these comparisons, some assumptions are currently studied.

=> Understand why CO_2 and O_3 bands behavior is different from the rest of band 1 (already the case before the change).

IASI-A/IASI-B intercalibration results are on the night soundings (~21:30), with the number of soundings "IASI-A before IASI-B" equivalent with "IASI-A after IASI-B".

Effect of the time gap between IASI-A and IASI-B

In the case of the use of the soundings only with the case "IASI-A before IASI-B",

we had (before the change) this kind of bias: Negative bias : IASI-A spectra, measured first, is warmer than IASI-B spectra.

Surface channels: the surface is getting colder (start of the night).

In CO₂ channels (stratosphere) in band 1 and 3: opposite effect: stratosphere is getting warmer.





3 – On-going studies

In the case of the use of the soundings only with the case "IASI-A after IASI-B", we had this kind of bias Positive bias : almost the opposite than in the case "IASI-A before IASI-B", IASI-B spectra being acquired first.

- When we take into account an equal number of the 2 cases, we have the mean of these 2 curves.
- When we have a temporal difference of 50 minutes between the 2 IASI (on a 100 minutes orbit), the geophysical bias is canceled.



- But, in real life, the temporal difference between the 2 IASI is not exactly 50 minutes. When this delay increases, for example 55 minutes for "IASI-A before IASI-B", and 45 minutes for "IASI-A after IASI-B", the final inter-calibration curve shows the instrumental bias + the atmosphere variation in 55 45 = 10 minutes.
- During ~10 minutes, do the stratosphere temperature evolves enough to lead to a spectrum variation of ~0,15 K in CO₂ band beginning of band 1, and around 0,05/0,1 K in O₃ band ? Or can it be due to concentration variations of these gases (ozone concentration linked with UV) ?

3 – On-going studies

- We are currently working on the following axis:
- Study of the time delay of "IASI-A before IASI-B" and "IASI-A after IASI-B", that is evolving due to the MetOp-A orbit drift.
- Study of the scenes used for CO and SNO: lat/lon distribution and evolution of the stratosphere temperature in these areas, mean temperature of these scenes.
- > Study of IASI-A/IASI-B inter-calibration using massive means.

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 For IASI-A, the update of non linearity tables was proposed for February 2019 (6 months later than IASI-B)

The mean radiometric error between proposed new NL correction and the operational one, expressed in NedT @ 280 K, for all the pixels

For a mean scene temperature of 263 K

For different scene temperatures



Double difference: {IASI-B new - IASI-B ope} – {IASI-A new - IASI-A ope}



- The difference between the 2 corrections of IASI-B and IASI-A is around 0.1 K, but not exactly, it depends on the wavenumber and scene temperature.
- If we change IASI-A, the inter-calibration bias between IASI-A and IASI-B will decrease from ~0 to ~ 0.05 K.
 So, IASI-A change is still a question mark...





4 – IASI-A status



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- The theoretical study on non-linearity has been validated in flight with IASI-B observations: differences in accordance with the study
- The current results of inter-calibration between IASI-A and IASI-B are satisfactory in band 1; the CO₂ and O₃ bands have a different pattern that need to be deeper analysed and are supposed to be geophysical effects.
- The theoretical study on non-linearity encourages us to change the correction tables on IASI-A too.
- If we change IASI-A, the inter-calibration with IASI-B and CrIS will be worse (with its limitation, we have no absolute reference, what is the reality ?)
- The difference between new and operational correction tables for IASI-A is low
- We can wonder if it is worth to change IASI-A, as it is a reference since 11 years, and the current inter-calibration with other sounders is satisfactory.

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6 – ISSWG feedback

• These results have been presented at the last ISSWG in December 5th.

- ISSWG have created a "task force" to better analyse the change on IASI-B and try to understand the signatures seen in CO₂ and O₃ bands.
- The 1st meeting of this task force will be held in March 2018.
- An action has been raised to ask CNES to exchange informations on IASI non linearity with CrIS team.
- The update of IASI-A non linearity tables is postponned, and will be considered again after the conclusions of ISSWG task force.

