



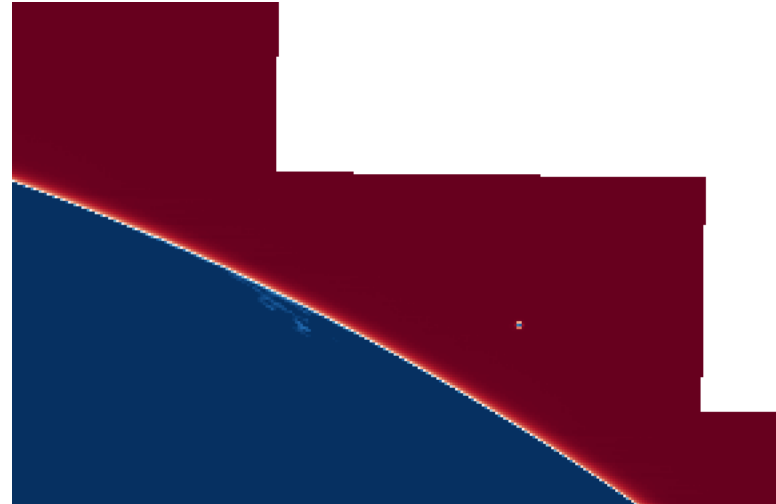
Universität Hamburg  
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**CENTRUM**  
FÜR ERDSYSTEMFORSCHUNG  
UND NACHHALTIGKEIT (CEN)

# Inter-Calibration of Infrared Imagers Using Observations of Venus

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# Agenda

- “Third Point“ for Calibration of Infrared Imagers
- What about Venus?
- First Results From Inter-Calibration
- Conclusions

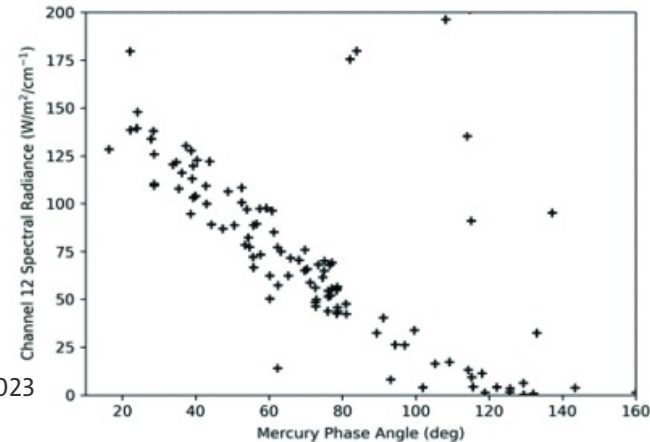
# Celestial Targets Considered Fit for Calibration

## Moon (Used With Solar Bands)

- Temperatures 100 - 400 K on surface
- Model needed close to New ☾
- Inconsistent spatiotemporal sampling in data stitching

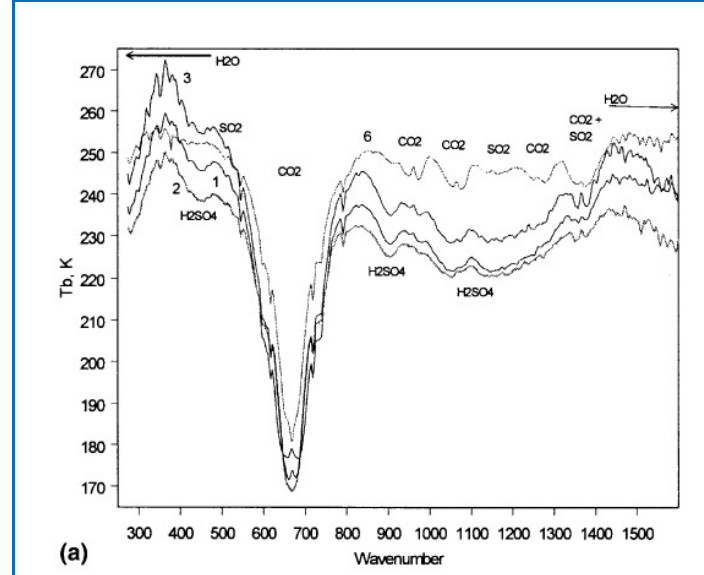
## Mercury (Bremer, SPIE, 2010)

- Large dynamic range needed
- Exact knowledge of PSF and pointing essential



# Venus: a Possible Radiometric Calibrator?

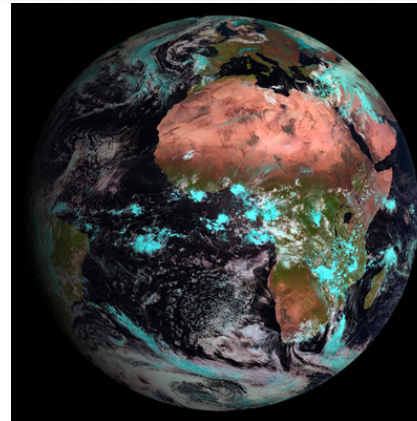
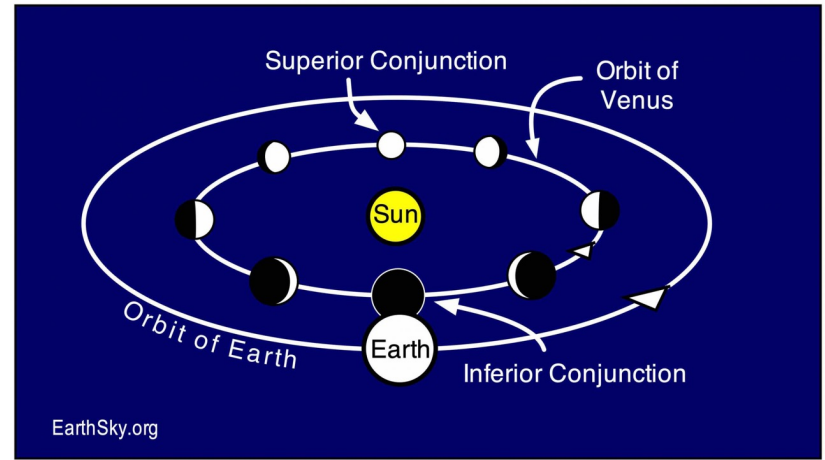
- Venus has an atmosphere, so
- Q: Its  $T_b$  is variable
- A: Reproducible diurnal variations?
- Q: There are spectral features – availability of model?
- A: No model needed for inter- and cross-calibration!
- Q: Spatial variations?
- A: Average over large fraction of disk!



Zasova et al. (2004)

# Observing Strategy: Inferior Conjunctions of Venus

Date and Time	Dec	Diameter
2007 Aug 18	+5°	60"
2009 March 27	+10°	59"
2015 Aug. 15	+6°	58"
2017 March 25	+9°	59"
2023 Aug. 13	+7°	58"
2025 Mar 23	+8°	60"



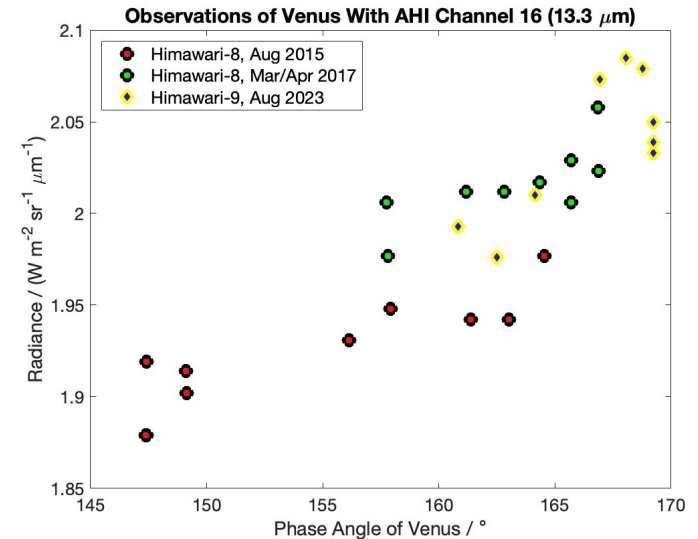
# Inter-Calibration of SEVIRI on METEOSAT-8...-11

Satellite	Date	Phase Angle	Counts (13.4 $\mu\text{m}$ )
METEOSAT-11	2023/08/13	L169°	207.4 $\pm$ 1.8
METEOSAT-10	2023/08/30	L143°	205.2 $\pm$ 1.7
METEOSAT-10	2017/03/30	L164°	203.9 $\pm$ 0.7
METEOSAT-9	2017/04/02	L159°	204.1 $\pm$ 0.9
METEOSAT-9	2015/08/21	L163°	203.5 $\pm$ 0.5
METEOSAT-8	2017/04/11	L144°	207.5 $\pm$ 1.3

Counts (raw data) change by less than 0.7% in radiance or 0.4 K in brightness temperature.

# Inter-Calibration of AHI on Himawari-8 and -9

- Example: AHI Channel 16 ( $13.3 \mu\text{m}$ )
- Phase angle range  $155^\circ - 165^\circ$  covered three times
- Himawari-8 in 2015:  $1.948 \pm 0.008 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1}$
- Himawari-8 in 2017:  $2.005 \pm 0.007 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1}$
- Himawari-9 in 2023:  $1.993 \pm 0.01 \text{ W m}^{-2} \text{ sr}^{-1} \mu\text{m}^{-1}$
- Agreement between Himawari-8 and -9
- Calibration drift of  $\approx 1 \text{ K}$  with Himawari-8 after launch?
- Radiance at  $13.3 \mu\text{m}$  with max. at local midnight?



# Summary and Conclusions

- Venus' night side: a homogeneous extended source with a radiance slightly lower than Earth
- STDEV in radiance  $< 1\%$  at  $\lambda > 8 \mu\text{m}$   $\Rightarrow$  Venus is suitable for vicarious calibration
- Spot checks of radiometric stability in different years: SEVIRI on METEOSAT-9 and -10 better than  $0.5 \text{ K}$  *on counts level*.
- Inter-calibration of Himawari-8 and -9: agreement within  $0.5\%$  in radiance or  $0.3 \text{ K}$  in brightness temperature, six years apart.
- Venus' only drawback is the low frequency of inferior conjunction

