The background features a dark blue gradient with a starry sky pattern. On the left side, there are several technical diagrams, including circular gauges with scales (ranging from 140 to 260) and various circular paths with arrows, suggesting a scientific or engineering context. The main title is centered in white, bold, sans-serif font.

CALIBRATION AND CHARACTERISATION OF MICROWAVE SOUNDERS WITH THE MOON

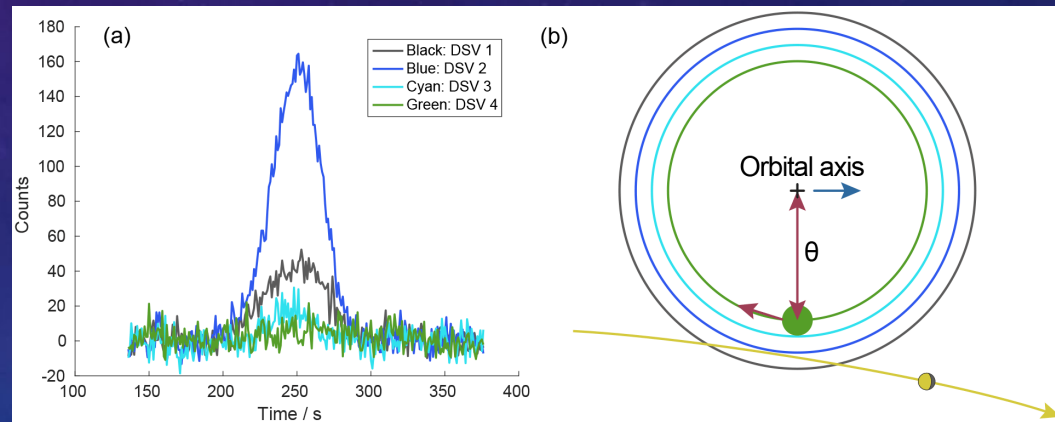
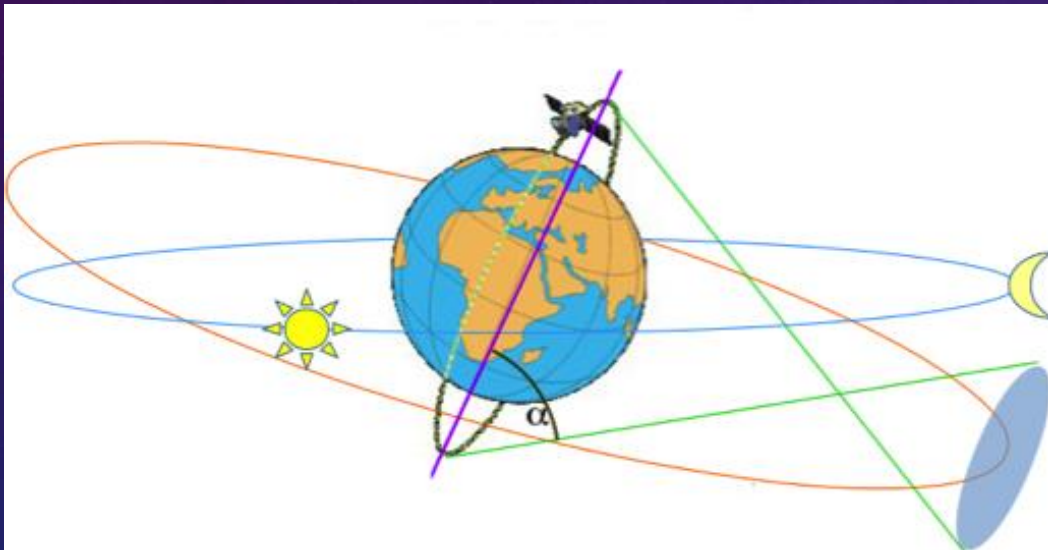
MARTIN BURGDORF

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OBSERVING THE MOON WITH AMSU-B AND MHS

Position of   

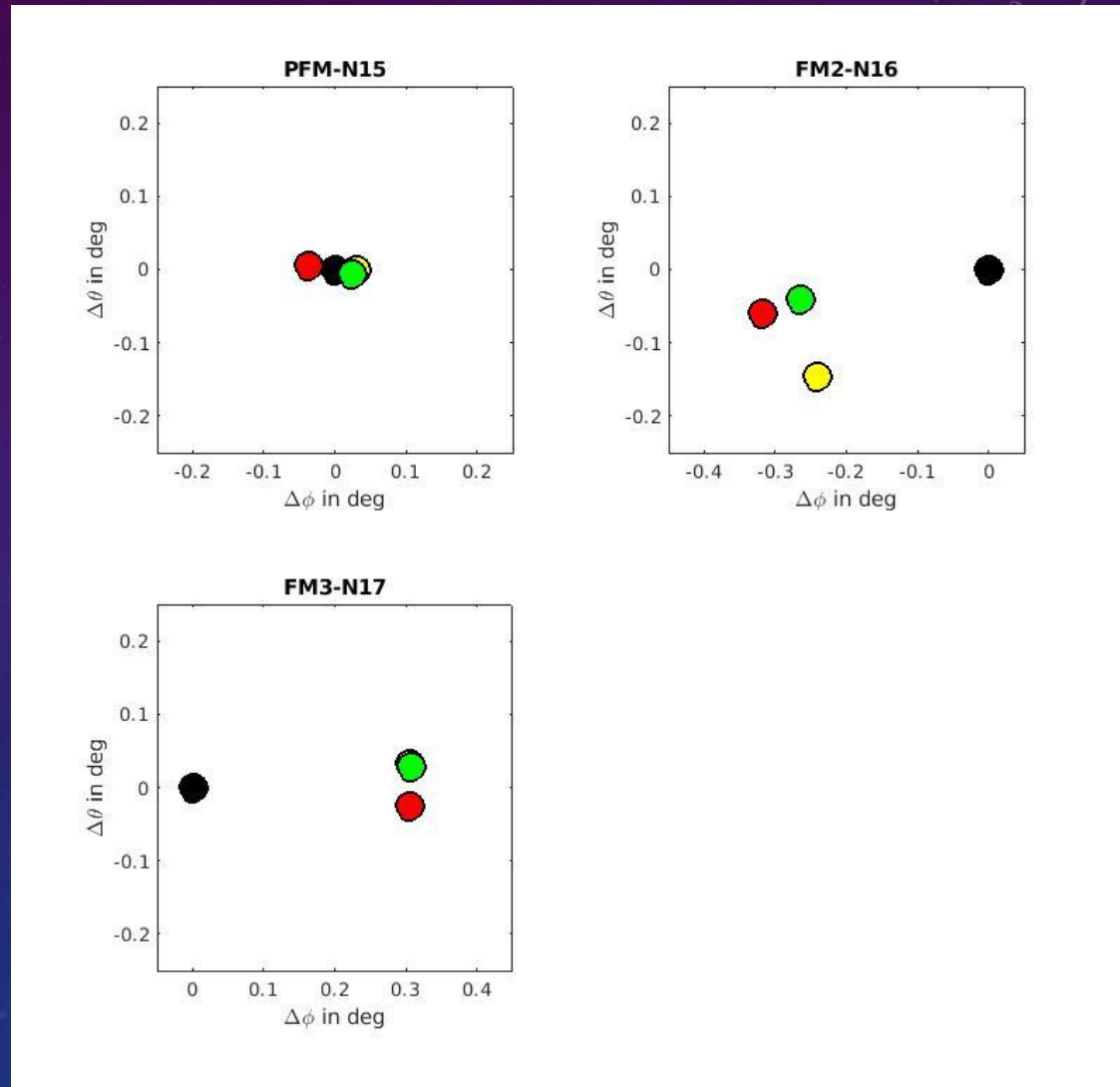
Light Curves



POINTING ERROR OF AMSU-B

It is possible to determine the pointing direction in both the along-track and the scan direction with high accuracy from the amplitude and the position of the peak in each pixel of the deep space view.

Channel 1: red, channel 2: yellow, other channels: green, nominal position: black

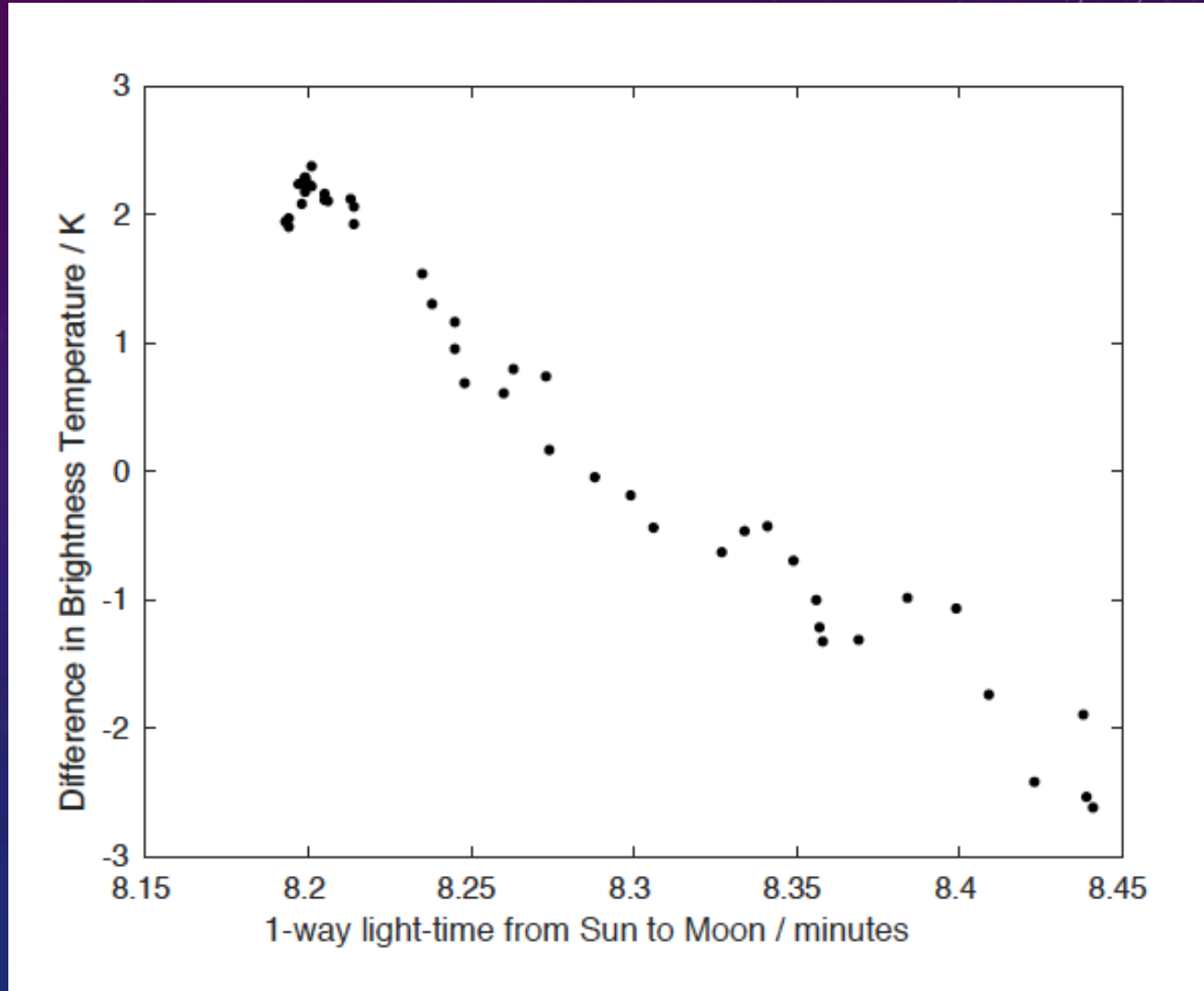


MEAN HALF POWER BEAMWIDTH

- Duration of Moon intrusion → beam size
- Check with measured radiance
- Results from ground tests could not always be reproduced in flight
- Mean half power beamwidth ⚡ requirements in some cases
- MHS on Metop-C: issues with both beam size and pointing error
- Beam diameter for sounding channels $\geq 1.21^\circ$ (except for N16 and Metop-B)

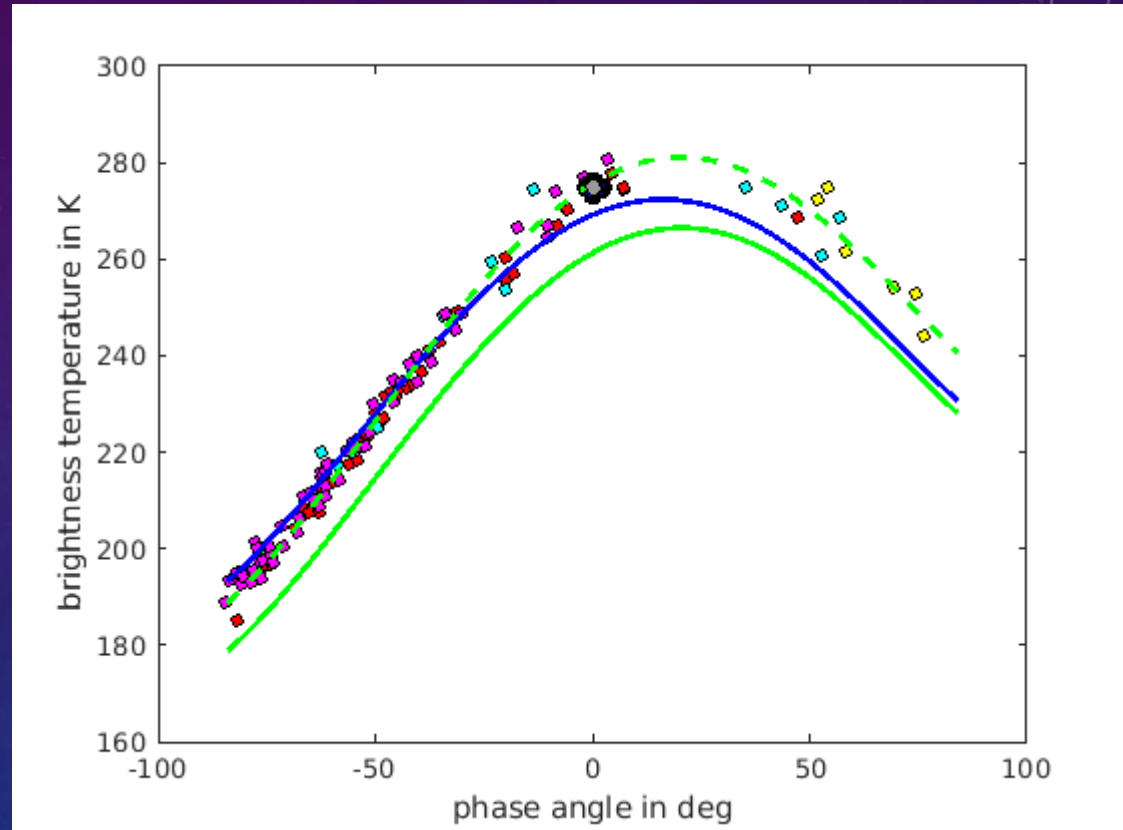
DISTANCE SUN - MOON

The disk-integrated TB is almost 5 K higher at perihelion than at aphelion. This agrees well with the value of 4.5 K by Liu.



OBSERVATION AND MODEL

The measured brightness temperatures of the Moon at 89 GHz from AMSU-B on NOAA-16 (cyan) and NOAA-17 (yellow) as well as from MHS on NOAA-18 (red) and NOAA-19 (magenta) are shown as a function of phase angle. The large, grey dot represents NOAA-20 ATMS. The solid lines give the predictions by Keihm (1984, blue) and Liu & Jin (2020, green). The dashed, green line represents the model by Liu & Jin, but scaled by a factor 1.055.



SUMMARY

- Pointing accuracy at DSV
 - Requirement: $\pm 0.1^\circ$ for AMSU-B, $\pm 0.09^\circ$ for AMSU-B
 - Not compliant in 1/3 of the cases, more than $\pm 0.3^\circ$
- Beamwidth at DSV
 - Requirement: $1.1^\circ \pm 10\%$
 - Not compliant in half of the sounding channels, discrepancies to ground tests \geq ten sigma
- Photometric calibration
 - Need to take distance of Moon to Sun and Observer and phase angle into account
 - No model agrees with measurements, but Liu's comes very close after scaling upwards by 5.5%

CONCLUSIONS

- Moon intrusions in the deep space view are helpful for characterising **in flight**
 - Pointing error of sounding channels
 - Beamwidth
 - Photometric stability
 - Agreement between satellites or channels
 - Noise and baseline variations on timescale of minutes
- **Performance differs in flight significantly from ground tests**