

Updates to the Absolute Radiometric Accuracy of the AIRS on Aqua

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Earth Observing Missions and Sensors

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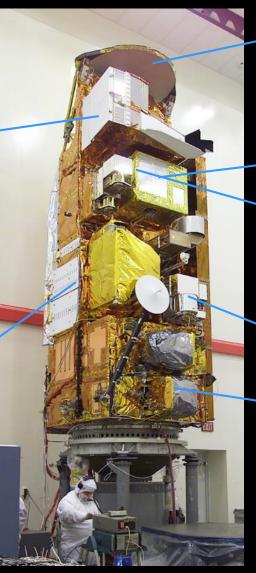
The Aqua Spacecraft Launched May 4, 2002



Moderate Resolution Imaging Spectroradiometer (MODIS) GSFC/Raytheon



Atmospheric Infrared Sounder (AIRS) JPL/BAE SYSTEMS



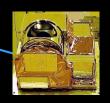
AQUA Spacecraft
GSFC/NGST



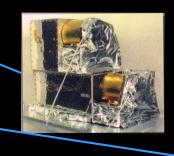
Advanced Microwave Scanning Radiometer (AMSR-E) MSFC/JAXA



Advanced Microwave Sounding Units (AMSU-A/B) JPL/Aerojet



Humidity Sounder from Brazil (HSB) JPL/Aerojet

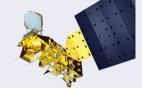




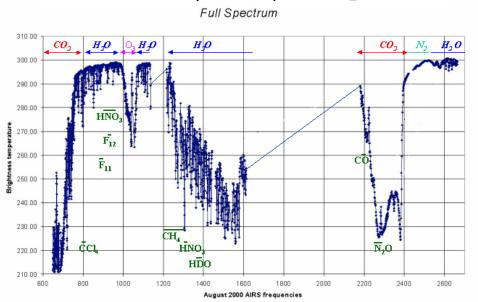
Clouds and Earth Radiant Energy System (CERES) LaRC/NGST

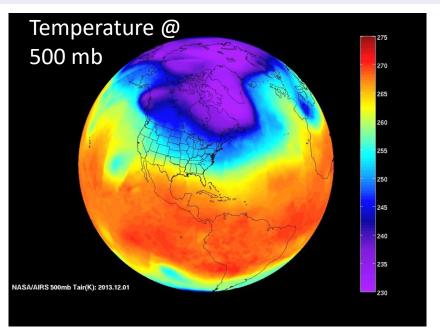


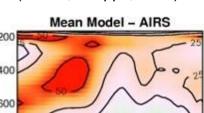
AIRS Supports Weather Forecasting and Climate Science



AIRS Channels for Tropical Atmosphere with T_surf = 301K

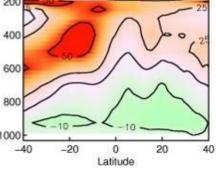






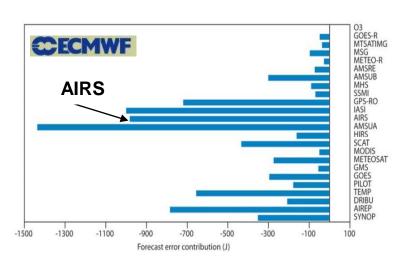
Water Vapor Climatology

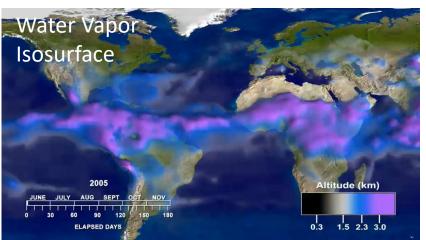
(Pierce, Scripps, 2006)



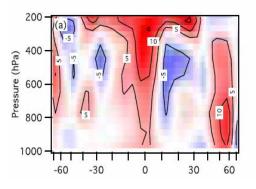
JPL/GSFC





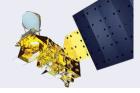


Water Vapor Feedback (Dessler, Texas A&M, 2008)





AIRS Global Daily Coverage Sees Critical Atmospheric Events



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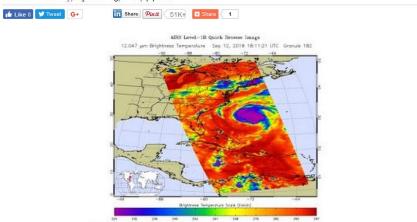
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✓ JPL Watches From Above as Hurricane Florence Creeps Closer to Landfall, Wreaking Havoc

Published: Thursday, September 13, 2018 | 4:40 AM



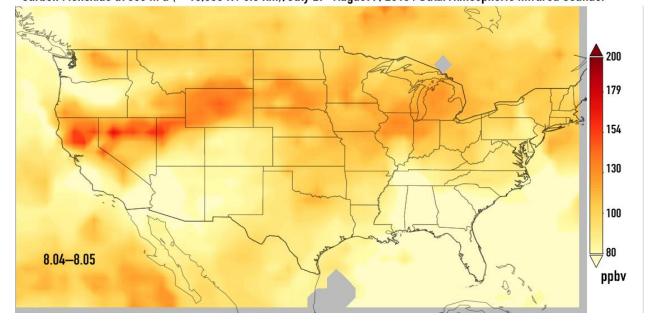
This image shows Hurricane Florence in infrared light, and was taken at 1:35 p.m. local time on Wednesday, September 12, 2018 by the Atmospheric Infrared Sounder (AIRS) on board NASA's Agua satellite. Florence underwent rapid intensification from Category 2 to Category 4 yesterday and was a Category 3 storm as of Wednesday evening. Image credit: NASA/JPL-Callech

All eyes were on Hurricane Florence Wednesday as the Category 3 storm barreled toward the U.S. East Coast. NASA's Atmospheric Infrared Sounder (AIRS) instrument was watching, too, and captured new imagery of the storm's approach.

AIRS, in conjunction with the Advanced Microwave Sounding Unit (AMSU), senses emitted infrared and microwave radiation from Earth to provide a three-dimensional look at weather and climate. It acquired infrared and visible light images at 1:30 p.m. EDT Wednesday. In the infrared image, a symmetrical ring of deep, cold rain clouds is shown in purple. Warmer areas, including a well-defined eye, are shown in blue. Shallower rain clouds are shown in green, while the red areas represent mostly cloud-free air moving away from the storm. The visible light image shows Florence much as our eyes would see it. It showcases the storm's thick cloud shield with clouds that extend far from the eye of the storm.

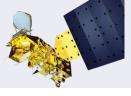
Thumisana Flananca undominant manid intensification from a Category a stamp to a Category 4 stamp scaling this week

Carbon Monoxide at 500 hPa (~ 18,000 ft / 5.5 km), July 29-August 7, 2018 / Data: Atmospheric Infrared Sounder





AIRS Designed for High Accuracy and Stability

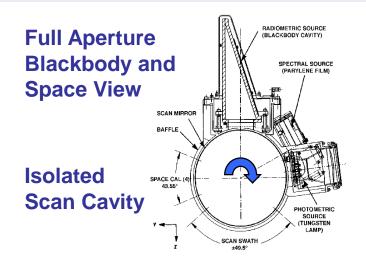


AIRS Features

- Orbit: 705 km, 1:30pm, Sun Synch
- Pupil Imaging IFOV: 1.1° x 0.6° (13.5 km x 7.4 km)
- Scanner Rotates about Optical Axis (Constant AOI on Mirror)
- Full Aperture OBC Blackbody, ε>0.998
- Full Aperture Space View
- Solid State Grating Spectrometer
- Temperature Controlled Spectrometer: 158K
- Actively Cooled FPAs: 60K
- No. Channels: 2378 IR, 4 Vis/NIR
- Mass: 177Kg,

Power: 256 Watts,

Life: 5 years (7 years goal)





Temperature Controlled Instrument



Active Detector Cooling



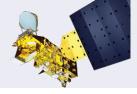
Grating Spectrometer

IR Spectral Range: $3.74\text{-}4.61~\mu\text{m}, 6.2\text{-}8.22~\mu\text{m}, \\ 8.8\text{-}15.4~\mu\text{m}$ IR Spectral Resolution: $\approx 1200~(\lambda/\Delta\lambda)$

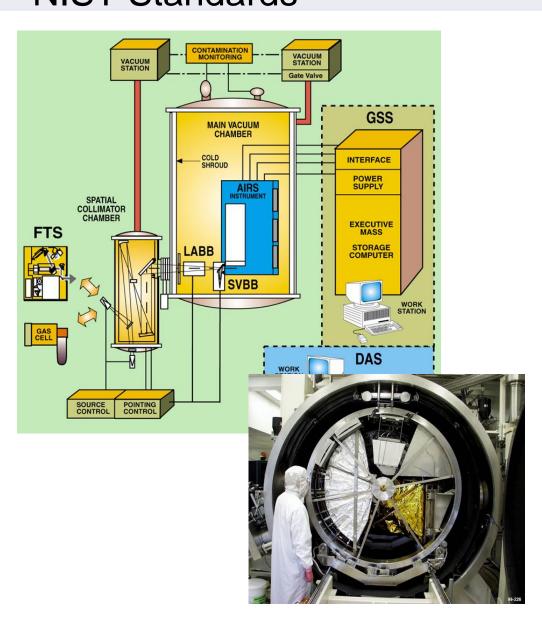
No. IR Channels: 2378 IR



Extensive Pre-flight Calibration on AIRS Ties to NIST Standards

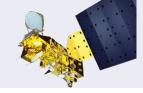


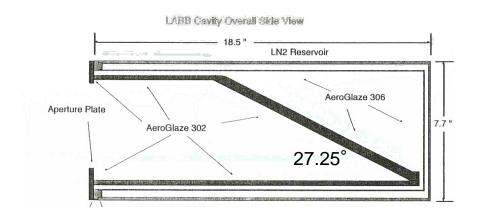
- Radiometric Response
 - Emissivity, Nonlinearity
 - Stray Light, Polarization
 - Scan Angle Dependence in TVAC
 - Transfer to On-Board Blackbody
 - 2 TVAC Cycles
- Spectral Response
 - SRF Characterization with FTS
 - Channel Spectra Characterized
- Spatial Response
 - Top-hat Functions All Channels
 - Stray Light Excellent
 - Far Field Response Excellent
- Good Documentation
 - Over 400 Design File Memos



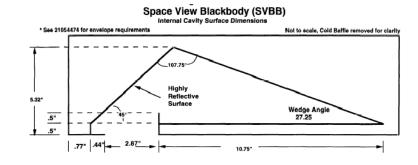


NIST Traceable External Large Area Blackbody (LABB) and Space View (SVBB)





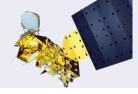
- Large Area Blackbody (LABB)
- T = 190K to 360K
- $\varepsilon > 0.99998$
- NIST Traceable PRTs (Rosemont)
- T_precision = 0.01K
- T_accuracy = 0.027K



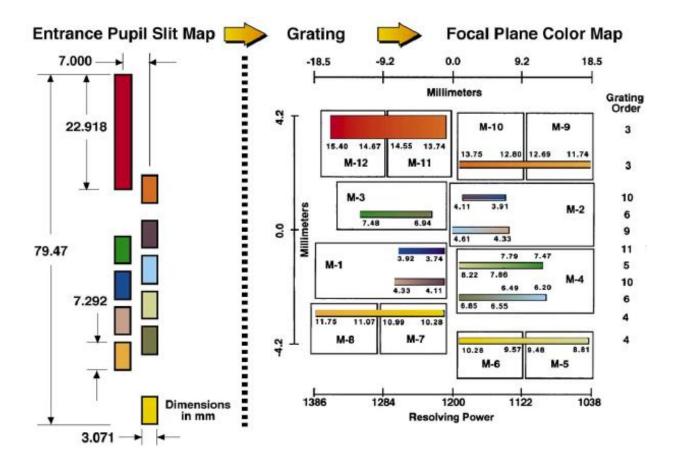
- Space View Blackbody (SVBB)
- T < 80 K
- $\varepsilon > 0.99998$
- T_precision = 0.01K
- T_accuracy = 0.5K



AIRS Images Pupil with Multiple Linear Arrays

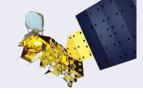


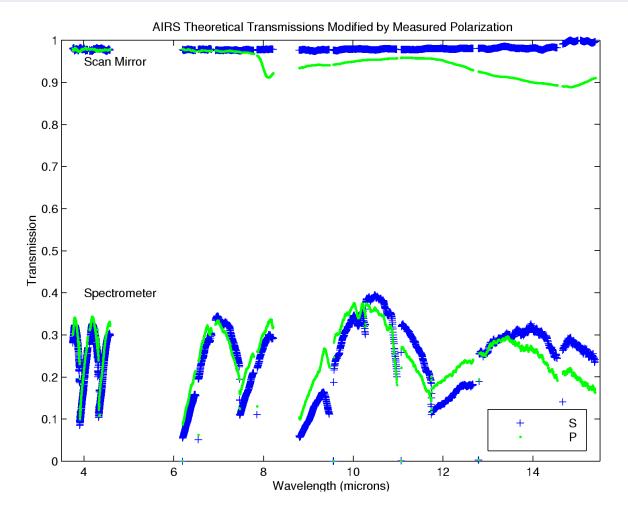
AIRS Module Spectral Band Limits						
		λ1	λ2	v1	v2	
		(µm)	(µm)	(cm ⁻¹)	(cm ⁻¹)	
1	M1a	3.752	3.934	2541.9	2665.2	
2	M1b	4.127	4.348	2299.8	2422.8	
3	M2a	3.891	4.088	2446.2	2569.8	
4	M2b	4.301	4.584	2181.5	2325.0	
5	M3	6.930	7.473	1338.2	1443.1	
6	M4a	6.196	6.489	1541.1	1613.9	
7	M4b	6.549	6.848	1460.3	1527.0	
8	M4c	7.469	7.786	1284.3	1338.9	
9	M4d	7.858	8.217	1217.0	1272.6	
10	M5	8.798	9.469	1056.1	1136.6	
11	M6	9.558	10.269	973.8	1046.2	
12	M7	10.264	10.974	911.2	974.3	
13	M8	11.065	11.744	851.5	903.8	
14	M9	11.731	12.670	789.3	852.4	
15	M10	12.790	13.735	728.1	781.9	
16	M11	13.728	14.543	687.6	728.4	
17	M12	14.663	15.394	649.6	682.0	

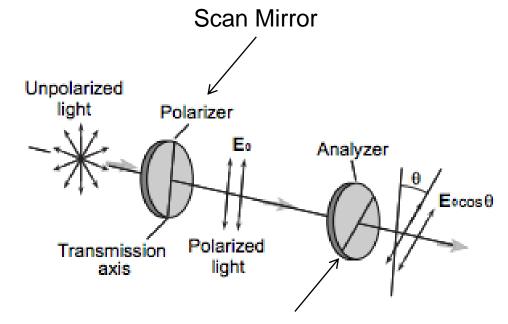




AIRS Scan Mirror and Spectrometer Act Like Polarizer and Analyzer







Spectrometer

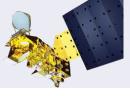
$$L_{ev} = L_{o}(\theta) + \frac{c_{0} + c'_{1}(dn_{ev} - dn_{sv}) + c_{2}(dn_{ev} - dn_{sv})^{2}}{[1 + p_{r}p_{t}cos2(\theta - \delta)]}$$

Assumes all space views at 90°

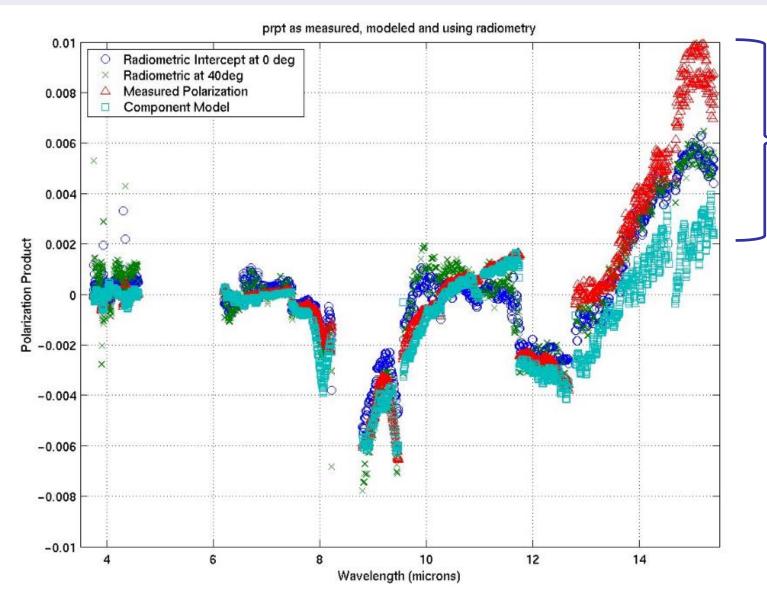
$$L_{o}(\theta) = \frac{L_{sm}p_{r}p_{t}[\cos 2(\theta - \delta) + \cos 2\delta]}{[1 + p_{r}p_{t}\cos 2(\theta - \delta)]}$$



Multiple Methods Used to Determine p_rp_t in V5



V5 Coefficients
Determined PreLaunch and Not
Changed Since

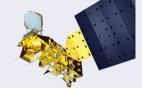


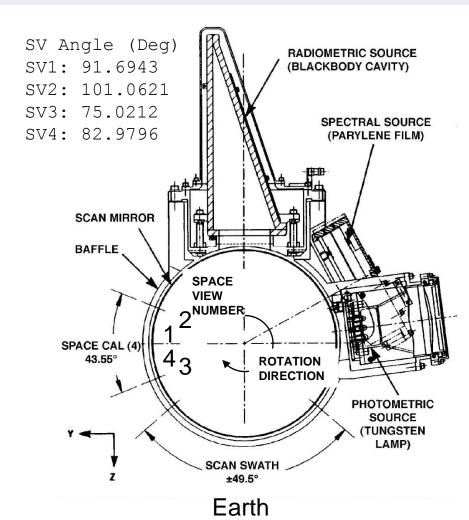
High Uncertainty Especially in LWIR

V5 is average of "Measured" and "Modeled" polarization, with Phase = 0

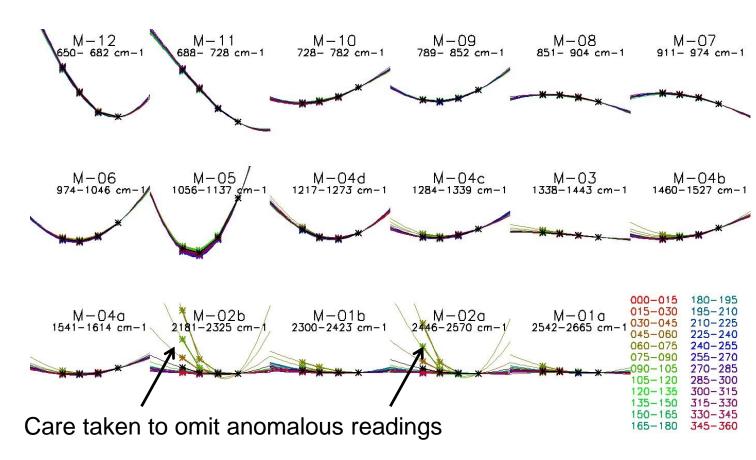


New Method Uses Multiple Space Views





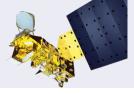
Every space view in the mission used to give 171 monthly averages

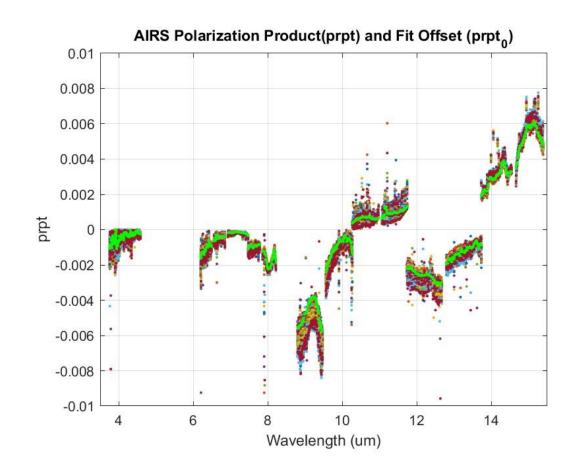


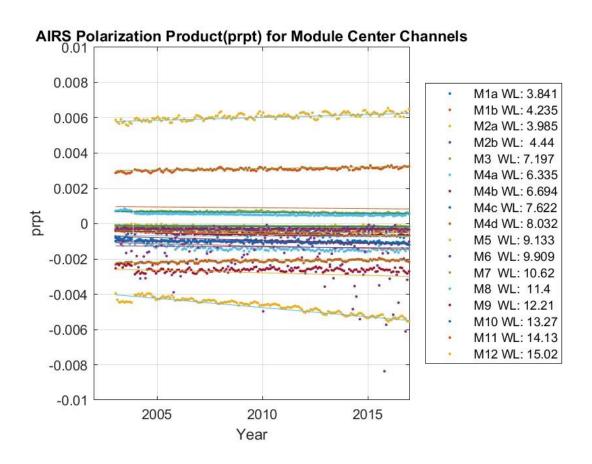
$$(dn_{sv,i} - dn_{sv,i})c_1' = -L_{sm}p_r p_t [cos2\theta_{sv,i}cos2\delta + sin2\theta_{sv,i}sin2\delta + cos2\delta]$$



Polarization Parameters Calculated Over Mission Life



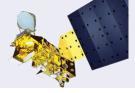


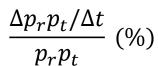


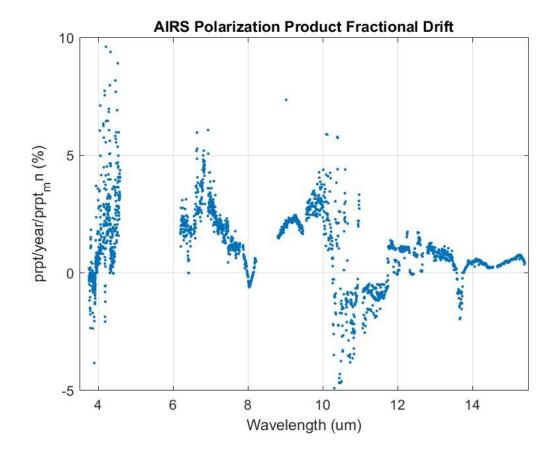
Polarization and Phase in V7k are Time Dependent Using Linear Fit Offset and Slope



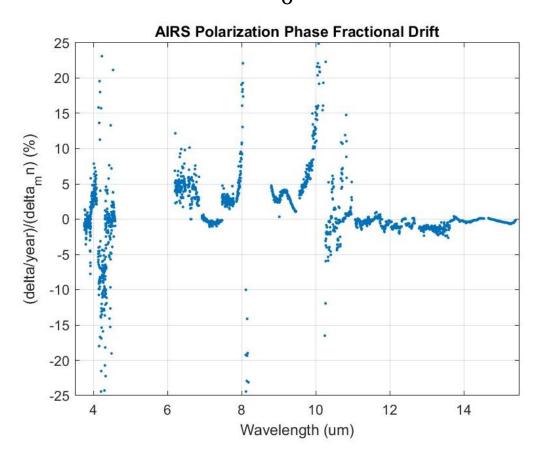
V7k Annual Change in Polarization and Phase





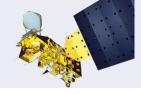


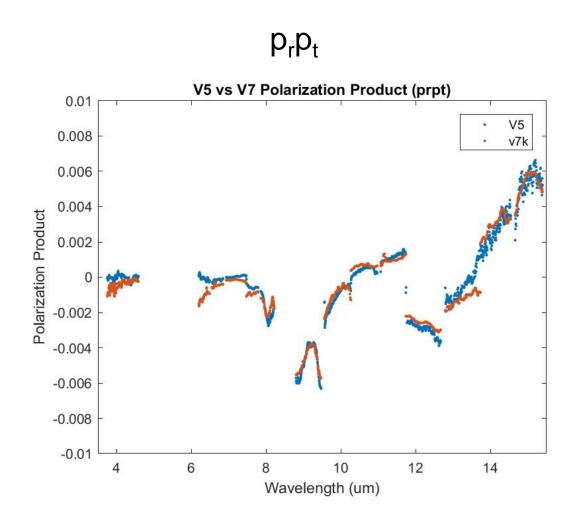
$\frac{\Delta\delta/\Delta t}{\delta}$ (%)

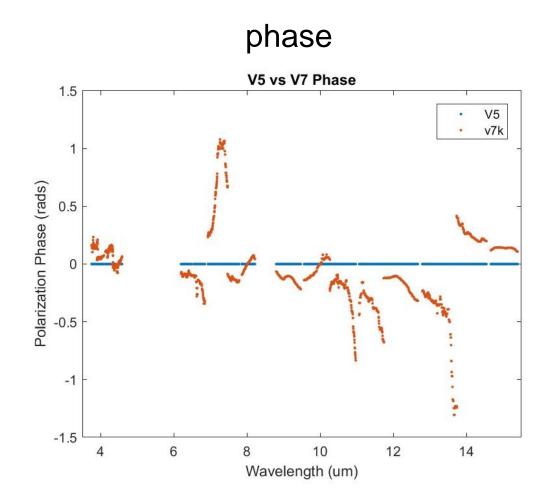




V7k Compared to V5: Polarization and Phase

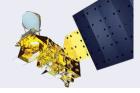


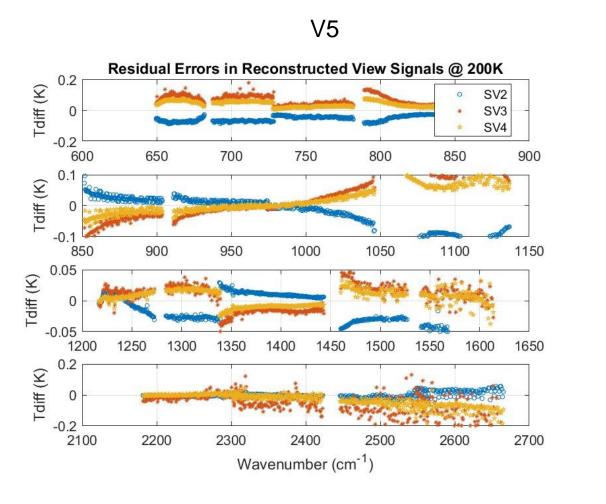


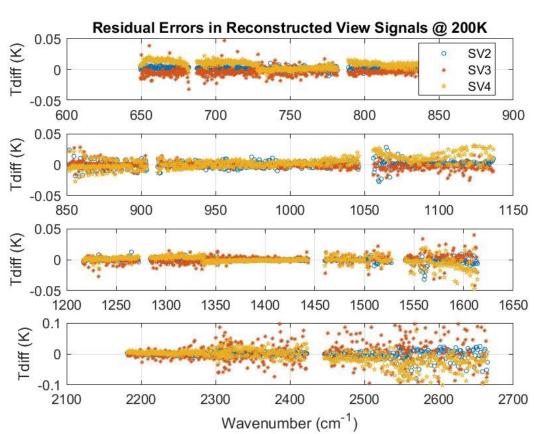




Residual fit to space views better in V7k



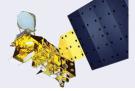




V7k

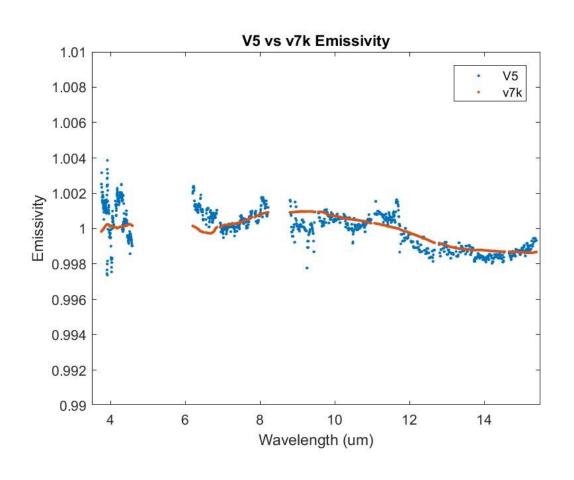


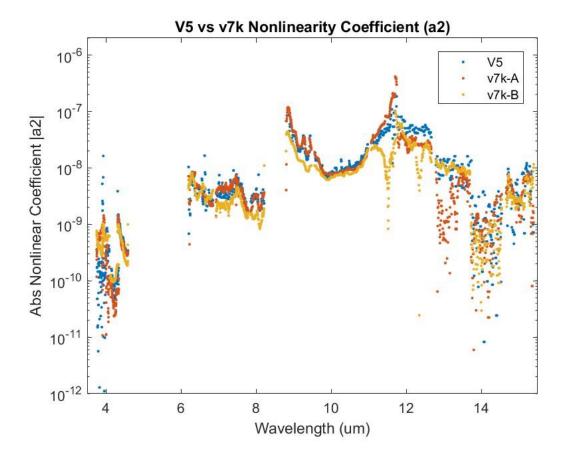
Changes Made to Emissivity and Nonlinearity



V7k Emissivity Smoothed over Channels

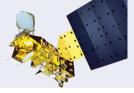
V7k Uses Separate A & B Nonlinearity

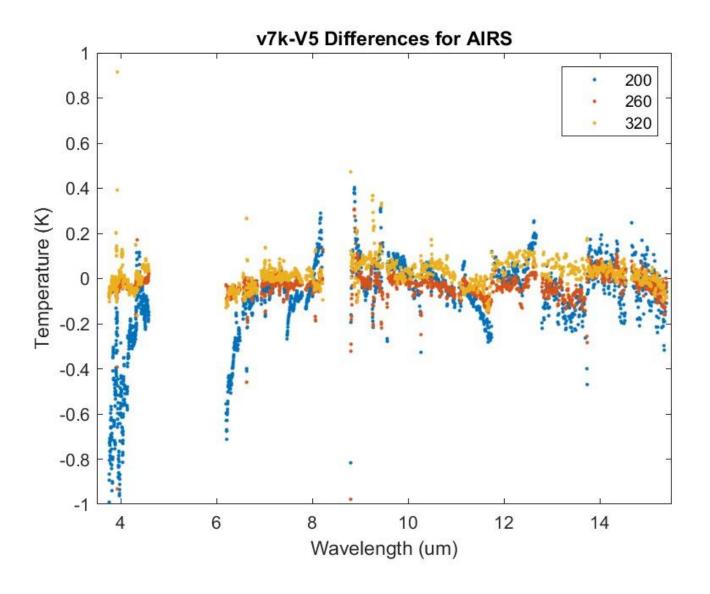






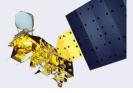
Det Propulsion Laboratory California Institute of Technology V7k Shows Biggest Differences for Cold Scenes



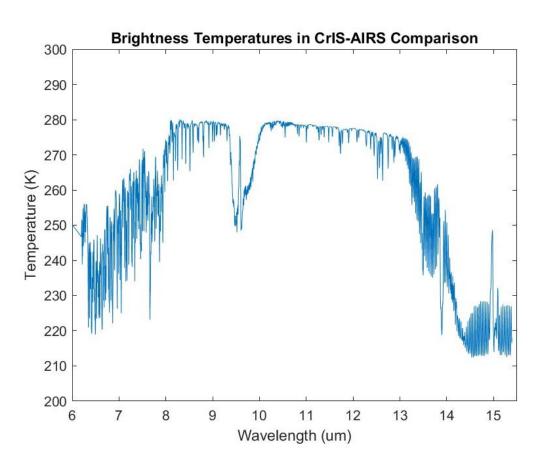




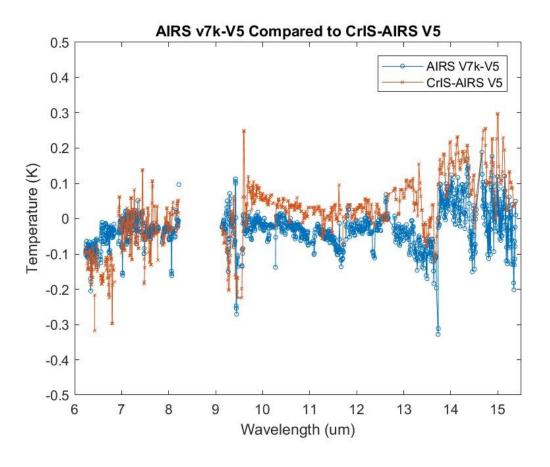
New Version Agrees Better with CrIS on SNPP



Average Scene Radiance of Difference



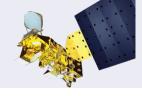
Random Sample Differences



Crosstrack Infrared Sounder (CrIS) on the Suomi NPP Satellite



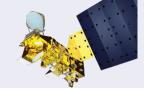
Contributors to Accuracy at 9.1µm

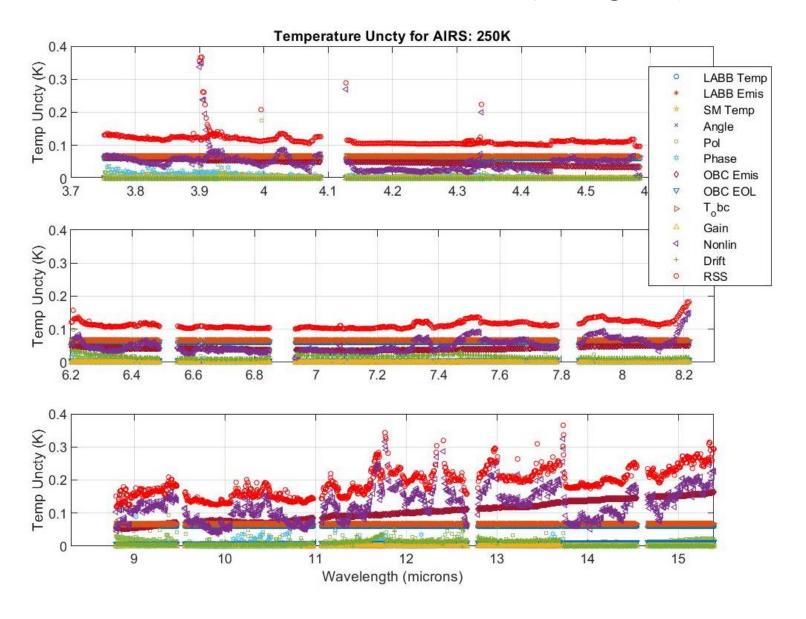


Parameter	2-σ	2-σ	2-σ	2-σ
	Parameter	Parameter	Radiometric	Radiometric
	Uncty	Uncty	Uncty (K)	Uncty (K)
Version	V5	V7k	V5	V7k
Uncertainty in LABB Temperature	0.03K	0.06K	0.030	0.060
Uncertainty in LABB Emissivity	0.0001	0.0001	0.004	0.004
Uncertainty in Scan Mirror Temperature	1.00K	1.00K	0.007	0.007
Uncertainty in Polarization Amplitude	0.0013	0.0003	0.076	0.017
Uncertainty in Polarization Phase	0.118 r	0.013 r	0.015	0.001
Uncertainty in OBC Blackbody Emissivity	0.0017	0.0015	0.065	0.059
Uncertainty in OBC Blackbody Emissivity (EOL)	0.0002	0.0002	0.008	0.008
Uncertainty in OBC Blackbody Temperature	0.05K	0.10K	0.033	0.067
Uncertainty in Nonlinearity	4.09%*	1.29%*	0.340	0.120
Uncertainty in drift in space view	0.08dn	0.08dn	0.004	0.004
Total Uncertainty at 250K			0.371	0.165



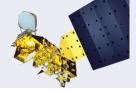
Contributors to Radiometric Uncertainty All Channels (2-Sigma)





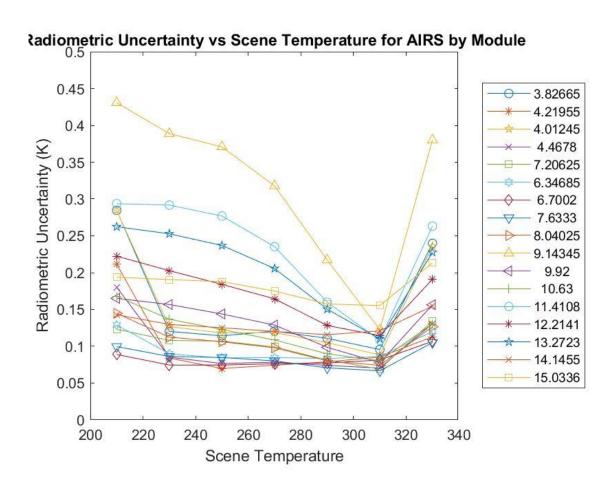


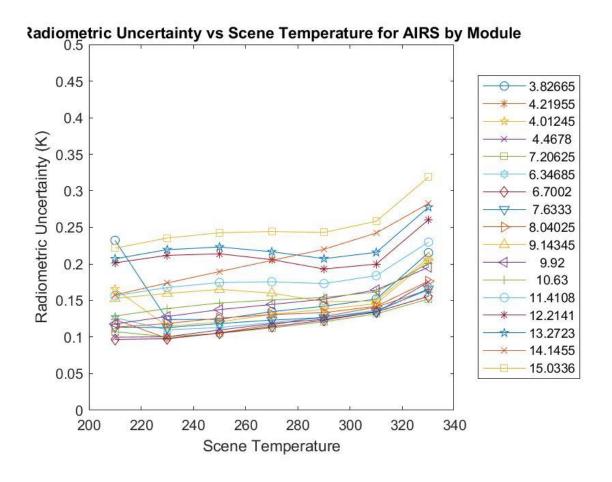
Radiometric Error vs Scene Temperature



V5

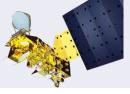
V7k







Summary and Conclusions



- AIRS designed for high accuracy and stability
- 16+ years of data now available for science investigations
- New Version of radiometric calibration coefficients developed: V7k
 - New time dependent polarization and phase
 - Smoothed emissivity
 - Separate A and B side nonlinearity
 - Better agreement with CrIS on SNPP
 - Expect Version 7k coefficients to be part of Level 1B V7 release for AIRS in 2019
- Radiometric Accuracy
 - Revisited accuracy of all contributors compared to 2008 analysis
 - New estimates relate to V7k. Compared to V5.
 - Some areas worse: Increase in LABB and OBC temperature uncertainty
 - Some areas better: Reduction in errors due to polarization
- AIRS, CrIS agree well, within the uncertainties estimated