

GK2A AMI Lunar Calibration Results

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Geo-KOMSAT-2A (GK2A) AMI



Geo-KOMSAT-2A (GK2A)



- GK2A was launched 4 December 2018
- The operation and public release started 25 July 2019
- Location : 128.2°E
- 16 channels (VIS 4, NIR 2 and IR 10)
- Radiometric Calibration Targets
 - Solar Diffuser for VIS/NIR channels and Blackbody for IR channel calibration
- In-orbit solar cal. validation & Instrument degradation trending of AMI VIS/NIR Channels
 - Vicarious calibration using ocean, desert, water cloud, DCC with RTM
 - Ray Matching method (MODIS and VIIRS)
 - and **Lunar Calibration** using GIRO

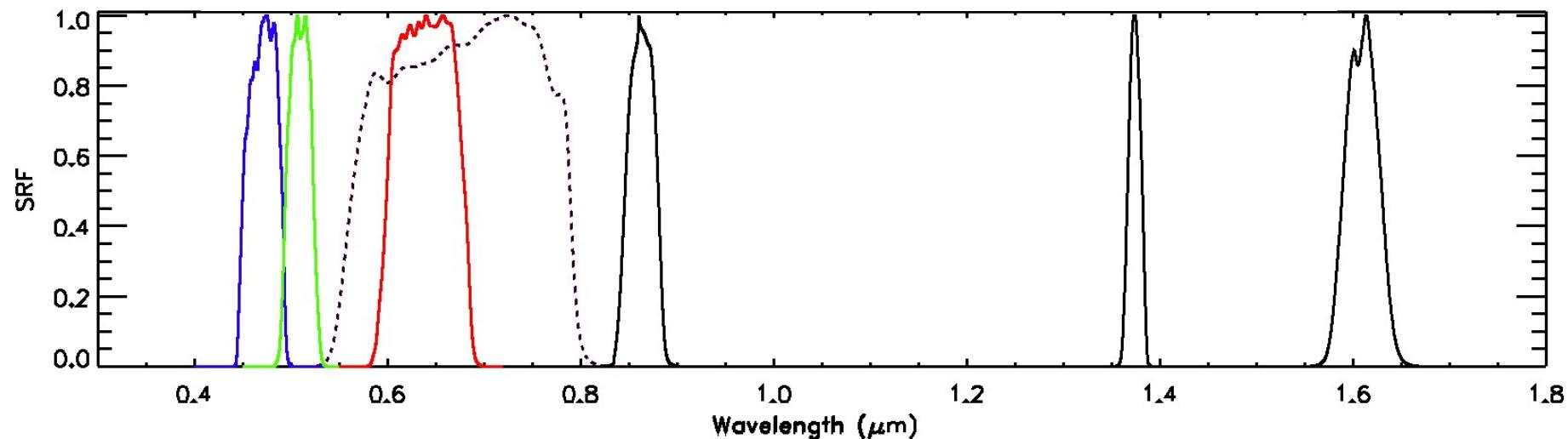
GK2A AMI VNIR Channel



◆ VNIR Channel Configuration

| | VI004 | VI005 | VI006 | VI008 | NR013 | NR016 |
|-------------------------|-------|-------|-------|-------|-------|-------|
| Central Wavelength [μm] | 0.47 | 0.51 | 0.64 | 0.86 | 1.3 | 1.6 |
| Spatial Resolution [km] | 1 | 1 | 0.5 | 1 | 2 | 2 |
| # of Detector | 632 | 632 | 1380 | 632 | 348 | 348 |

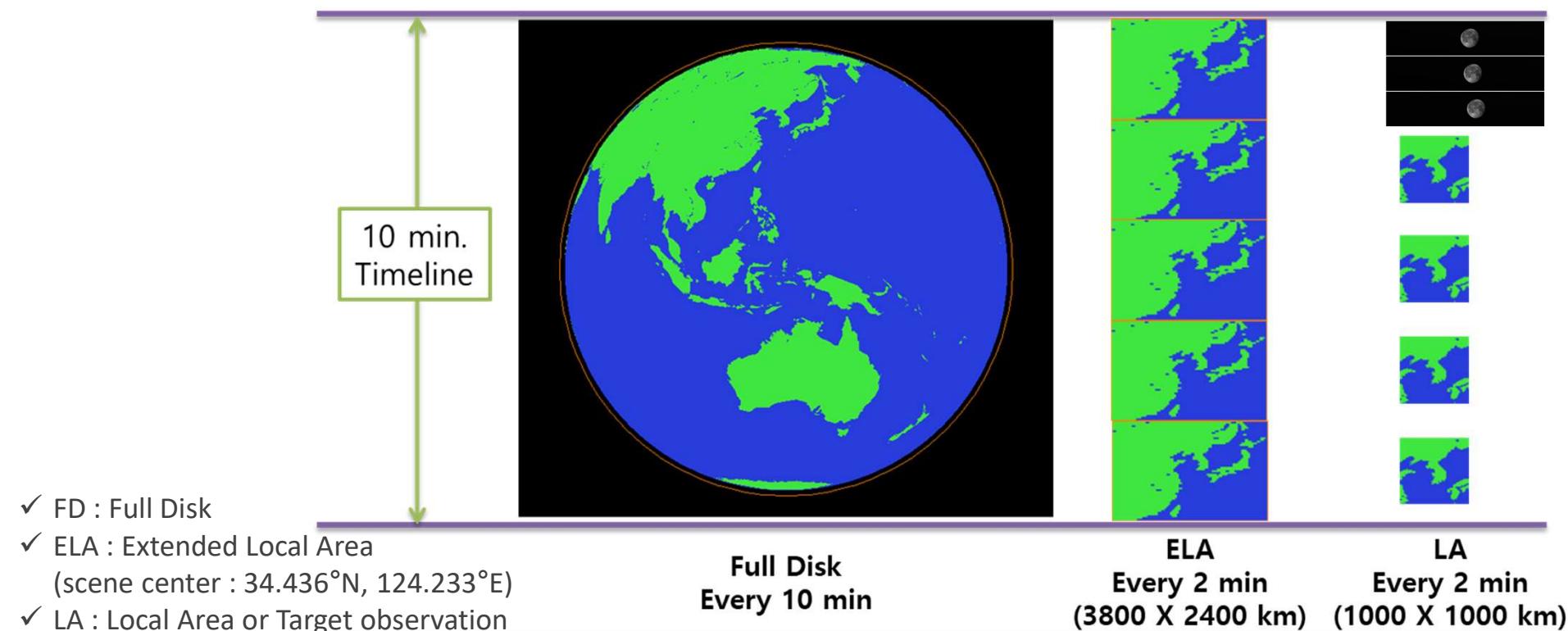
Spectral Response Functions of GK2A/AMI and COMS/MI (dashed line)



GK2A AMI Moon Observation



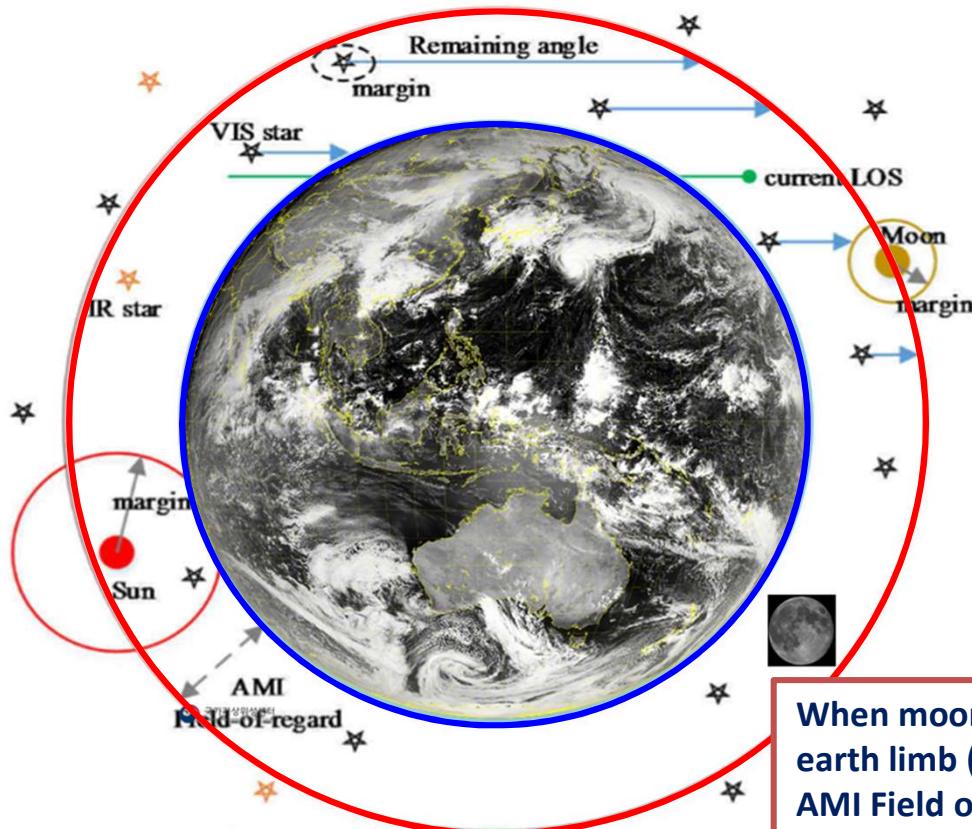
- ◆ Normal Operation : FD 1 + ELA 5 + LA 5 (10 min.)
- ◆ Lunar observation : FD 1 + ELA 5 + LA 4 + Moon 3 (10 min.)
 - LA 1 time (3 swath) → Moon 3 times



GK2A AMI Moon Observation



- ◆ GK2A AMI has observed the Moon about 40 times every month since May 2019

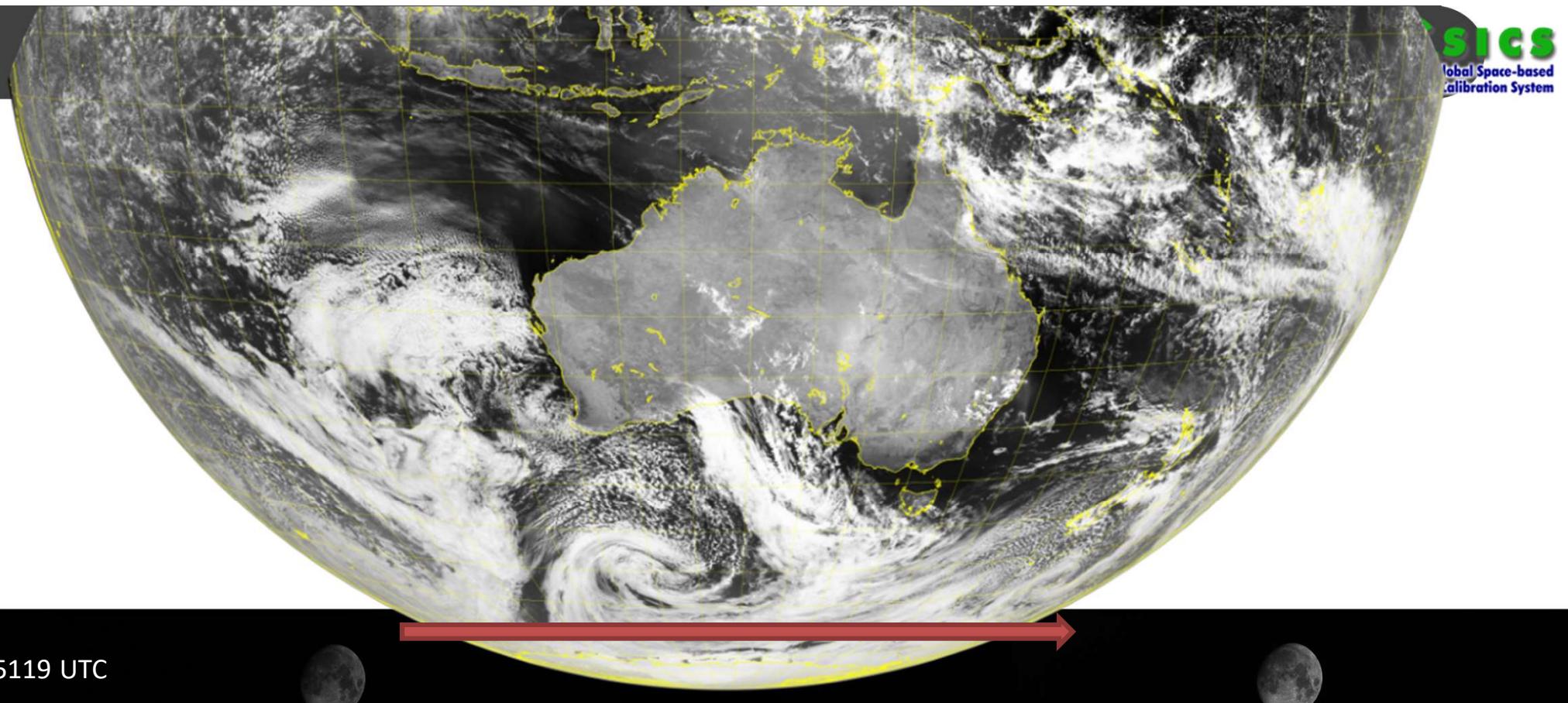


- AMI Moon observation condition
 - Moon Brightness > 50 (half moon)

| Period | No. of Moon Observation | No. of Moon Calibration Data set | Phase Angle Range | No. of Data Set |
|------------------------|-------------------------|----------------------------------|---------------------|-----------------|
| May 2019 ~ April. 2021 | 1132 | 936 | $ PA \leq 30$ | 319 |
| | | | $30 < PA \leq 60$ | 338 |
| | | | $60 < PA $ | 279 |

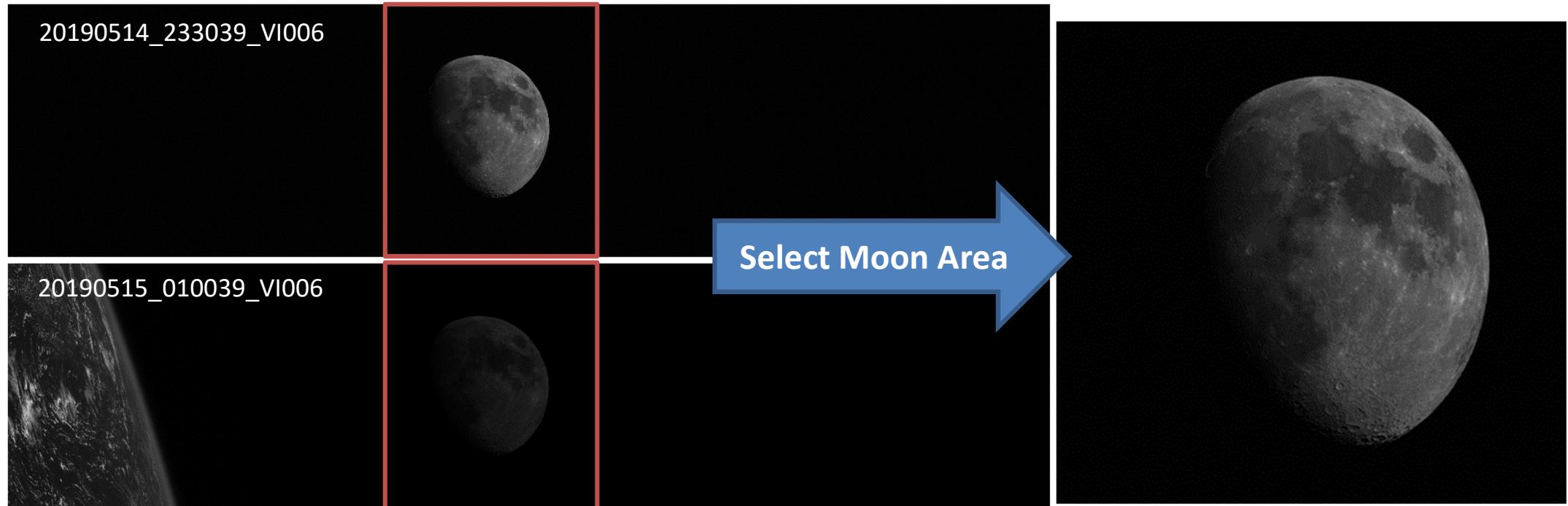
- Not use Moon data set when moon locate near earth or moon phase angle > $\pm 92^\circ$

When moon locate between
earth limb (blue circle) and
AMI Field of Regards(red circle)



06 Aug 2020
051039~055039 UTC

AMI Moon Data Set



| Channel | VI004 | VI005 | VI006 | VI008 | NR013 | NR016 |
|-------------------------------|---------|---------|-----------|---------|---------|---------|
| Moon data set (pixel size) | 600*632 | 600*632 | 1200*1380 | 600*632 | 300*348 | 300*348 |

✓ Phase Angle (absolute value) Range of observed Data set : min : 3.73 deg. / max : 91.82 deg.

GK2A AMI Lunar Irradiance



◆ Lunar Irradiance Calculation

$$I = \Omega \sum_i^{row} \sum_j^{col} Radiane_{i,j} \cdot \frac{1}{Oversampling\ Factor}$$

- Ω : Sample solid angle = sr(EW_ASD * NS_ASD)=EW_ASD*NS_ASD
- NS_ASD = NS IFOV
- $Radiane_{i,j}$: calibrated radiance at (i,j) image coordinate

| Channel Name | ASD (μrad) | IFOV (μrad) | | Solid Angle (EW_ASD*NS_ASD) (rad) |
|--------------|------------|-------------|------|-----------------------------------|
| | | NS | EW | |
| VI004 | 22 | 22.9 | 22.9 | 5.038E-10 |
| VI005 | 22 | 22.9 | 22.9 | 5.038E-10 |
| VI006 | 11 | 10.5 | 12.4 | 1.155E-10 |
| VI008 | 22 | 22.9 | 22.9 | 5.038E-10 |
| NR013 | 44 | 42 | 51.5 | 1.848E-09 |
| NR016 | 44 | 42 | 51.5 | 1.848E-09 |

GK2A AMI Oversampling Factor



◆ Calculate oversampling factor

- Using the NOAA Proposed Calculation Method
- ✓ Reference : GOES-16 ABI Lunar Data Preparation to GIRO, Fangfang Yu, 2nd Lunar Calibration Workshop, Xi'an, China, Nov. 13-17, 2017

| Channel Name | ASD (μrad) | IFOV (μrad) | | Solid Angle (EW_ASD*NS_ASD) (rad) | Oversampling Factors |
|--------------|------------|-------------|------|-----------------------------------|----------------------|
| | | NS | EW | | |
| VI004 | 22 | 22.9 | 22.9 | 5.038E-10 | 1.0056809602 |
| VI005 | 22 | 22.9 | 22.9 | 5.038E-10 | 1.0055607934 |
| VI006 | 11 | 10.5 | 12.4 | 1.155E-10 | 1.0028090287 |
| VI008 | 22 | 22.9 | 22.9 | 5.038E-10 | 1.0057331638 |
| NR013 | 44 | 42 | 51.5 | 1.848E-09 | 1.0049463375 |
| NR016 | 44 | 42 | 51.5 | 1.848E-09 | 1.0051906216 |

GK2A AMI Moon Results



◆ AMI Lunar Irradiance vs. GIRO

- Moon pixel select based on L1A DC thresholds for each channel
- Set the thresholds to remove bright pixel in space area
(e.g. star, or Contaminated pixel by Earth limb)

| | VI004 | VI005 | VI006 | VI008 | NR013 | NR016 |
|-----------------------------|-------|-------|-------|-------|-------|-------|
| Thresholds | 30 | 30 | 70 | 120 | 70 | 50 |
| mean counts of dc_offset | 19.8 | 19.8 | 39.7 | 79.6 | 39.7 | 19.5 |

- dc_offset : The digital count offset from the ith Moon observation of the channel k derived as the average value over the deep space portion of the image

GK2A AMI Lunar Results (May 2019. 5 ~ April 2021)



➤ Normalized Lunar Ratio

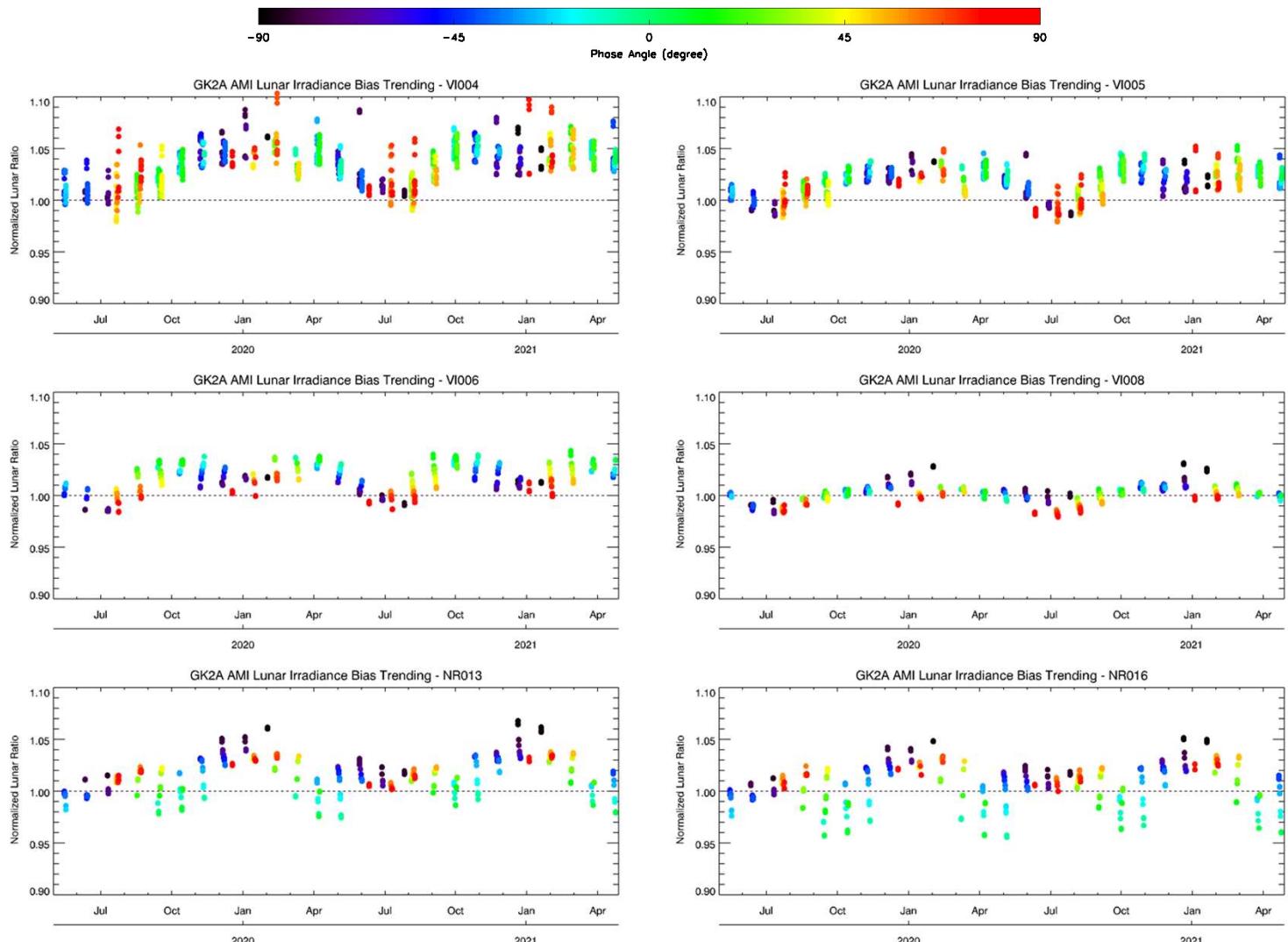
$$R_{nor} = \frac{Ratio_t}{Ratio_{t0}}$$

$$Ratio_t = \left(\frac{Irr_{obs,t}}{Irr_{GIRO,t}} - 1 \right) \cdot 100$$

Difference of lunar irradiance
between the observation and GIRO

✓ Note :

- Sub-sampled lunar observations (#936) are shown
- Ratio is normalized by $Ratio_{t0}$ (t0; 14 May 2019)



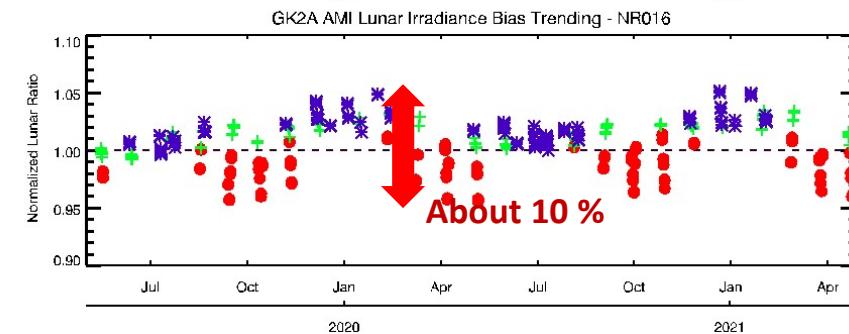
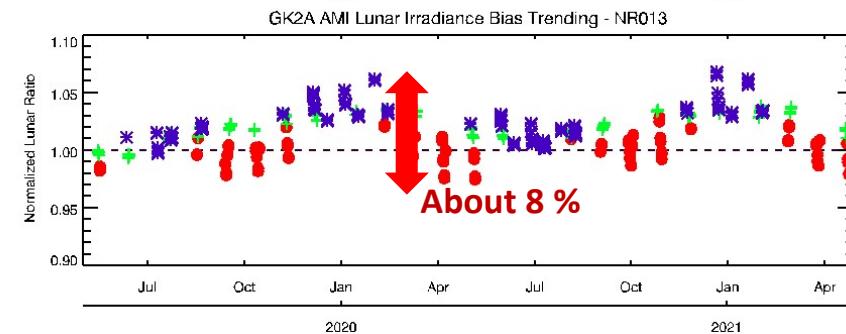
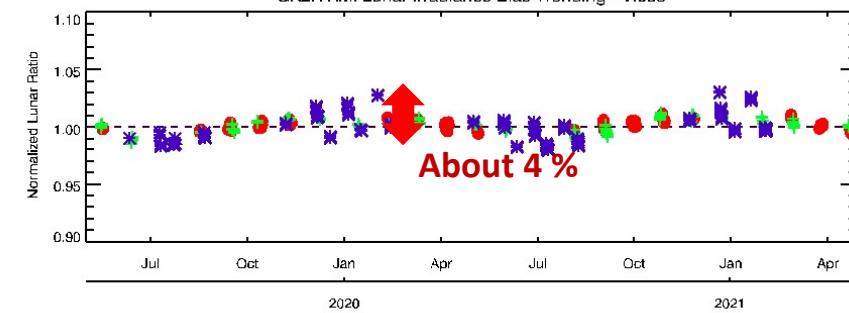
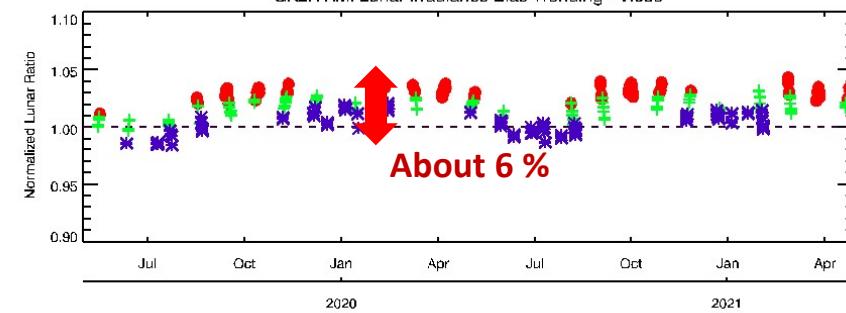
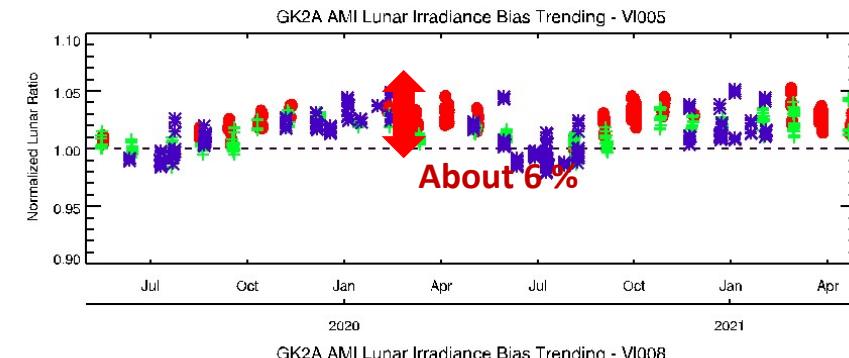
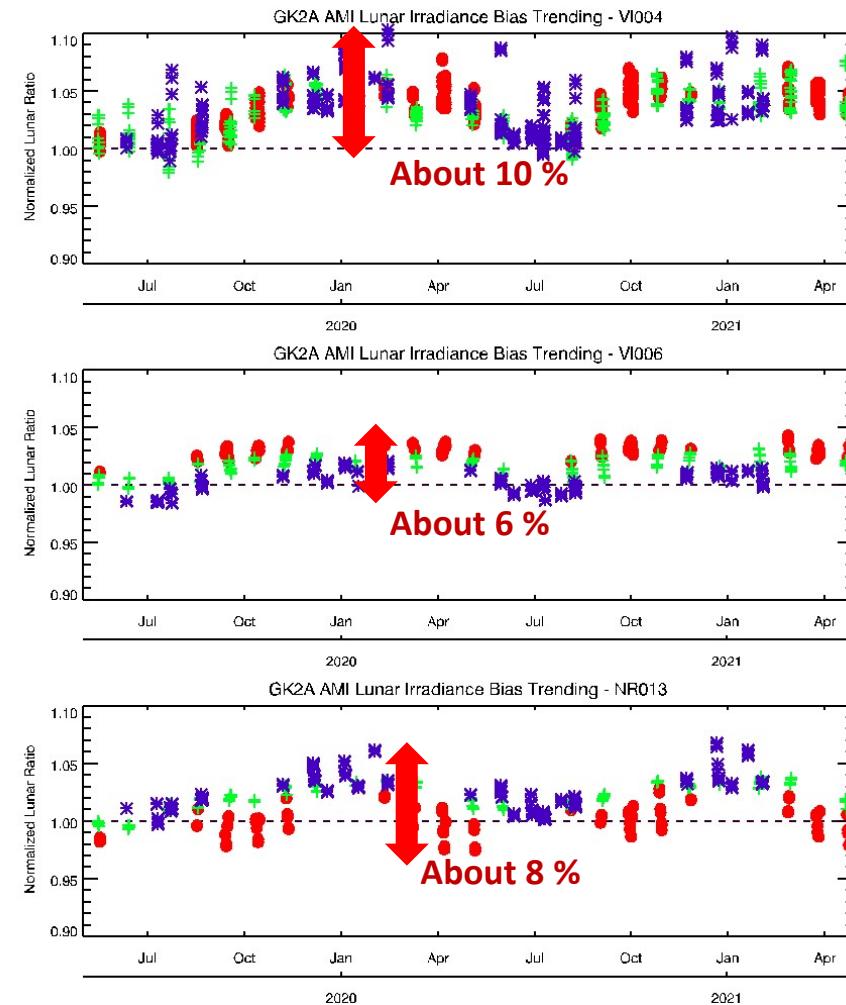
GK2A AMI Lunar Results (May 2019. 5 ~ April. 2021)



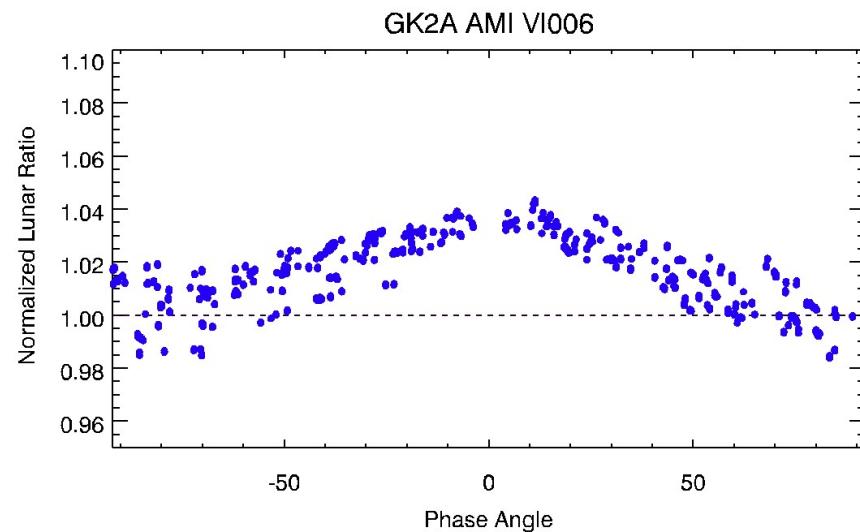
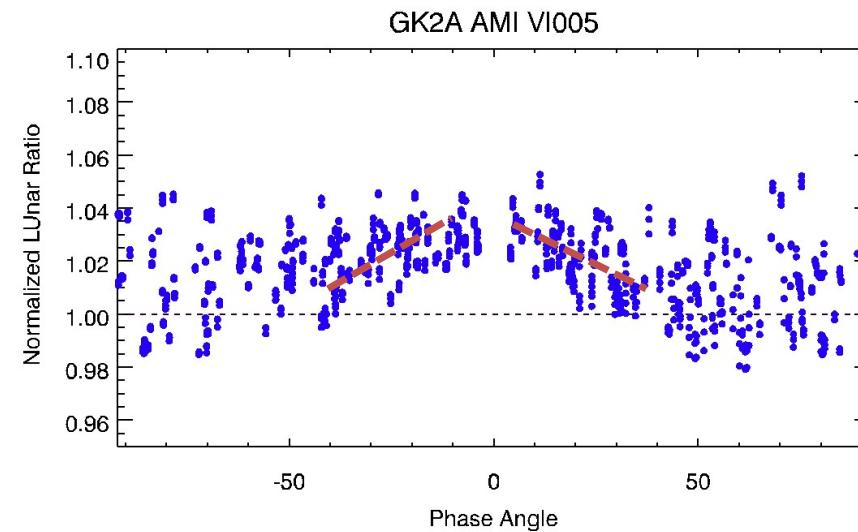
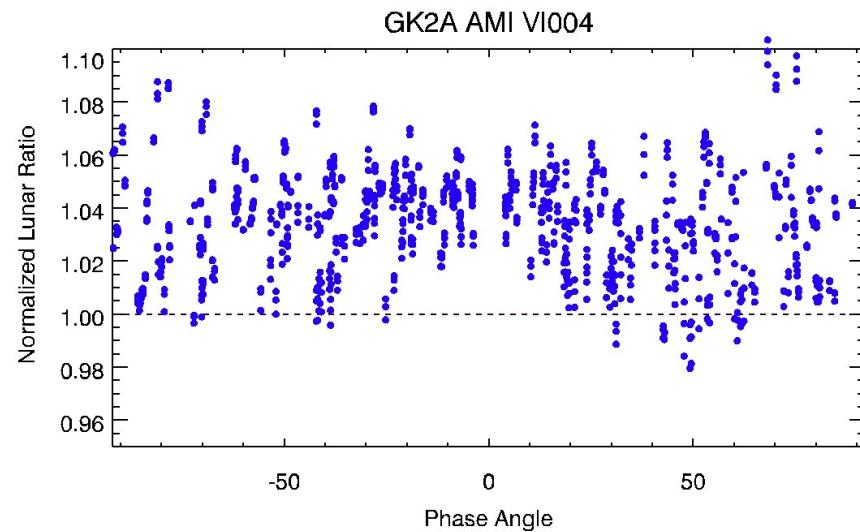
Red dot :
absolute PA $\leq 30^\circ$
(192 data set)

Green + sign :
 $30^\circ < \text{absolute PA} \leq 60^\circ$
(228 data set)

Purple Asterisk (*) :
absolute PA $> 60^\circ$
(224 data set)

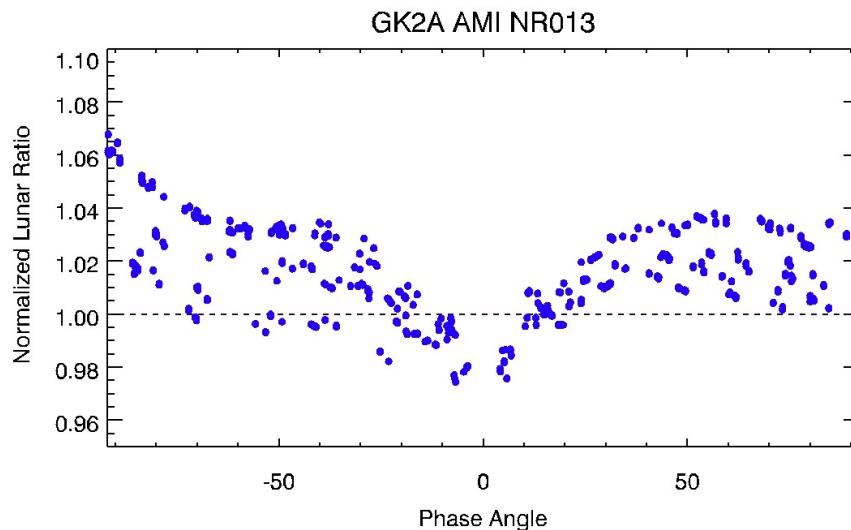
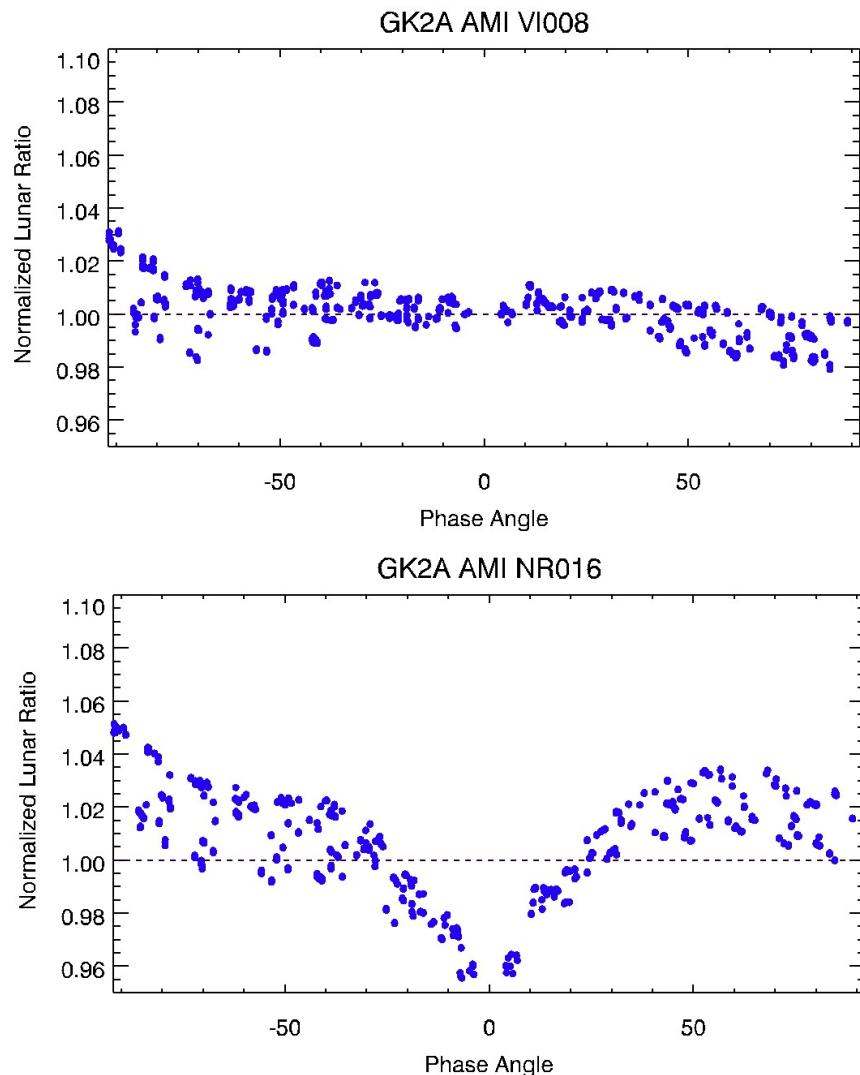


GK2A AMI Lunar Results (Ratio_{nor} vs. Phase Angle)



- ◆ Date : May 2019 – April 2021
 - VI004: Much scattered, no strong apparent phase angle dependent ratio
 - VI005: Phase angle dependent at relatively small absolute phase angles, but $> 40^\circ$ and $<-40^\circ$?
 - VI006: Phase angle dependent appears

GK2A AMI Lunar Results (Ratio_{nor} vs. Phase Angle)



- VI008 : Phase angle dependent bias appears
- NR013 and NR016 : Strong phase angle dependent calibration accuracy and Magnitude Increases with the wavelength

◆ GK2A AMI lunar observation

- 6 VNIR bands (0.47, 0.51, 0.64, 0.86, 1.3, 1.6 μm) and available since May 2019
- Lunar observation: the quality is under checking

◆ GK2A AMI Lunar calibration Results

- Show unexpected **seasonal variation about 5~10%** depend on channel (SRF)
 - Also VNIR DCC calibration (GSICS and Ray matching method) →**10 June 2021 webmeeting**
- **→ Need Long Term Monitoring and further investigation**
- Need revisit of oversampling factor and pixel IFoV
- Update moon pixel selection (e.g. use of new DC threshold and lunar shape fitting)

◆ Plan : Reprocess Lunar data

- Will re-process the lunar data without the updates of solar calibration coefficients to de-couple the SCT from the instrument performance

감사합니다.

