



Inter-calibration Requirements / Algorithms for the VIS/NIR:

Starting Points for Discussion

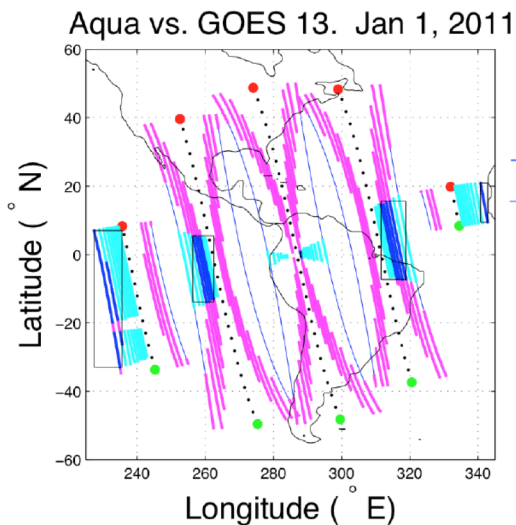
Constantine Lukashin
NASA Langley Research Center, Hampton, VA

GSICS Web Meeting

Inter-Calibration of Sensors in VIS/NIR: Current

Results from C. Roithmayr

Temporal matching: 30 minutes



• • • Aqua ground track

Aqua swath

- $|\Delta VZA| \leq 10^\circ$
- $|\Delta RAZ| \leq 20^\circ$
- $|\Delta VZA| \leq 10^\circ$ and $|\Delta RAZ| \leq 20^\circ$

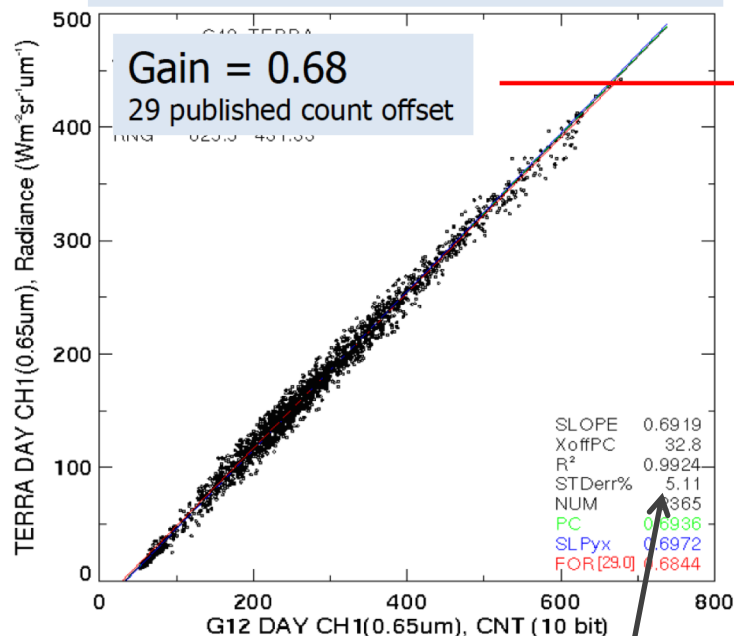
From the GSICS Report



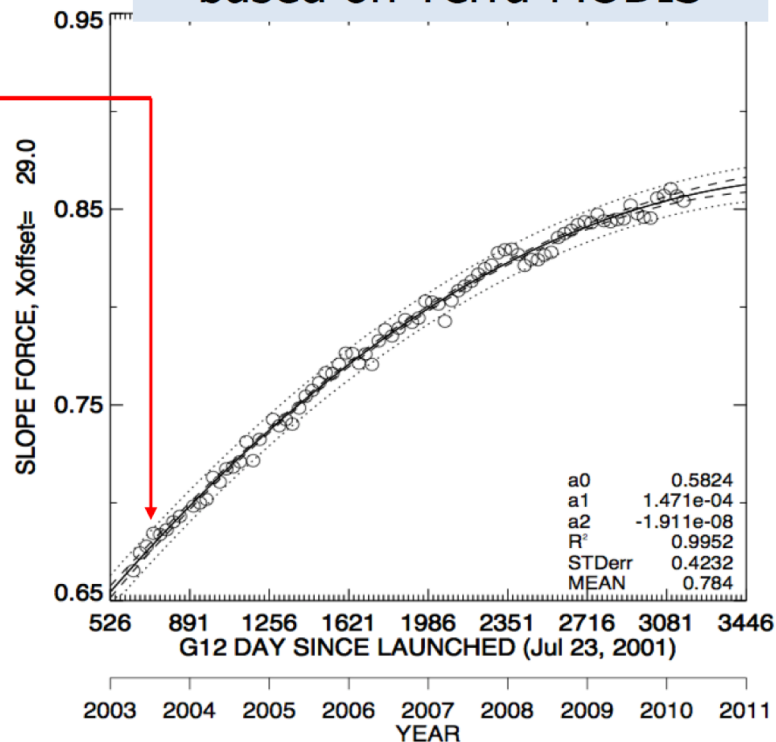
- ✧ LEO and GEO data matching when available (e.g. MODIS and GOES-13).
- ✧ Uniform and stable surface sites.
- ✧ Instrument stability by observing the Moon (e.g. SeaWIFS).
- ✧ Deep Convective Clouds, clear ocean & deserts: involve RT modeling.

Inter-Calibration of Sensors in VIS/NIR: Current

GOES-12/Terra-MODIS
July 2003



GOES-12 gain
based on Terra-MODIS



Results from the GSICS Report

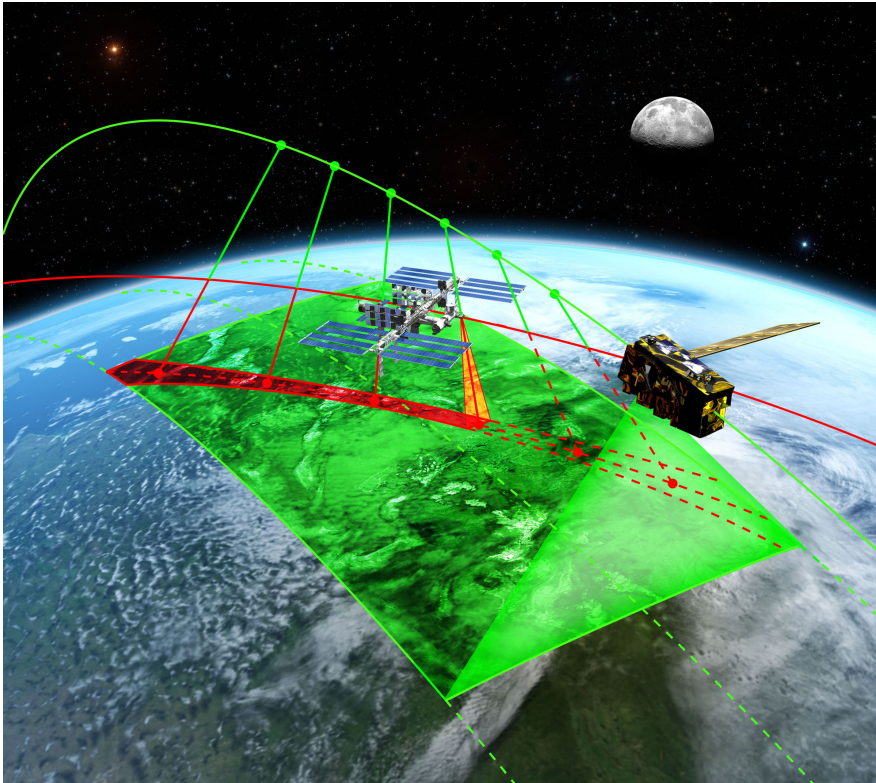
Inter-calibration of gain:

- ✧ Type A uncertainty (random) is 5.11% (k=1): due to data matching.
- ✧ Type B uncertainty (not random) is defined by the MODIS accuracy of 2% (k=1) [pre-launch].
- ✧ Spectral Type B uncertainty: due to difference in spectral response.

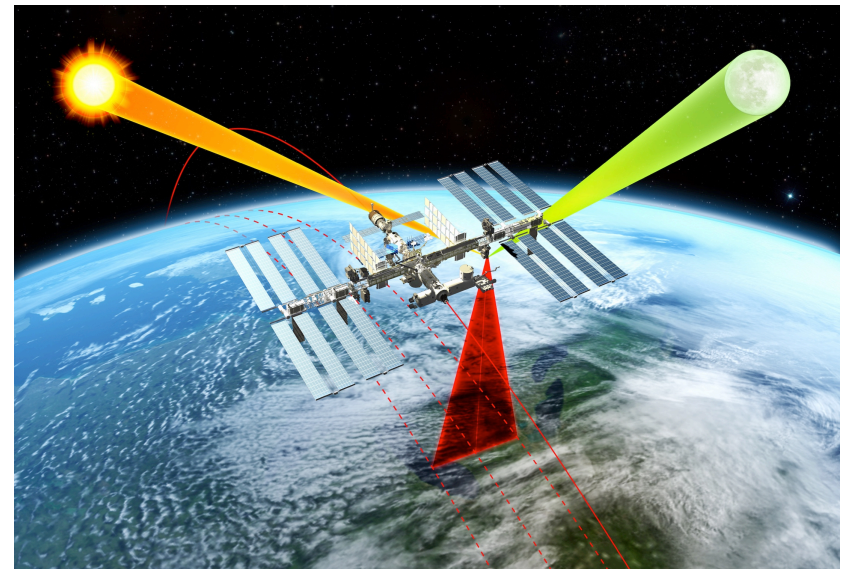
Fundamental limitation
of current approach

CLARREO Pathfinder Direct Inter-Calibration

Real time on-orbit operations for data matching



- ✧ CLARREO Pathfinder provides high-accuracy reference on orbit $\leq 0.3\%$ ($k=1$)
- ✧ CLARREO Pathfinder has 2D pointing ability for real-time data matching.



- ✧ CLARREO Pathfinder data matching with CERES and VIIRS on JPSS: temporal matching within 10 minutes, on-orbit angular matching $< 1^\circ$
- ✧ CLARREO Pathfinder location on ISS: ELC-1 Site 3

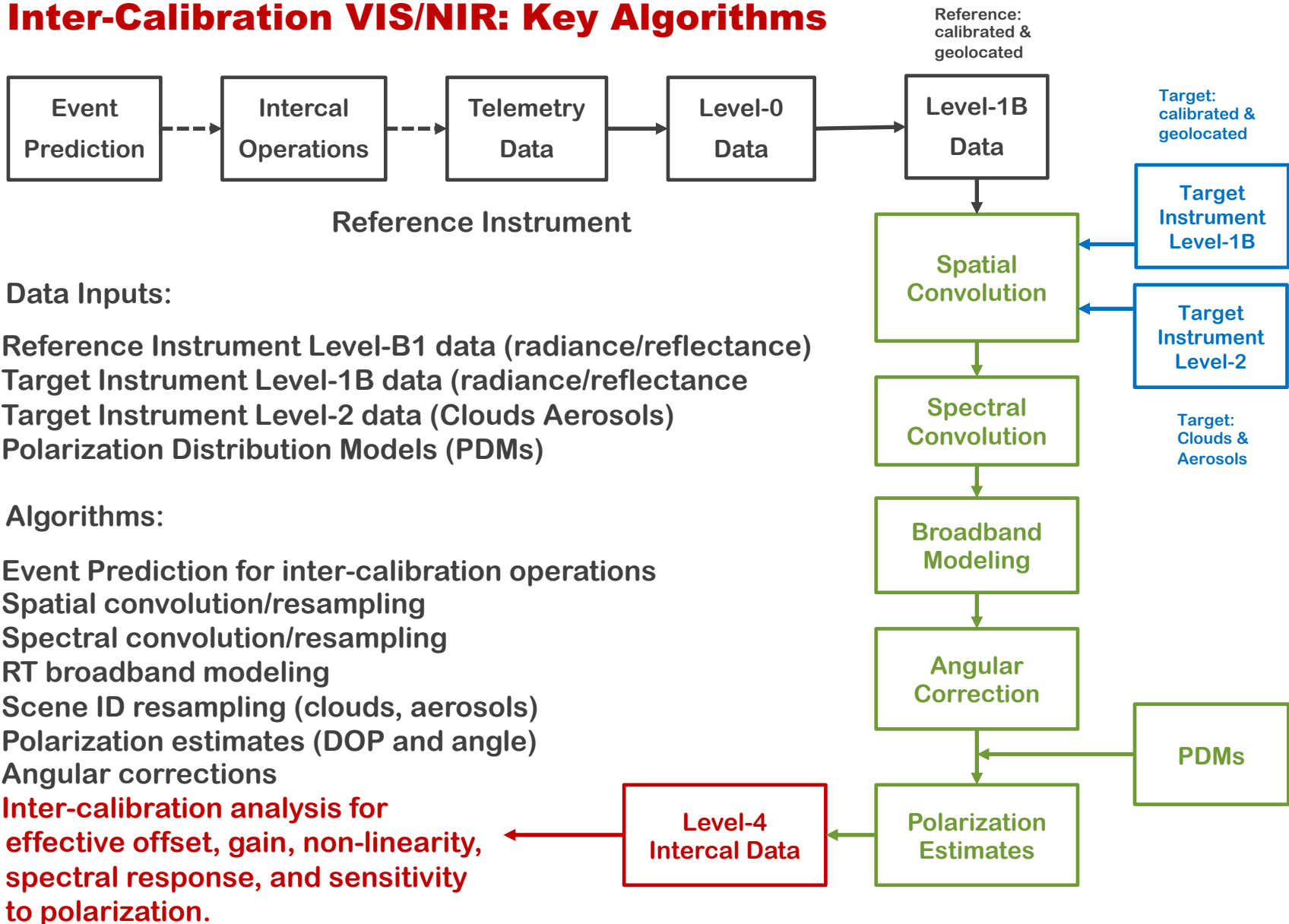


Inter-Calibration VIS/NIR: Instrument Key Requirements

- ✧ On-orbit Type B uncertainty for spectral reflectance $< 0.3\%$ ($k=1$)
Rationale: To provide an accurate reference
- ✧ On-orbit Type A uncertainty for spectral reflectance: can be relaxed
Rationale: It will be reduced by averaging over inter-calibration sample
- ✧ Spectral range: 350 nm to 2300 nm
Rationale: To cover most of RS broadband energy
- ✧ Nyquist spectral sampling < 4 nm
Rationale: To limit narrowband re-sampling error
- ✧ Spatial sampling < 1 km
Rationale: Comparable to inter-calibrated imagers (can be a trade)
- ✧ Spatial swath about 100 km
Rationale: To provide sufficient sampling for large-FOV sensors
- ✧ 2D Pointing ability
Rationale: To be able to match data on-orbit (time & angles), off-nadir sampling
- ✧ Orbit selection – LEO, polar versus low inclination
Rationale: Sampling distributions is more uniform for low inclination



Inter-Calibration VIS/NIR: Key Algorithms





Inter-Calibration VIS/NIR: Discussion

- ✧ Instrument Requirements & Algorithms: lots of trades
Target instruments parameters and their orbits are the drivers !
- ✧ Current CLARREO Pathfinder averaging approach:
 - *No by-detector inter-calibration of imagers (flat-fielding by imager team)*
 - *No high-spectral resolution inter-calibration (validation possible ?)*
- ✧ Broad spectral range:
Required only for broadband radiometers like CERES and GERB
- ✧ Lunar calibration and inter-calibration: very effective approach !
 - *Calibration target is already in space (Moon as solar diffuser)*
 - *Need for improvement in combined uncertainty $< 0.5\%$ ($k=1$)*
 - *Need for accurate measurement base for new model: hyperspectral with broad range*
 - *Need for modeling from hyperspectral to the shortwave broadband*
- ✧ Reference instrument calibration (and calibration generally) approach:
 - *Need for improvement in uncertainty and SI-traceability on-orbit*
 - *Improved calibration sources on-orbit: Solar Spectral and Total Irradiance (TSIS)*
 - *Calibration by direct viewing Sun and Moon*
 - *On-board calibration sources in VIS/NIR: a challenge !*