

# **Consideration of the On-board Calibration of Interferometric Synthetic Aperture Microwave Radiometer: *Using Geostationary Interferometric Microwave Sounder (GIMS) for Example***

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Chinese Academy of Sciences  
(MiRS, CAS)



# Outline

- **Synthetic aperture radiometer overview**
  - Real Aperture Vs. Synthetic Aperture
- **GIMS overview**
  - System concept
  - Demonstrator development
- **GIMS calibration consideration**
  - Overview of calibration for real aperture and synthetic aperture radiometer
  - Ground-based Calibration & Imaging Results
- **Conclusion**

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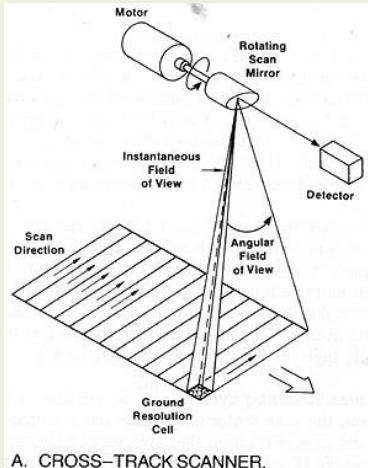
# Microwave radiometry: sounder & imager

## Sounders:

- AMSU-A/B, MHS, etc
- Cross-track scan
- 53GHz, 183GHz

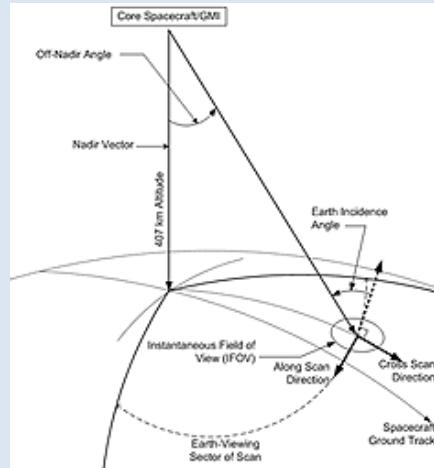


FY-3C/MWTHS



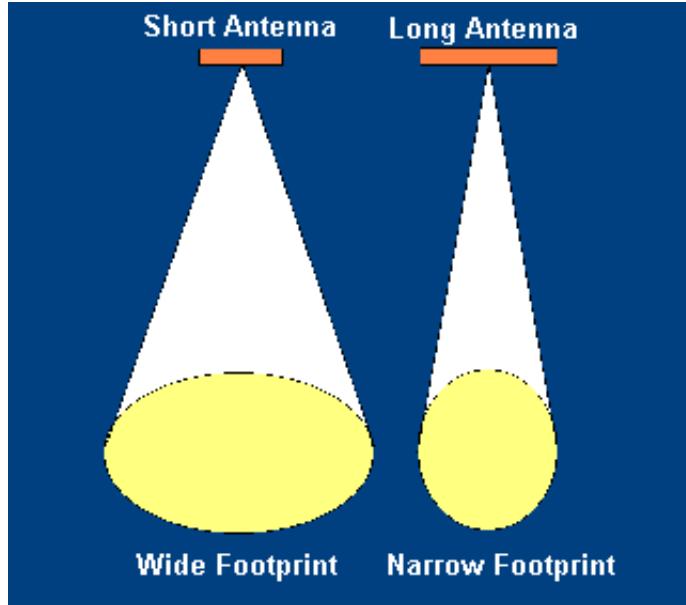
## Imagers:

- SSMI/S, TMI, GMI, AMSR2, etc
- Conical scan
- C/X band - 183GHz

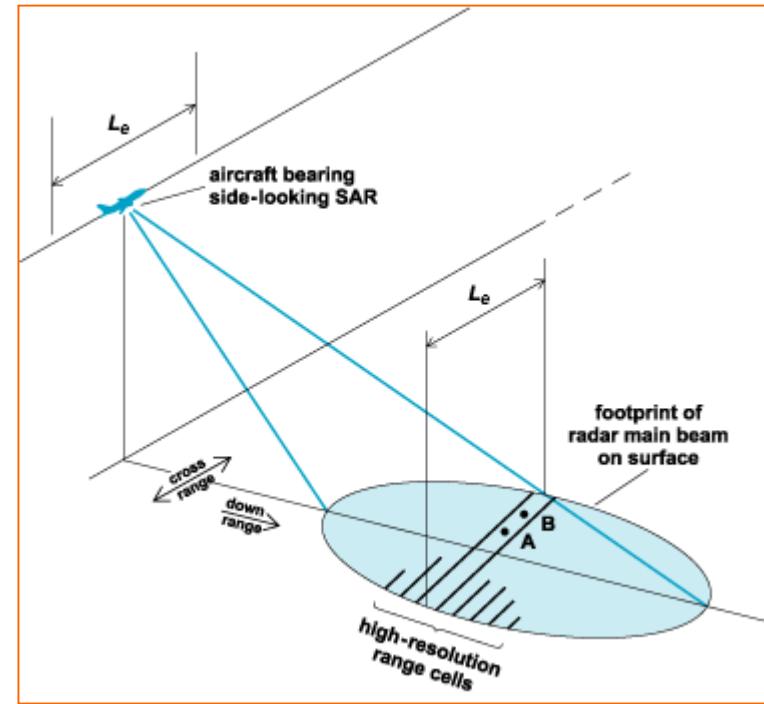


GMI

# The spatial resolution of passive microwave sensors



spatial resolution of **passive** sensors:  
totally determined by the antenna  
aperture size

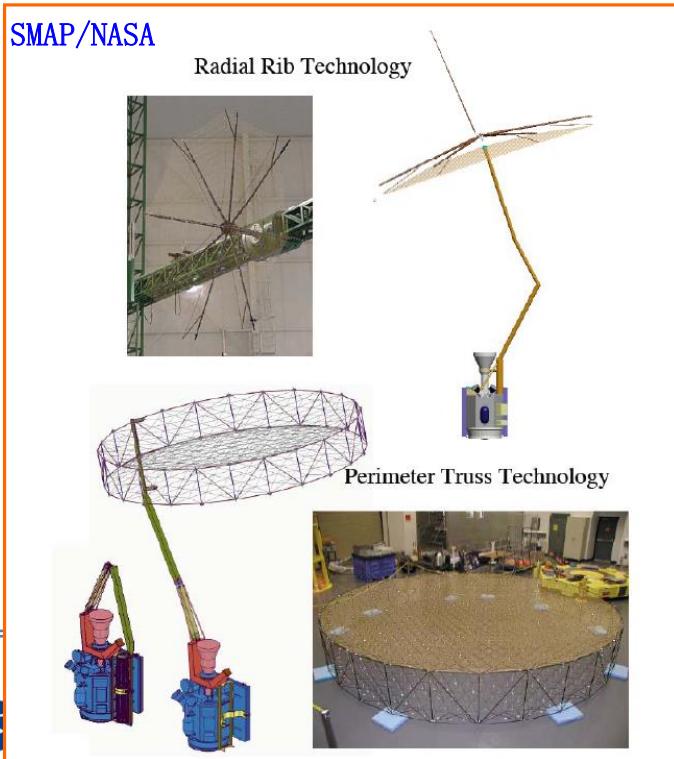


**Spatial resolution of active sensors:** can be improved by **range compression**, and **aperture synthesis**.

# How to improve the spatial resolution of the passive microwave sensors?

- Real aperture technologies:

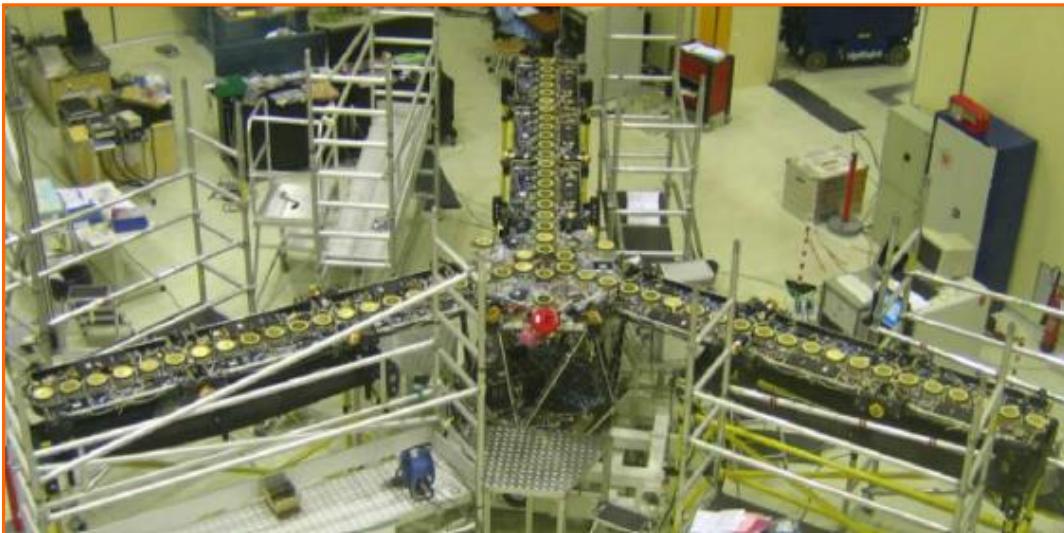
- large deployable antennas (low frequency)
- Large reflector antennas (high frequency)



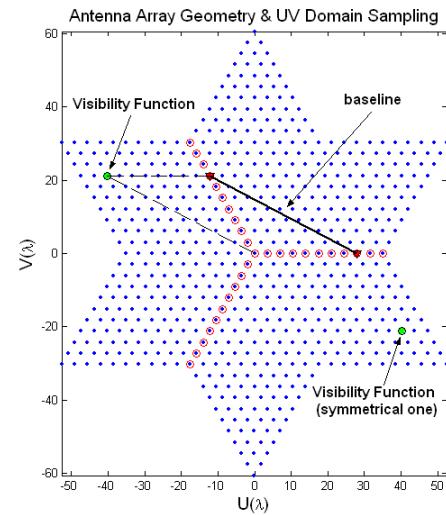
# Synthetic Aperture Radiometry

- Synthetic Aperture Radiometer (interferometric radiometer, synthetic thinned array radiometer):
  - Using **thinned array** to replace the **real aperture**
  - Measuring the Fourier transform of the brightness temperature distribution

MIRAS/SMOS/ESA, Y-shape thinned array



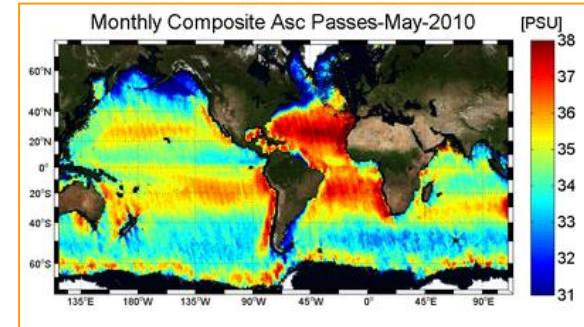
Equivalent Observation Aperture



# Applications of Interferometric Synthetic Aperture Radiometer

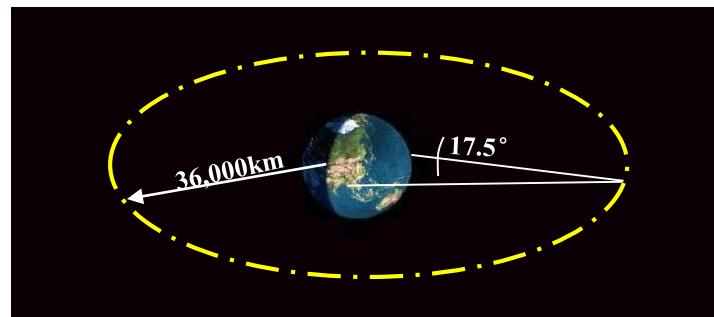
- **Soil Moisture & Ocean Salinity: L band**

- SMOS
- Chinese Salinity Mission
- WCOM



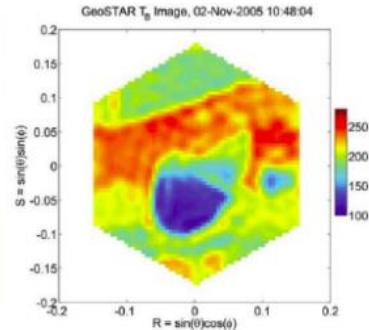
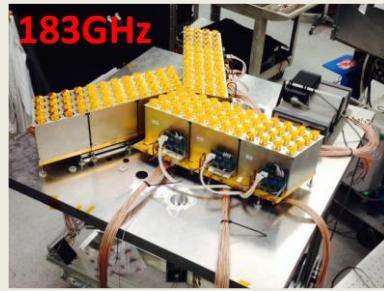
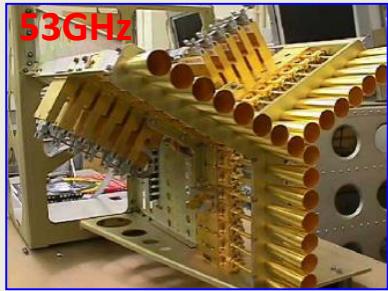
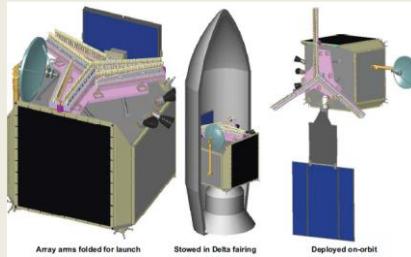
- **Geostationary Microwave Atmospheric Sounding: millimeter and sub-millimeter wave band**

- GeoSTAR, NASA/JPL
- GAS, ESA
- GIMS, CAS/NSSC



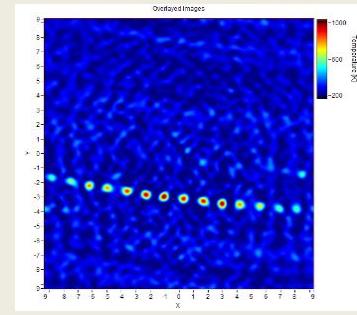
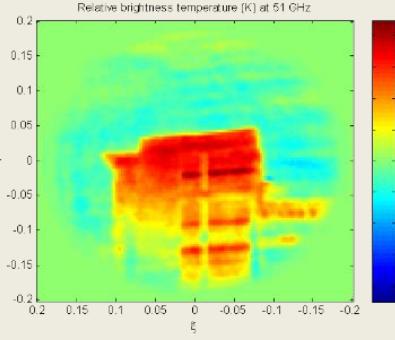
## Synthetic Aperture: Y-shape stationary array

**GeoSTAR (JPL/NASA)**  
**From 2000~**



## Synthetic Aperture: Rotating Y-shape array

**GAS (ESA)**  
**From 2002~**

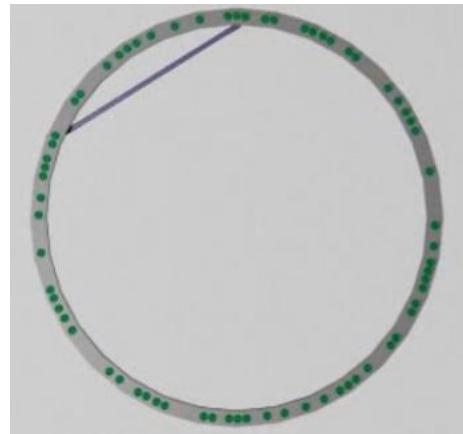
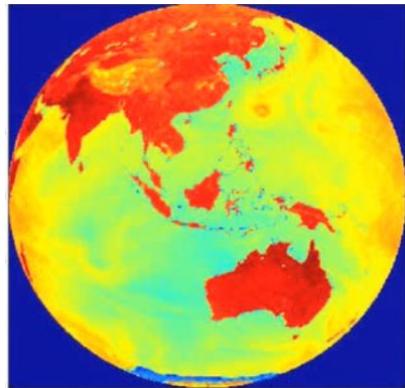


# Outline

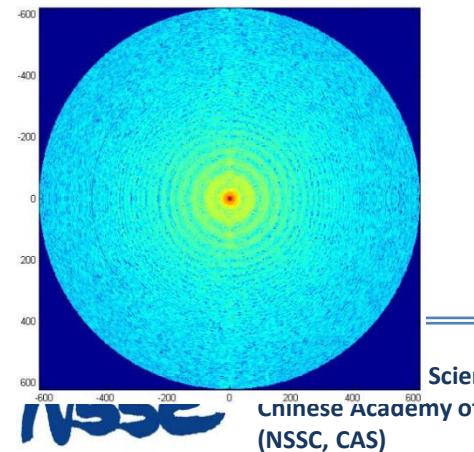
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# GIMS Concept

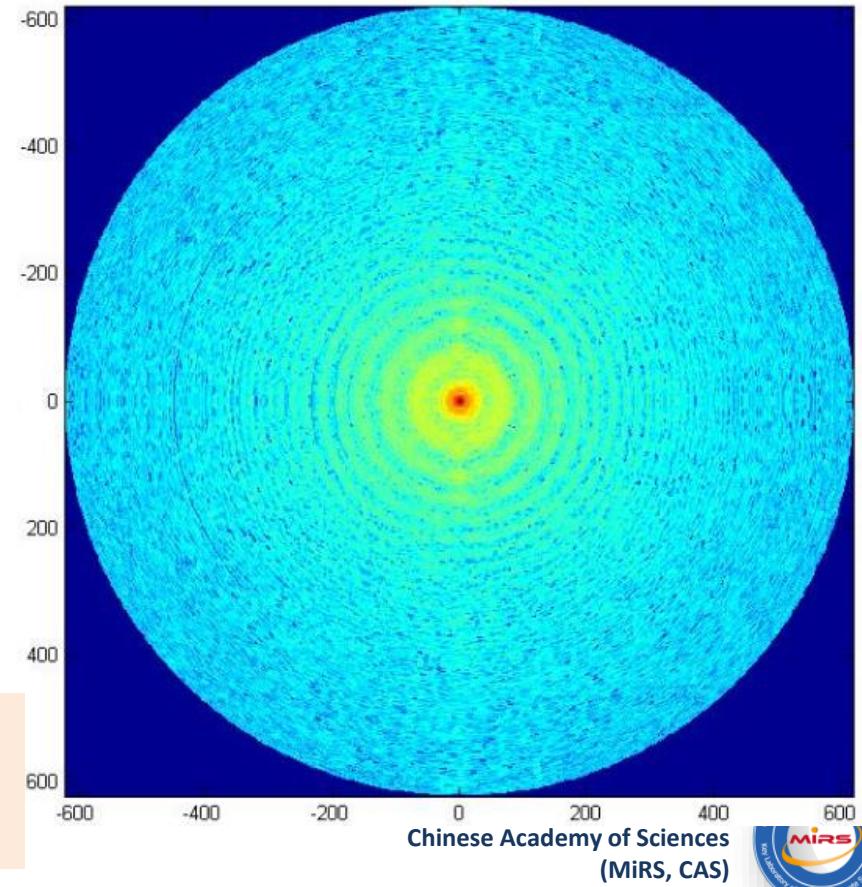
- GIMS ( Geostationary Interferometric Microwave Sounder )
  - Fourier Relationship: TB (Spatial Domain)  $\leftrightarrow$  VF (Spatial Frequency Domain)
  - Timeshared Sampling in SF Domain with Rotating Thinned Circular Array



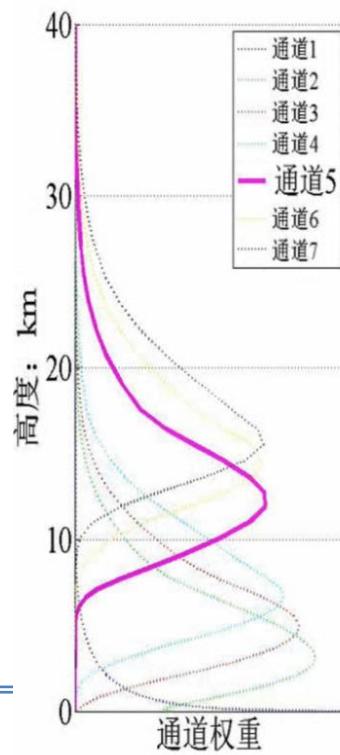
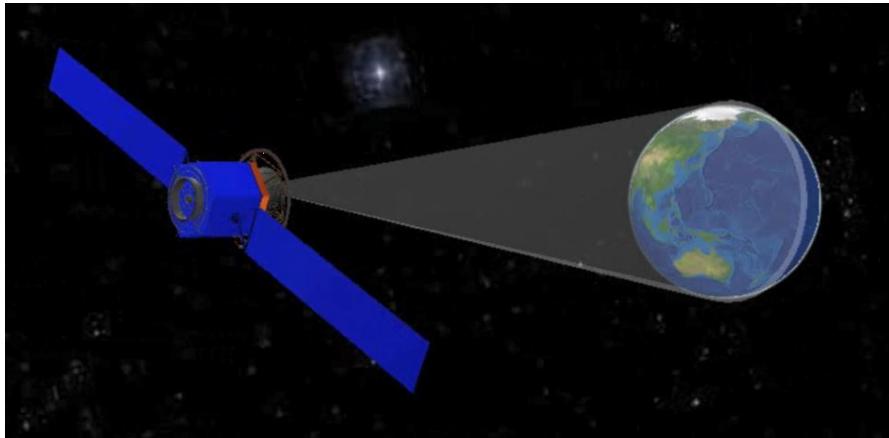
**Rotating circular thinned array**



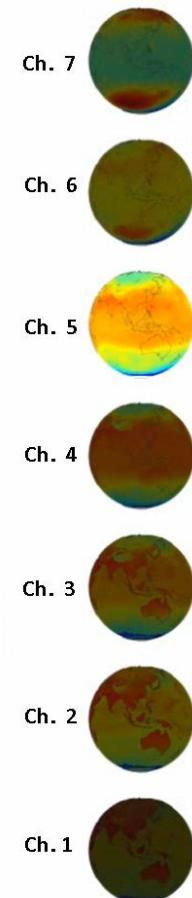
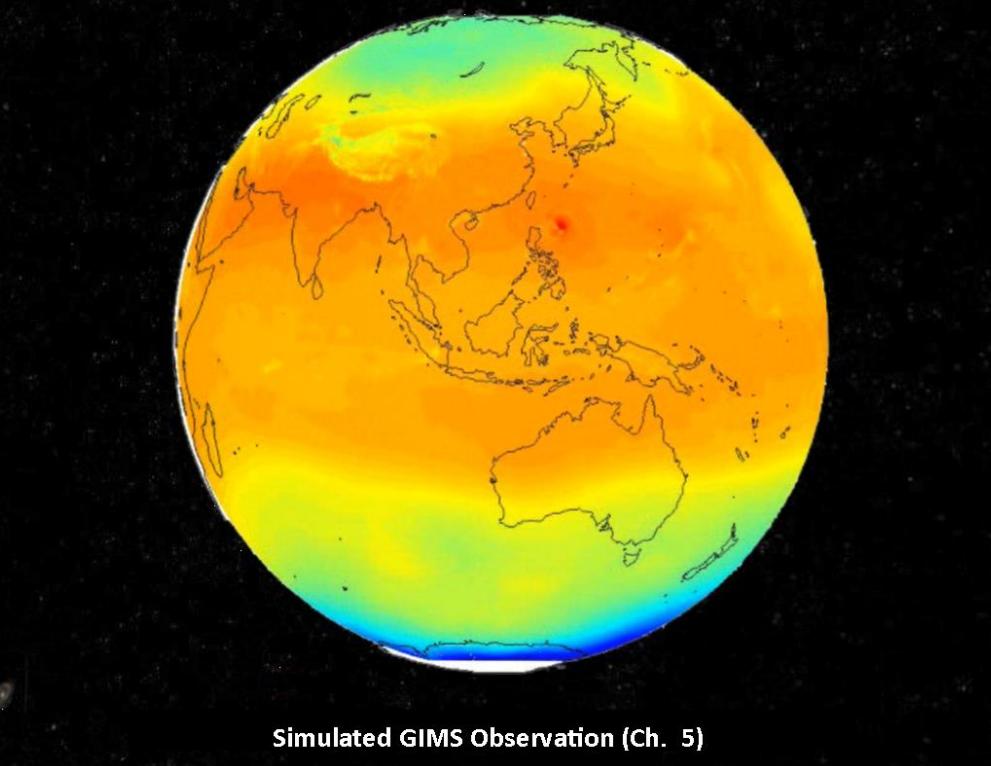
- Number of antenna units:  $N$
- Number of instantaneous VF samplers:  $C_N^2 = \frac{N(N-1)}{2}$



# Full-disk Coverage & 3D Sounding

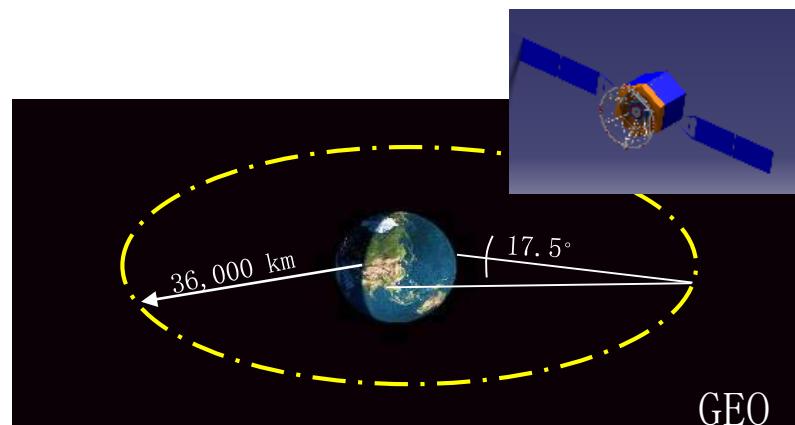
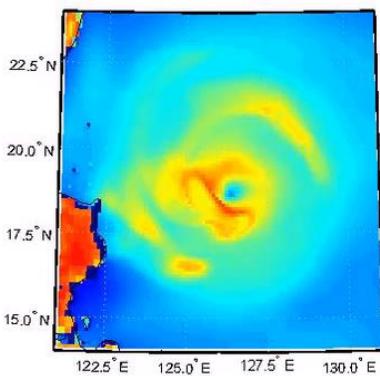
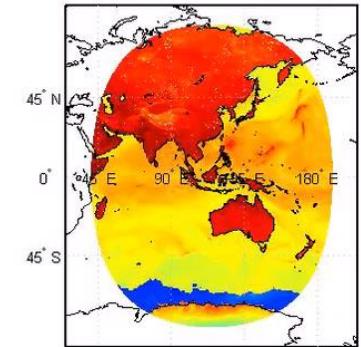


- Interferometric Synthetic Aperture Radiometry: **Full-Disk Coverage**
- Rotating Circular Thinned Array with 70 antenna units: **Fast Imaging Capability (5mins/channel)**
- 50-56GHz Sounding Channels: **3D Atmospheric Temperature Profiling**

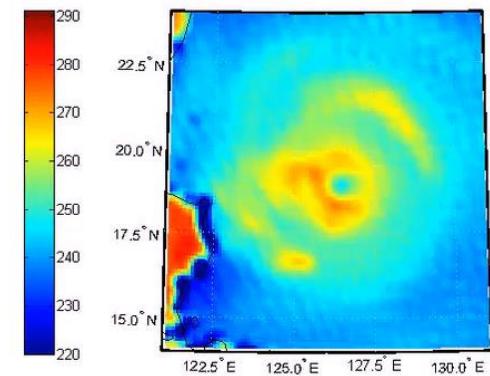
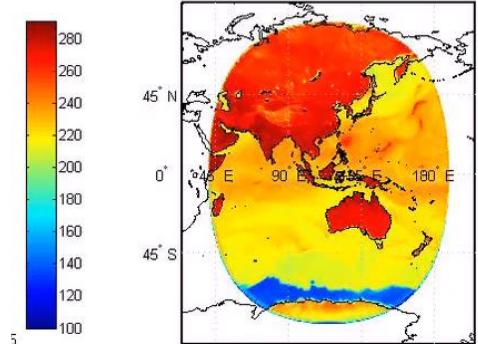


# Fast Imaging → Tracking the evolution of TCs

**Dynamic Target  
(FNL+WRF+RTTOV)**



**Simulated GIMS Observation  
(Full-disk, 30km res, 5mins)**

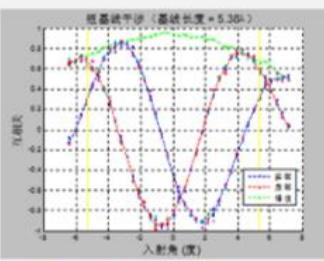
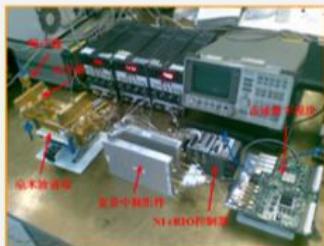


- 30-mins duration
- 5-min imaging period,
- 10s image refresh based on kalman-filtering & interpolation processing

# GIMS development roadmap

2006~2008:

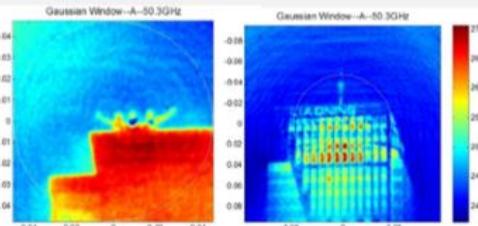
Two-element Interferometer



Point Source Fringe test

2008~2012:

28-element GIMS-I Demonstrator



GIMS-I Field Imaging Results

2014~2017:

Dual-mode microwave  
humidity/temperature sounder concept

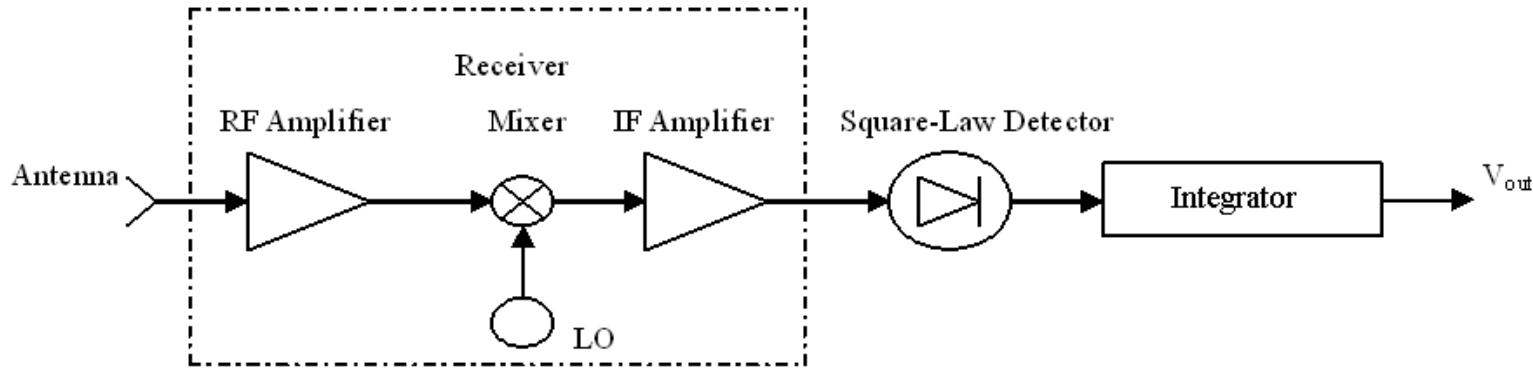


Where we are now!

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# Calibration of traditional total power radiometers



- Two-point calibration

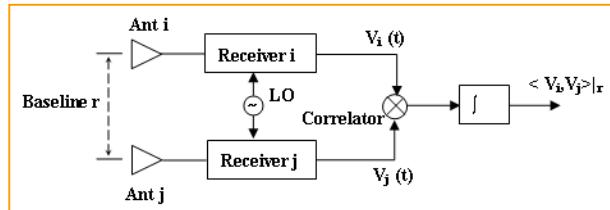
$$v = aT + b$$

● Receiver gain

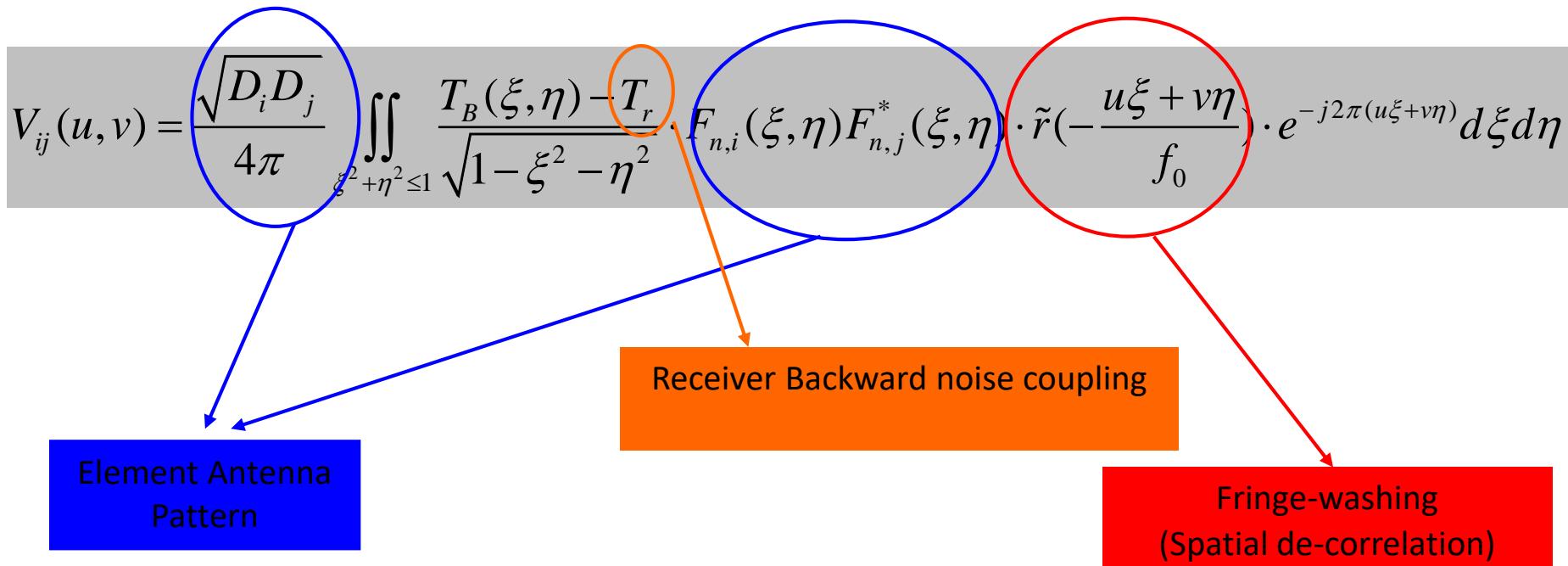
- Receiver noise temperature
- Loss before the first LNA, including antenna, isolator, switch, etc..
- Detector bias

# Calibration Considerations for Synthetic Aperture Radiometers (1)

- (1). Calibration is done on Visibility Functions, not on TB



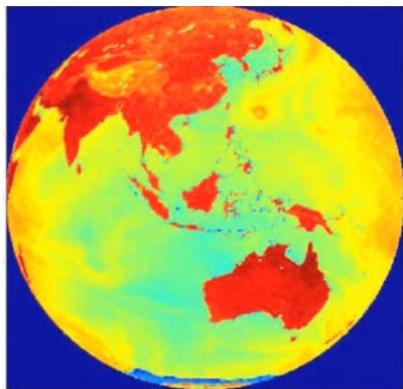
$$T = a_0 V_0 + a_1 V_1 + a_2 V_2 + \dots + a_n V_n + b$$



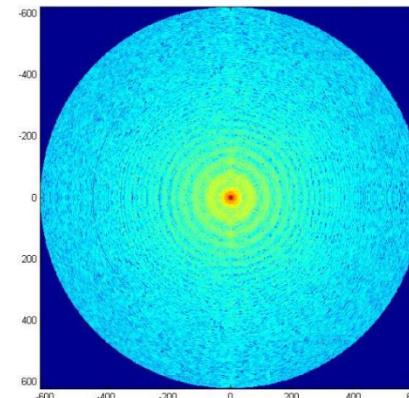
# Calibration Considerations for Synthetic Aperture Radiometers (2)

- (2). Short baselines is much more important than the long baselines
  - The value of the VF decrease very quickly from several tens K to  $10^{-3}$ K with the baselines length

TB distribution in spatial domain



VF samples in spatial frequency domain

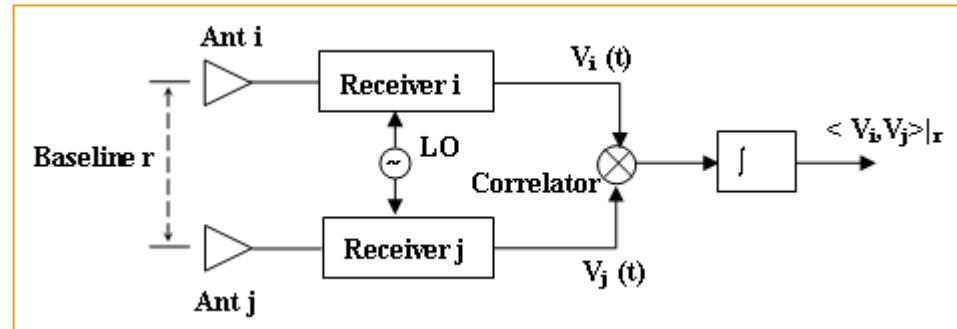


$$T = a_0V_0 + a_1V_1 + a_2V_2 \dots \dots + a_nV_n + b$$

# Calibration Considerations for Synthetic Aperture Radiometers (3)

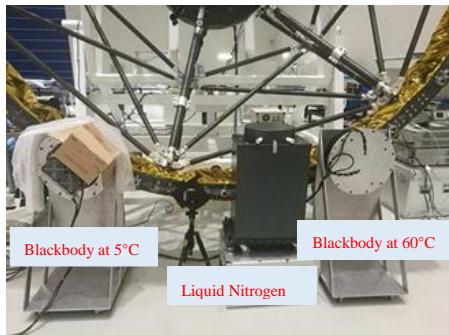
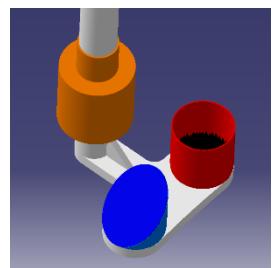
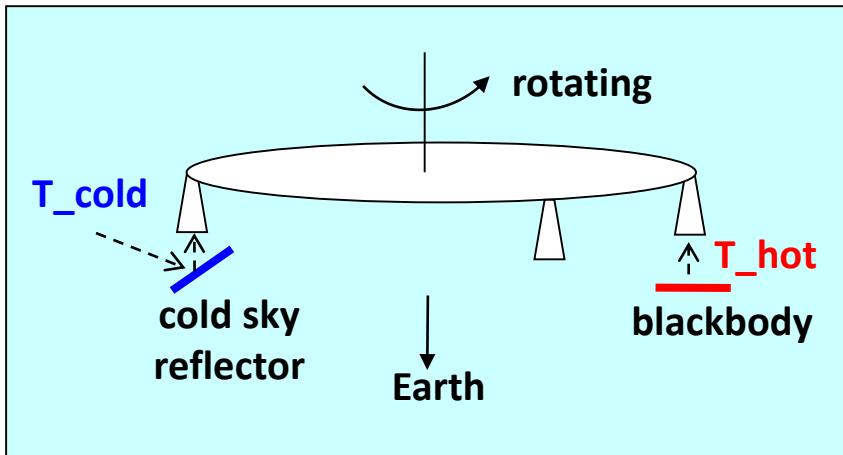
- (3). In-orbit Calibration approach
  - Inter-channel Amplitude Calibration
    - Noise injection
    - Two-point calibration
  - Inter-channel phase alignment:
    - Correlated Noise Injection
    - Redundant calibration
  - Antenna pattern calibration:
    - flat target response
  - Other issues:
    - Backward Noise coupling
    - IQ imbalance
    - Fringe-washing effects
    - Imaging algorithm error

$$T = a_0 V_0 + a_1 V_1 + a_2 V_2 + \dots + a_n V_n + b$$

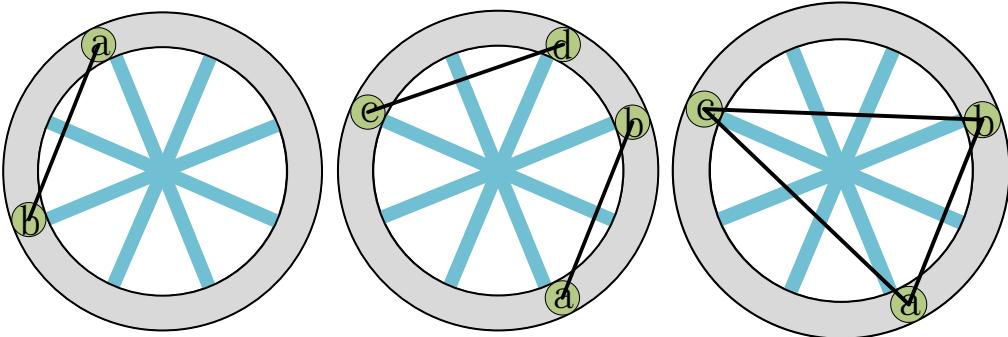


# GIMS Calibration Approach

- Amplitude:
  - two-point calibration



- Phase:
  - Self-calibration



$$\phi_{ab\_0}^{raw} + \phi_{ab\_180}^{raw} = 2(\alpha_a - \alpha_b)$$

$$\begin{aligned}\phi_{ab}^{raw} + \phi_{bc}^{raw} + \phi_{ca}^{raw} \\= \phi_{ab}^{id} + \alpha_a - \alpha_b + \phi_{bc}^{id} + \alpha_b - \alpha_c + \phi_{ca}^{id} + \alpha_c - \alpha_a \\= \phi_{ab}^{id} + \phi_{bc}^{id} + \phi_{ca}^{id}\end{aligned}$$

HAN D, LIU H, WU J, etc, "Inter-element Phase Calibration for Geostationary Interferometric Microwave Sounder (GIMS)", IEEE GRS Letter, 2016, 13(9)

# GIMS-II Demonstrator Calibration & Imaging Test

- **Calibration**

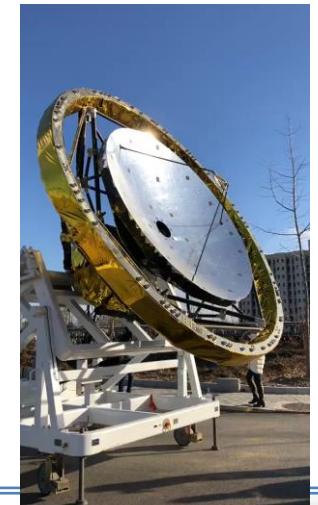
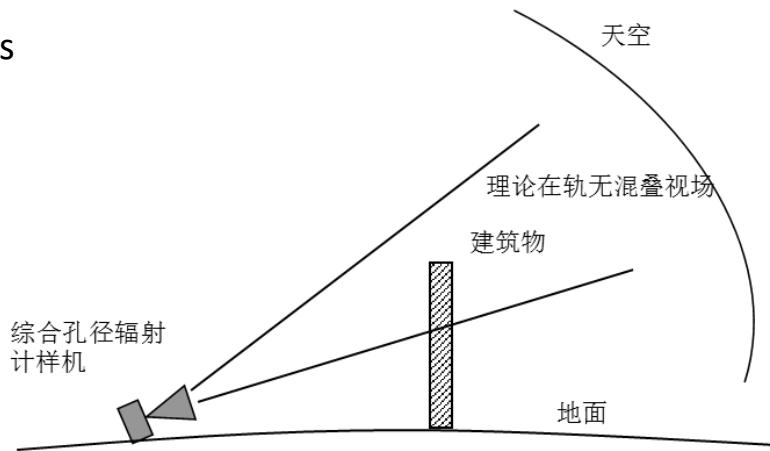
- ① Two-point Calibration of 70-element Receiving Channels
- ② Redundant phase calibration: point source imaging

- **Imaging Tests**

- ① Point Source: evaluation on spatial resolution
- ② Building
- ③ Solar transit
- ④ Imaging on Targets with changing temperature
- ⑤ Cold sky imaging: NEDT evaluation

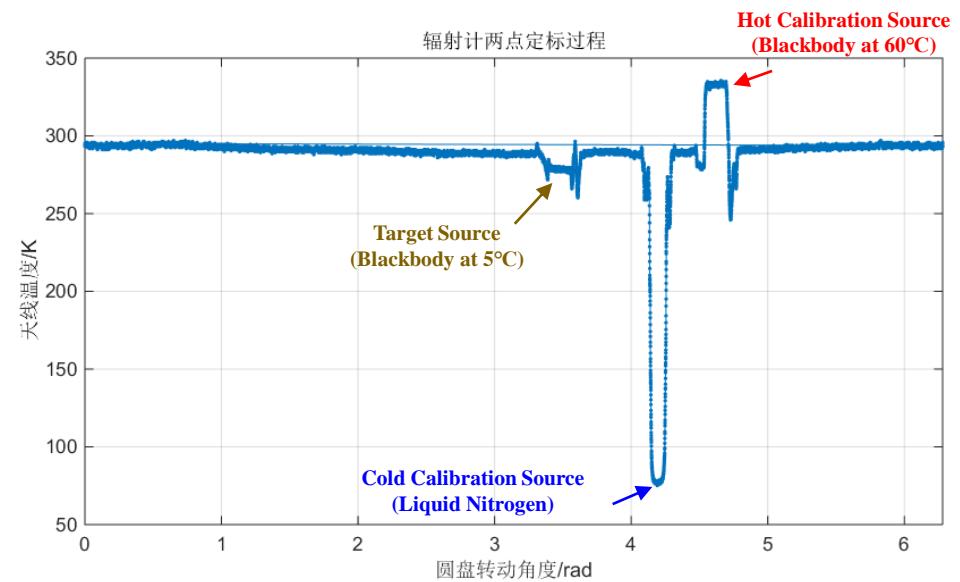
- **Restriction of On-ground Tests**

- Aliasing: no cold sky background  $> 18$  degree
- Near Field:  $D > 600$  wavelength  $\rightarrow$  Farfield  $> 5\text{km}$ .
- Calibration: No Cold Sky reference, Redundant Phase calibration in Near field
- Environment effects: temperature variation, wind, solar illumination, etc

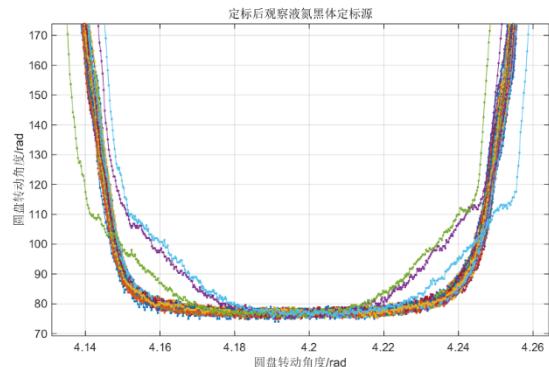


# End-to-end Two-point Calibration (70-elements)

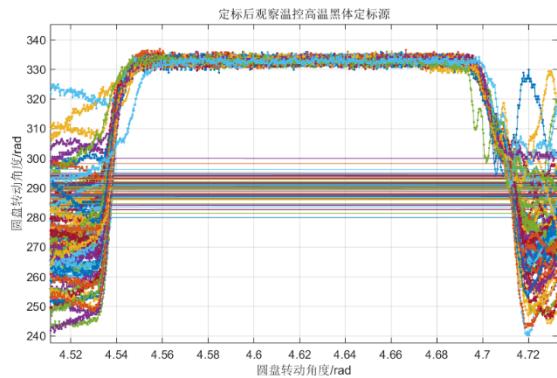
- Two calibration source + one target source



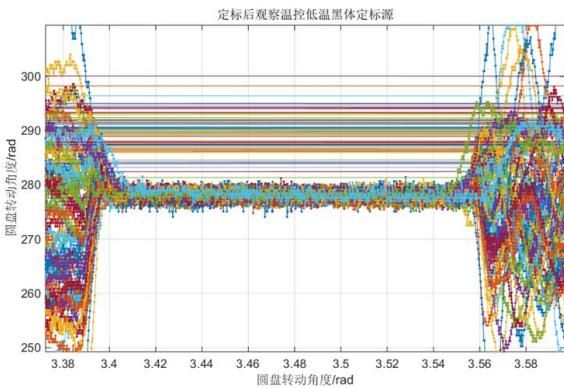
## Liquid Nitrogen



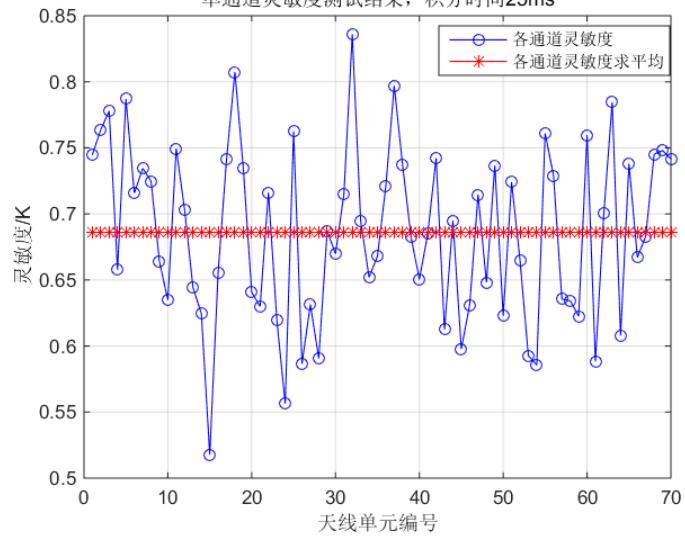
## Blackbody at 60 °C



## Blackbody at 5 °C

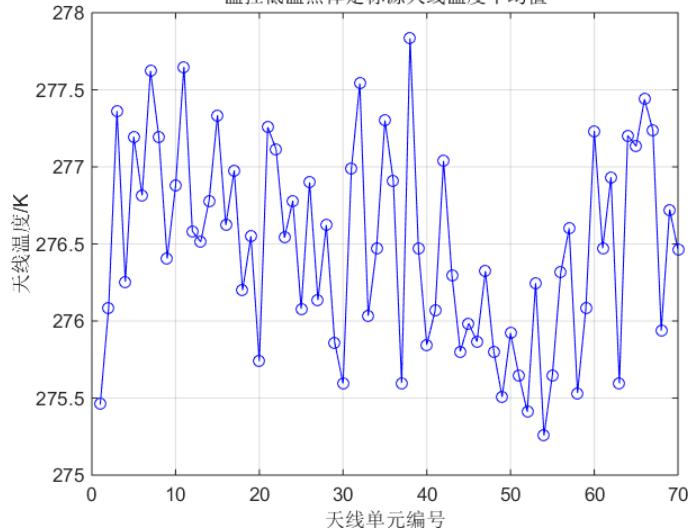


单通道灵敏度测试结果, 积分时间25ms



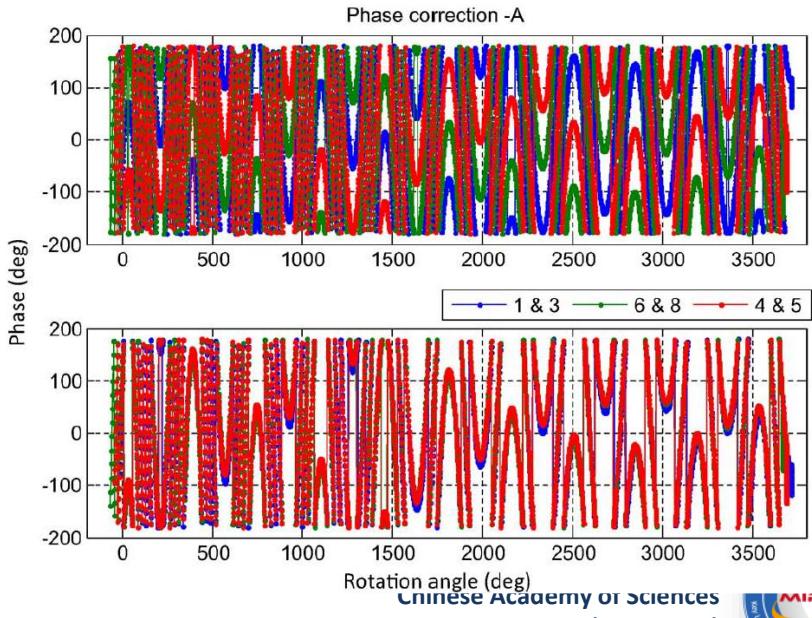
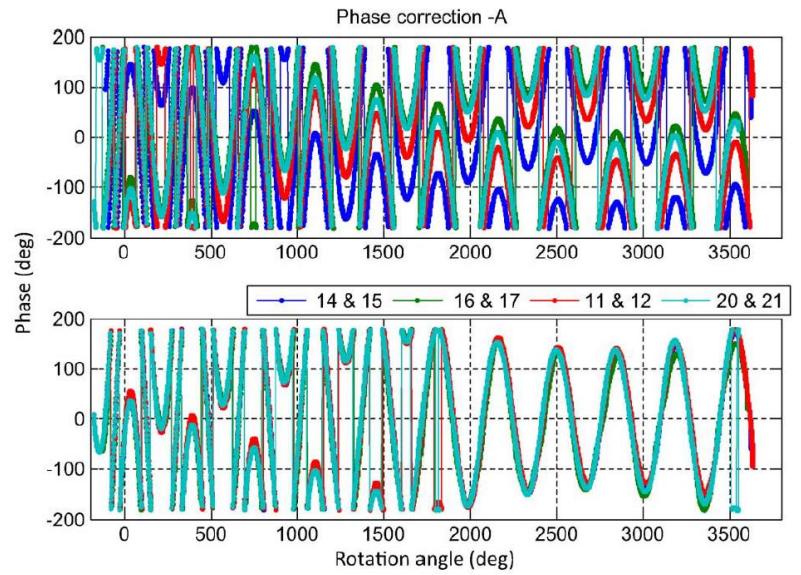
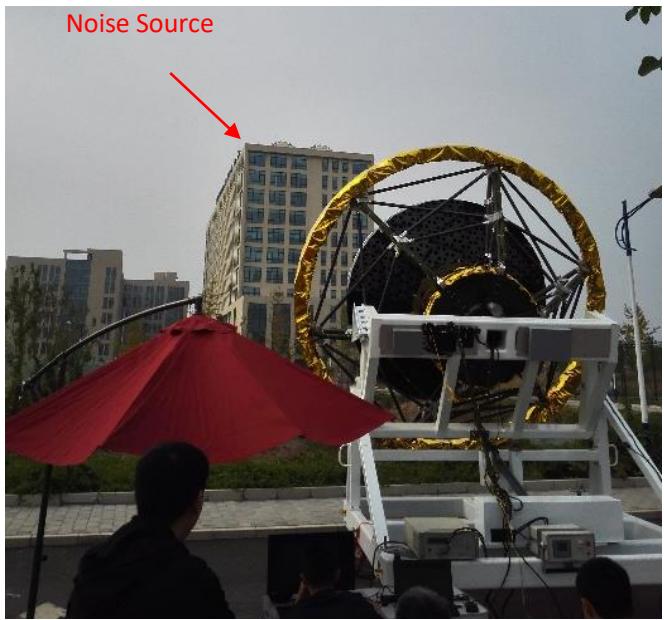
Mean NEDT of 70 receivers: 0.686K

温控低温黑体定标源天线温度平均值

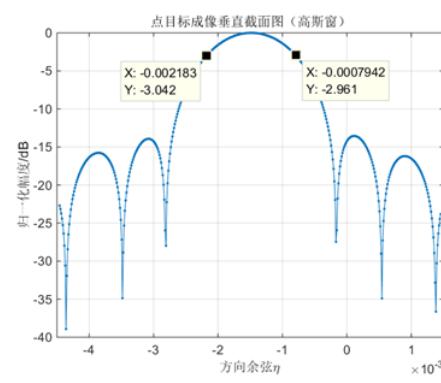
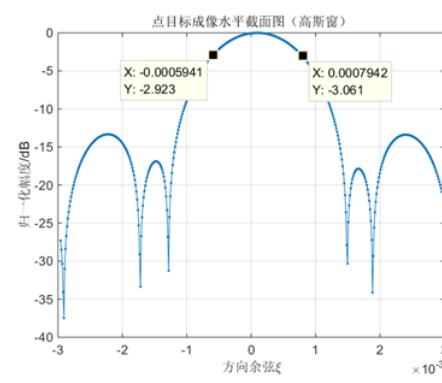
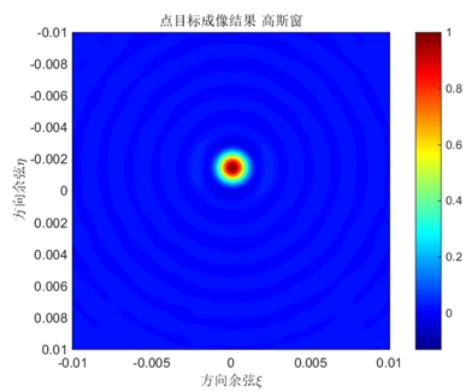
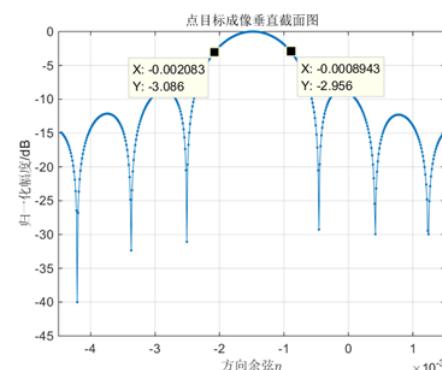
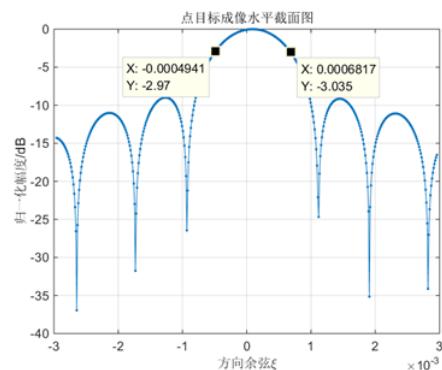
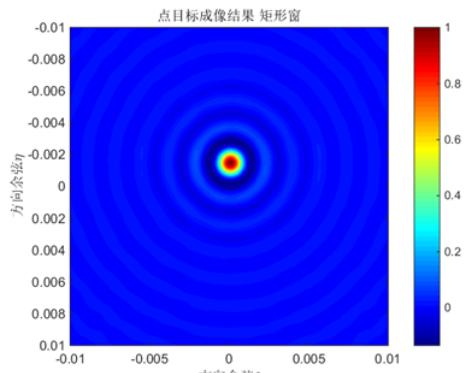


Mean TB of 70 receivers measurements: 276.48 K  
STD: 0.65K

# Redundant Phase Self-Calibration



# Point Source Imaging (spatial resolution)

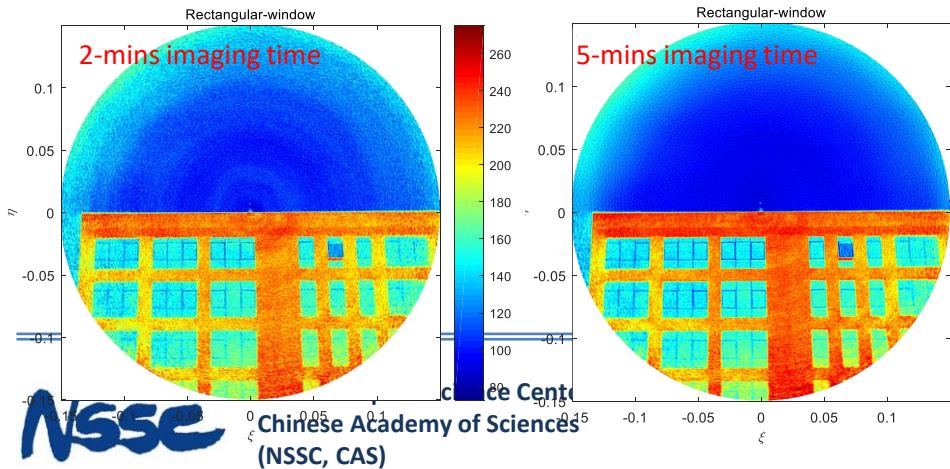
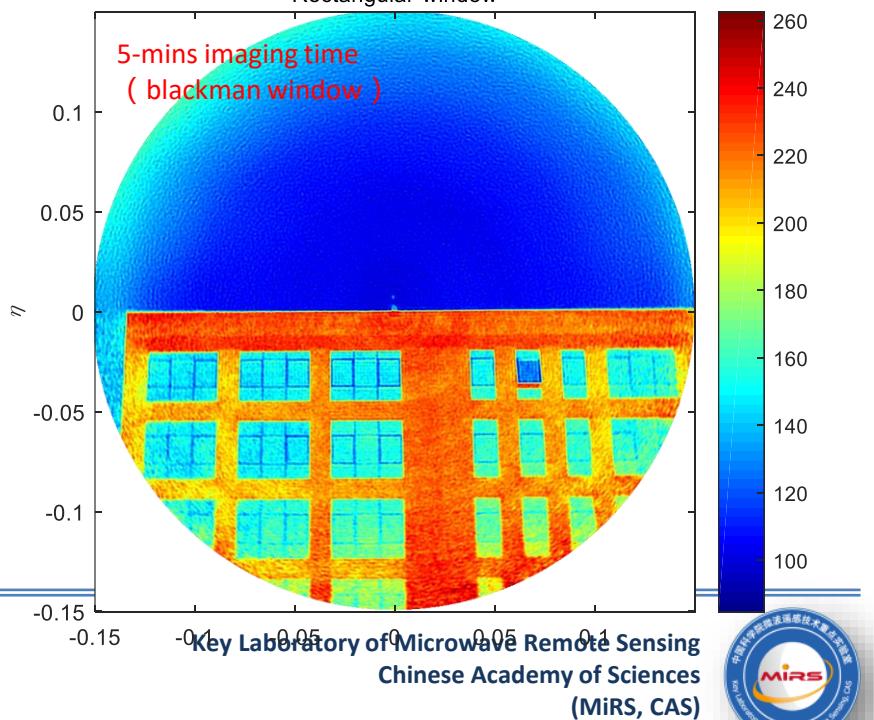
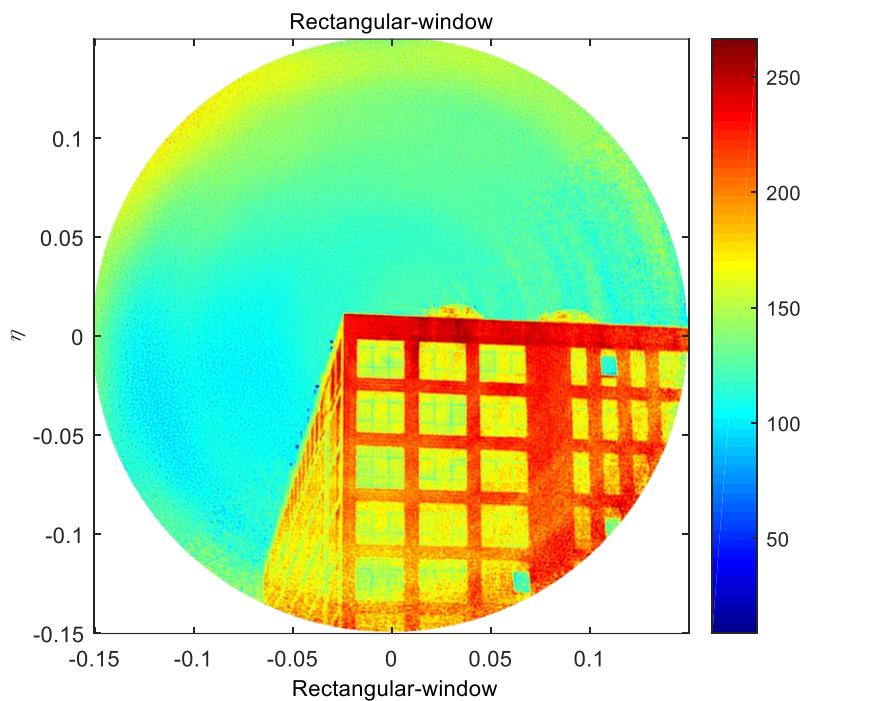


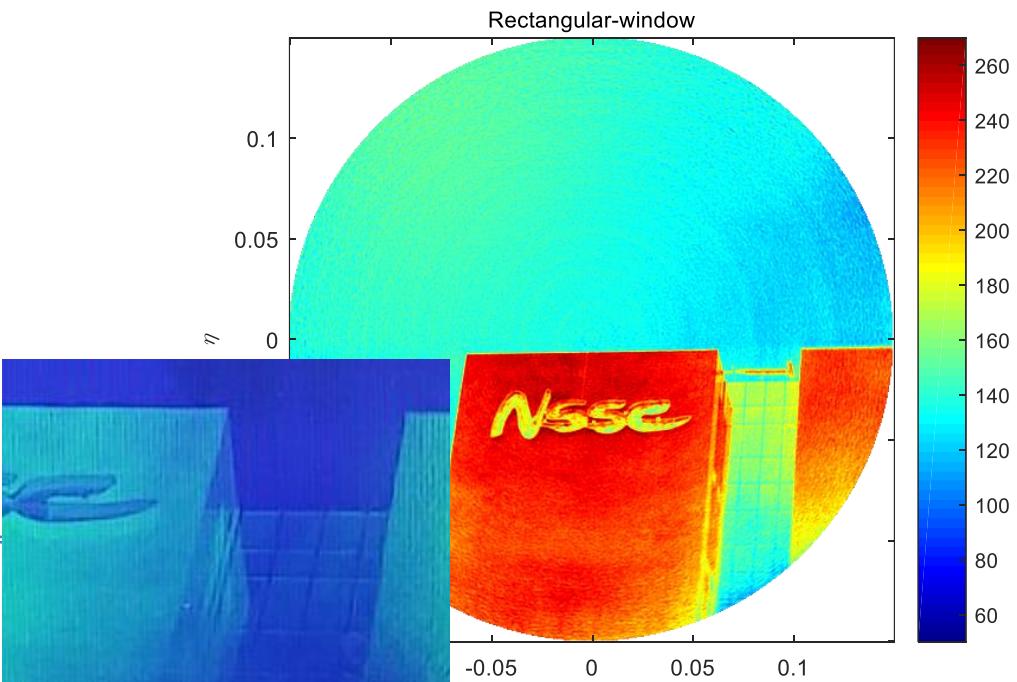
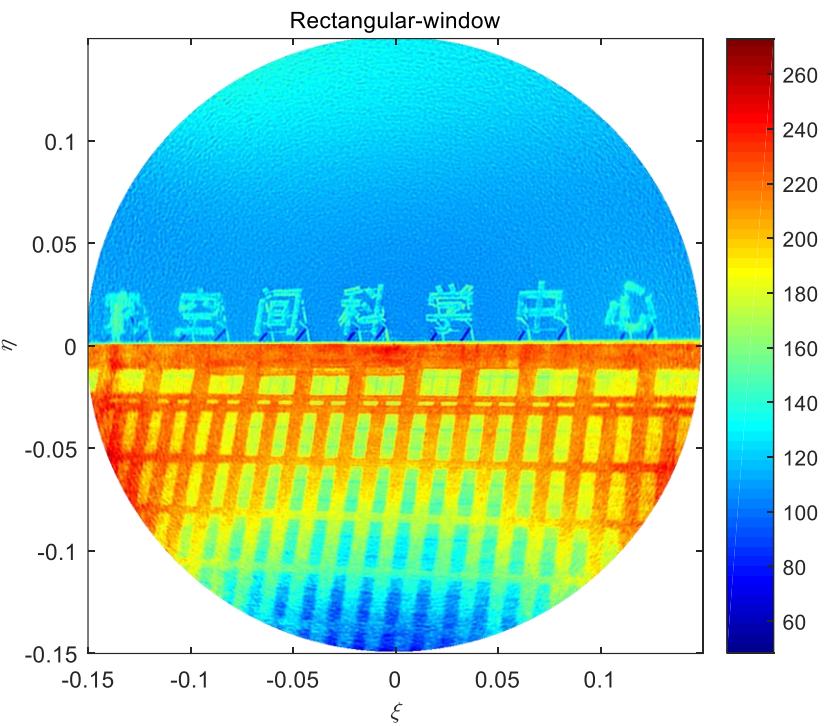
Noise Source



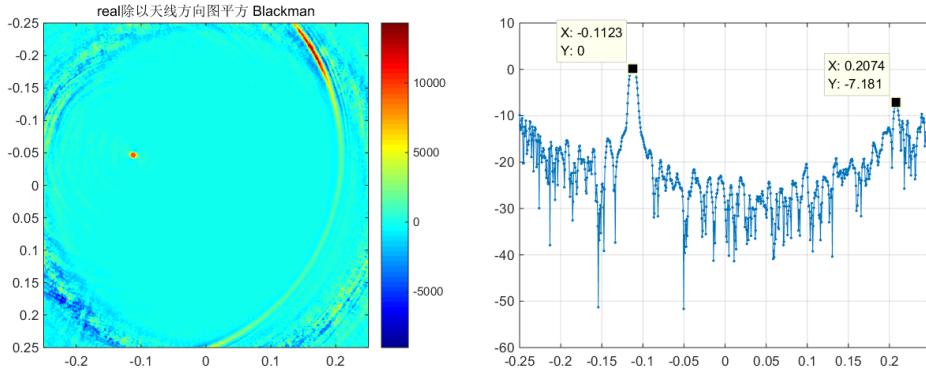
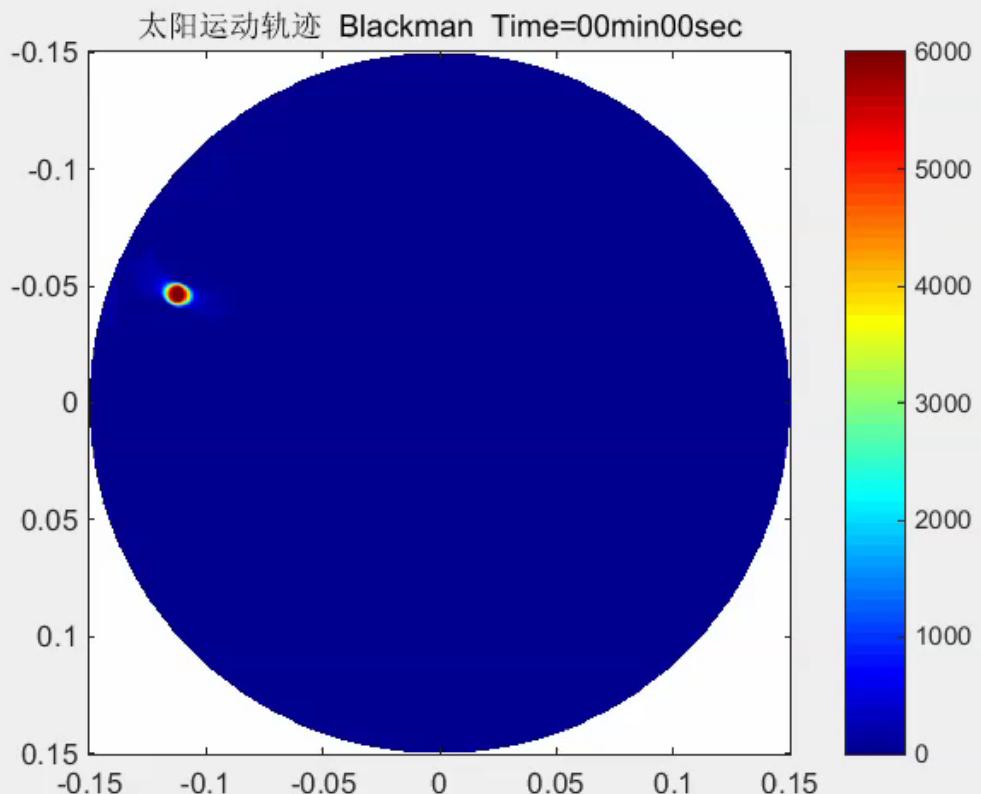
	测试结果	等效在轨分辨率
Rectangular Window	$(6.817+4.941)*1e-4*180/\pi=0.067^\circ$	42.1 公里
	$(20.83-8.943)*1e-4*180/\pi=0.068^\circ$	42.7 公里
Guassian Window	$(5.941+7.942)*1e-4*180/\pi = 0.080^\circ$	50.0 公里
	$(21.83-7.942)*1e-4*180/\pi = 0.080^\circ$	50.0 公里

# Building imaging





# Solar transit

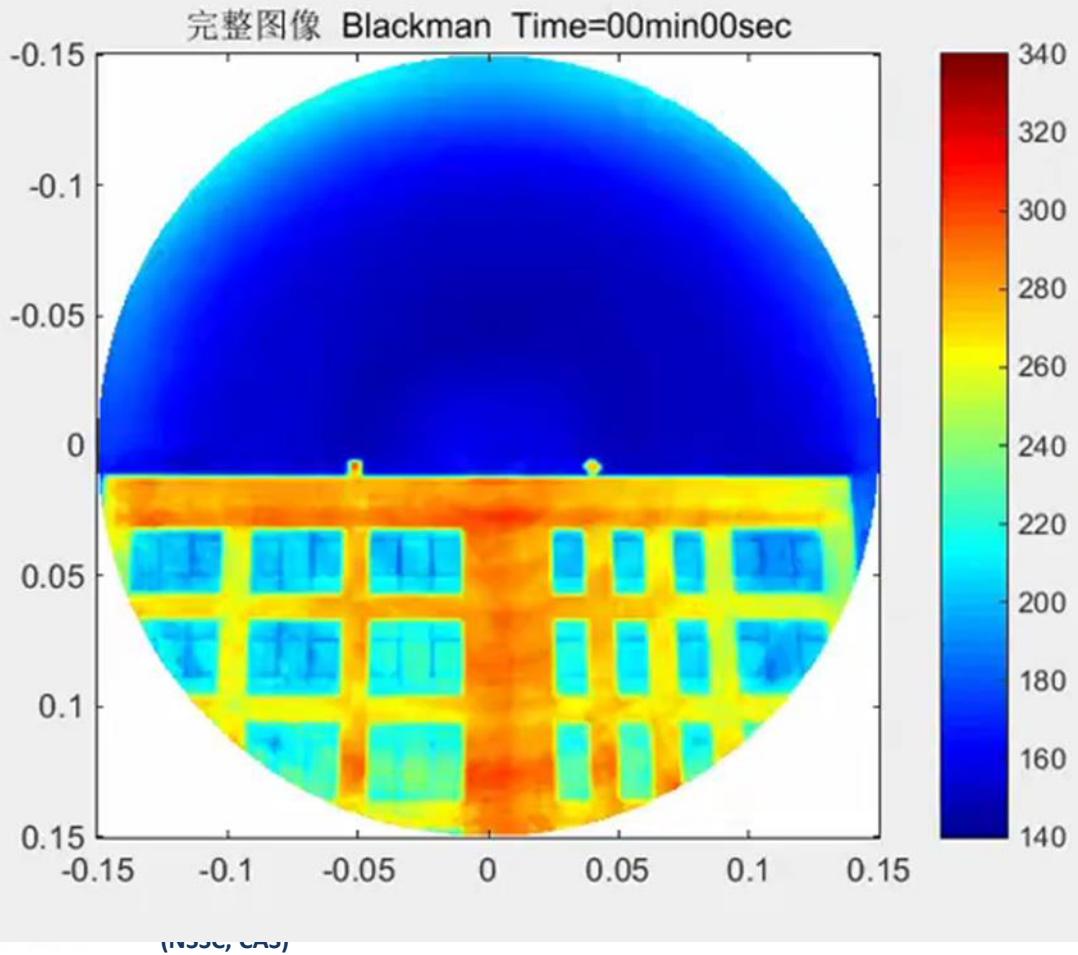
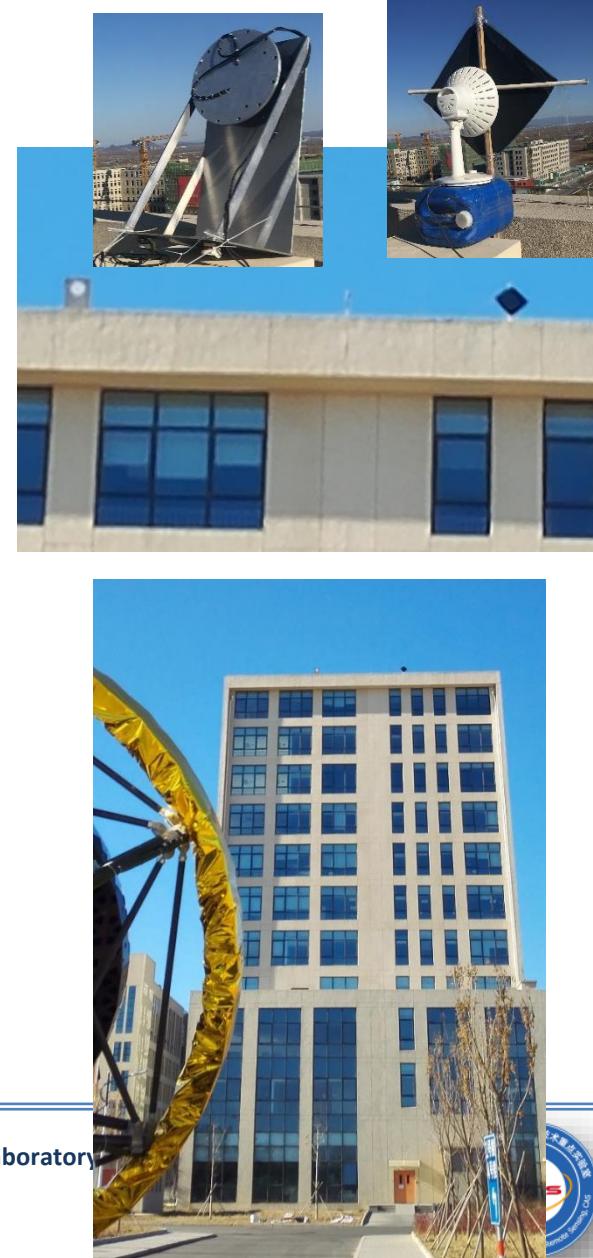


Aliasing free field of view evaluation

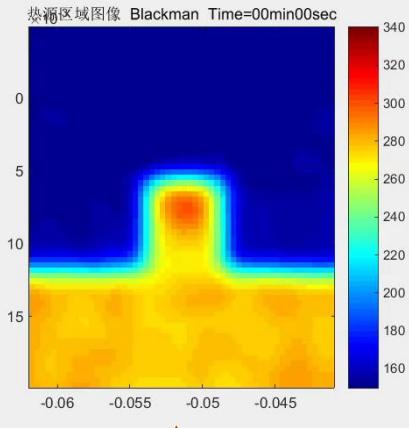
	Results
Left-side grating-lobe	$\pm \arcsin[(0.2074+0.1123)/2] = \pm 9.198^\circ$
Right-side grating-lobe	$\pm \arcsin[(0.1336+0.1849)/2] = \pm 9.163^\circ$

# Continuous Imaging Tests on Thermal Sources with Changing Temperature

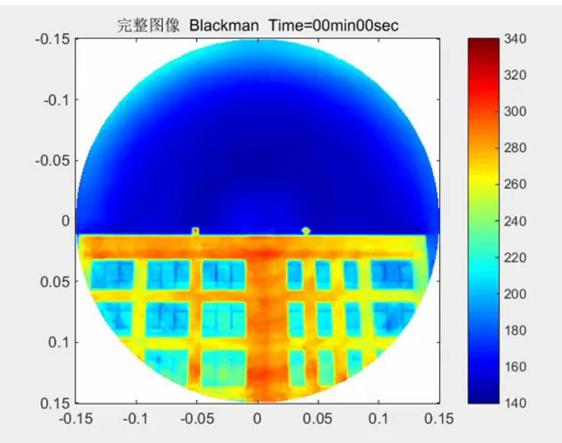
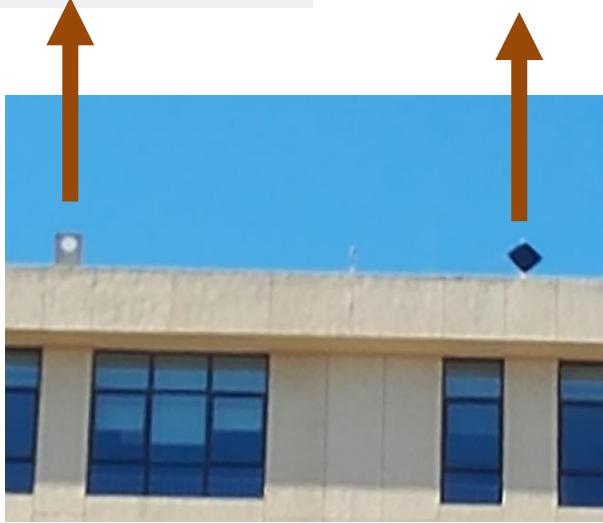
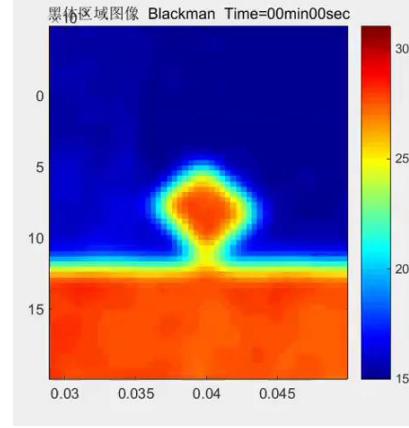
- Continuous Observation 110-mins
  - Hot Source ( $\sim 20^{\circ}\text{C} \rightarrow 60^{\circ}\text{C}$ , switch-off)
  - Blackbody (illuminated by an heater, switch-off)



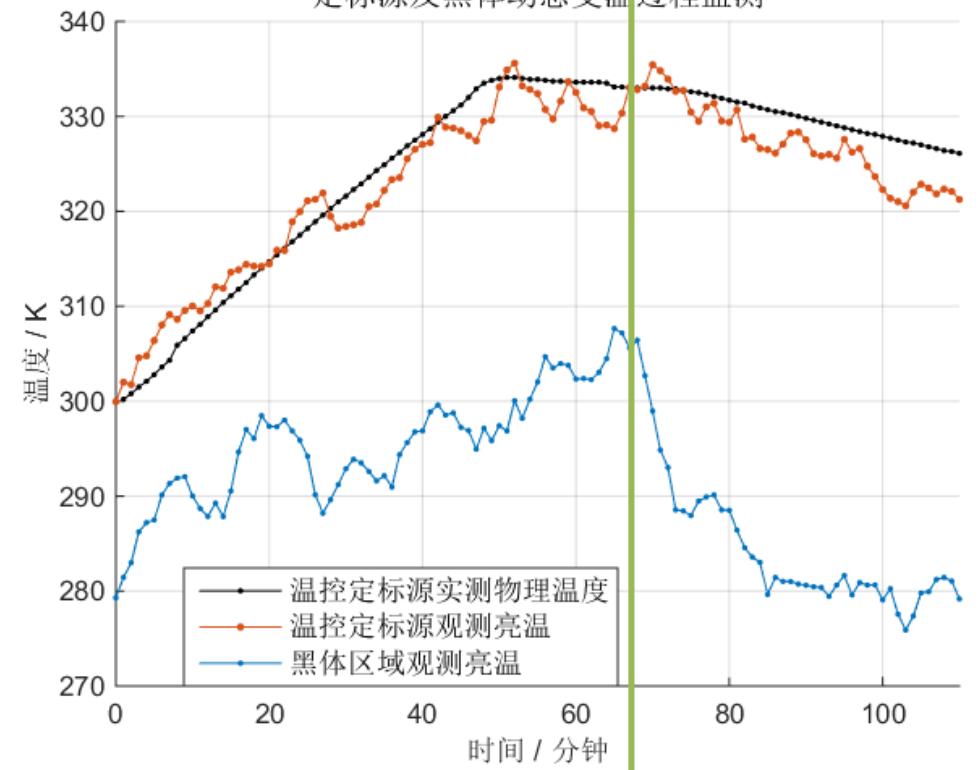
热源区域图像 Blackman Time=00min00sec



黑体区域图像 Blackman Time=00min00sec



### 定标源及黑体动态变温过程监测



Key Laboratory of Microwave Remote Sensing

Switch-off

Sciences  
LiRS, CAS)



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# Perspectives (1)

- 53/183GHz Joint-demonstration with ESA
  - 2015.01.30: PDR, ESTEC, Noordwijk, the Netherland
  - 2015.12.08: DDR, NSSC, Beijing, China
  - 2017.11.09: TRR, Omnisys, Gothenburg, Sweden

ESA > Our Activities > Space Engineering & Technology



## ESA AND CHINA TEAM UP ON TYPHOON-TARGETING IMAGER



Prototype for ground testing

17 January 2018 ESA has teamed up with the Chinese Academy of Sciences to test an instrument capable of peering down from orbit through dense clouds and rain to sound the depths of typhoons and storms.

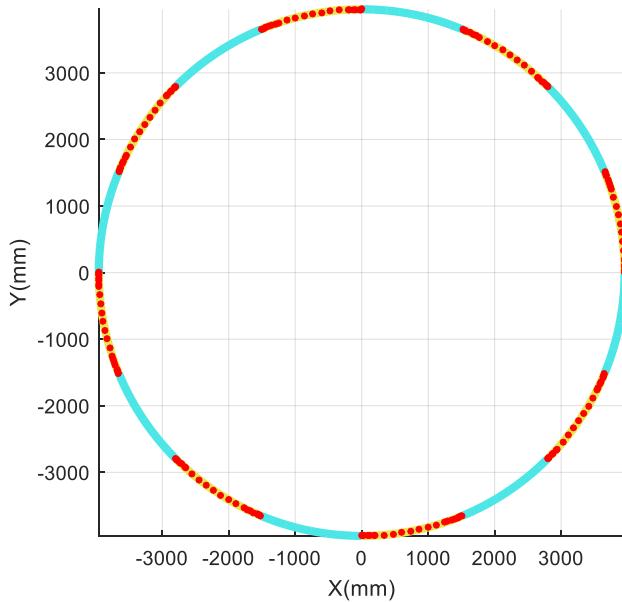
China's National Space Science Center – an entity under the Academy – in Beijing has built a 3 m-diameter prototype millimetre-wave instrument for ground testing. A smaller ESA-led instrument that works on a separate, complementary frequency band was slotted into it, then the combined instrument underwent ground testing.

"China has an obvious interest in typhoons, and enhanced weather and climate forecasting is important to everyone," explains Peter de Maagt, heading ESA's Antennas and Sub-Millimetre Wave section.

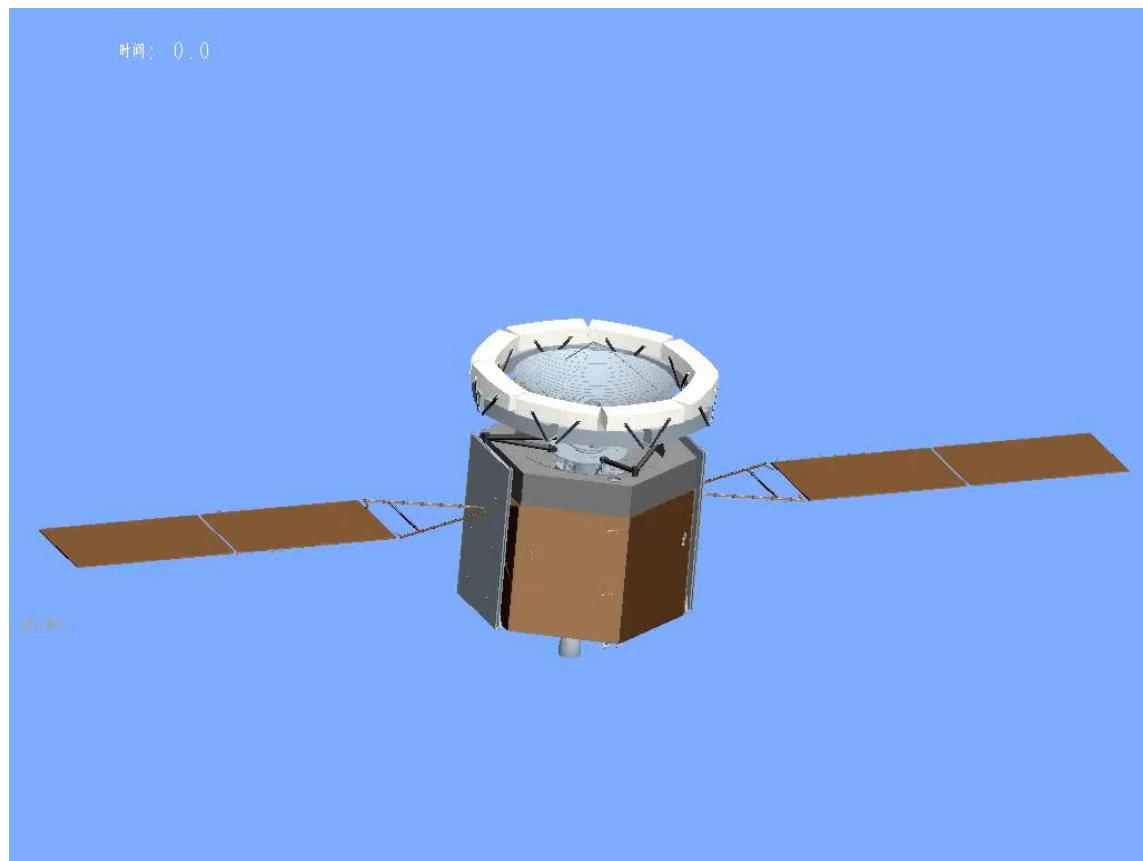


# Perspectives (2)

- Improved GIMS Design with deployable sub-array structure



- 50GHz
- Physical dimension  $D_a=7.91m$
- Equivalent dimension:  $D_b=14.82m$
- Ground resolution: 24km
- Antenna elements:  $8*15=120$



Thanks!

