### GSICS Annual Meeting 17 March - 21 March 2025

## Multi-Point Geostationary Observations of Major 2024 Space Weather Events : Data from GK2A, Himawari-9, GOES-16 and -18

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### Korean Space Weather Monitor on GK2A: KSEM



eter	Specification
ange	e: 100 keV - 3.8 p: 77 keV - 6 MeV
gy tion	∆E/E ≤ 0.2%
ection	6 directions
acy	≤ 1 nT
ength	1 m
je	±3 pA/cm <sup>2</sup>

### $\leq 0.01 \text{ pA/cm}^2$

**GK2A provides multiple continuous** measurements of space weather over the



### **Multi-Point Simultaneous Observations in Space Weather Events**

### Importance of Multi-Point Simultaneous Observations and Comparative Studies in Space Weather Events

- The Earth's magnetosphere is continuously influenced by the solar wind with sudden magnetospheric disturbances.
- During extreme space weather events, the magnetopause undergoes compression and restructuring, causing GEO satellites to transition between the magnetosphere and interplanetary space
- GK2A, Himawari-9, GOES-16, and GOES-18 provide simultaneous multi-point observations from different locations, allowing for a comprehensive analysis of global magnetospheric dynamics.
- In this time, we utilizes multi-point simultaneous observations from major space weather events in May and October 2024 to compare and analyze effects of magnetopause compression on magnetic field and electron flux measurements.





### **Electron flux correlations between GK2A and GOES-16 (Day-averaged)**



• The	R		β		α		Channel	
betw	2021	2020	2021	2020	2021	2020	No.	
Scat	0.844	0.882	-0.089	-0.085	0.909	0.914	2	
a rea	0.868	0.886	-0.488	-0.246	1.080	1.019	3	
• Enor	0.875	0.880	-0.915	-0.567	1.176	1.078	4	
and	0.891	0.907	-0.874	-0.610	1.222	1.147	5	
anu	0.904	0.925	-0.884	-0.607	1.233	1.123	6	
<ul> <li>Corr</li> </ul>	0.940	0.940	-0.363	-0.249	1.069	0.998	7	
• The	0.951	0.932	-0.095	-0.074	0.791	0.750	8	
_ betw	0.902	0.911	0.971	0.889	1.168	1.121	9	

comparisons were made using the conjunction condition of  $dL^* < 0.1$ veen two satellites

ter plots of daily-averaged electron particle flux from both detectors show asonable correlation for most channels except for Channel 9.

rgy based spectral distributions show good agreement between KSEM PD MPS-HI with similar power law distributions, except for Channel 9

relation coefficients obtained from  $\log[Flux_{KSEM PD}] = \alpha \log[Flux_{MPS-HI}] + \beta$ .

Pearson's R-values are ~0.85 or greater, indicating a strong correlation veen the two detectors' data.







### 2 Major Space Weather Events in 2024: May and October



#### May 9-12, 2024

- AR 3664 exhibited extreme activity, producing multiple X-class solar flares and Earthward CMEs.
- More than 6 CMEs merged into a single structure before reaching Earth.
- The noon-side magnetopause contracted to 4.66 RE (based on the Lin model, 10minute averaged data).
- The Kp index peaked at 9, and the Dst index dropped to -412 nT.
- The largest geomagnetic storm since the 2003 Halloween event.



#### October 9-11, 2024

- AR 3848 became highly active, producing multiple X-class solar flares and Earthdirected CMEs.
- Geomagnetic disturbances persisted due to the combined effects of high-speed solar wind and successive CME impacts.
- The dayside magnetopause contracted to 5.05 RE (based on the Lin model, 10minute averaged data).
- The Kp index reached 9-, and the Dst index dropped to -341 nT.



## May 2024 Case 1 (2024-05-10 17:00 - 18:00)



- Exposure to Interplanetary Space: GOES-16 was exposed before GOES-18.
- Exposure to Interplanetary Space: Not exposed. • Electron Flux Changes: GOES-16, being closer to the noon sector, showed a rapid decrease, while GOES-18 exhibited a slower decline. • Electron Flux Changes: Gradual and slight decrease.
- Magnetic Field Changes: Sudden and strong Bz reversal (w/ a shorter SSC) • Magnetic Field Changes: GK2A's Bz decreased but did not reverse duration for GOES-16). direction.
  - SSC Onset Time:
    - GOES-16: 2024-05-10 17:05
    - **GOES-18**: 2024-05-10 17:08 (or 17:06?)

### **GK2A**, Himawari-9 (Night)

SSC Onset Time: ?

## May 2024 Case 2 (2024-05-11 00:00 - 01:00)



#### GOES-16 (Night)

- Exposure to Interplanetary Space: Not exposed.
- Electron Flux Changes: No significant decrease.
- Magnetic Field Changes: No clear decrease or reversal in Bz.

#### **GOES-18 (Early Afternoon)**

- Exposure to Interplanetary Space: Exposed.
- Electron Flux Changes: Sharp decrease.
- Magnetic Field Changes: Bz showed repeated

- SSC Onset Time: 2024-05-11 00:16
- reversals, indicating instability  $\rightarrow$  Possibly located near the magnetopause boundary?

### **GK2A**, Himawari-9 (Morning)

- Exposure to Interplanetary Space: Exposed.
- Electron Flux Changes: Gradual decrease (likely due to prior multiple exposures).
- Magnetic Field Changes: GK2A recorded a sharp Bz decrease and reversal.
  - SSC Onset Time: ?





## May 2024 Case 3 (2024-05-11 06:00 - 07:00)



### GOES-16, GOES-18 (Night)

- Exposure to Interplanetary Space: Not exposed.
- Electron Flux Changes: No significant additional decrease, as most flux had **Electron Flux Changes**: No significant additional decrease, as flux had already been lost. already been largely depleted.
- Magnetic Field Changes: No noticeable changes.

#### **GK2A**, Himawari-9 (Early Afternoon)

• Exposure to Interplanetary Space: Exposed.

- Magnetic Field Changes: GK2A recorded a sharp Bz decrease and reversal.
  - SSC Onset Time: 2024-05-11 06:10
  - Shortly after the decrease, Bz fluctuated significantly  $\rightarrow$  Possibly indicating that the magnetopause boundary remained near GK2A.





## October 2024 Case 1 (2024-10-10 15:00 - 16:00)



### GOES-16 (Near Noon), GOES-18 (Morning)

- Exposure to Interplanetary Space: GOES-16 was exposed before GOES-18. GK2A, Himawari-9 (Night)
- Electron Flux Changes: GOES-16, being closer to the noon, showed a rapid • Exposure to Interplanetary Space: Not exposed. decrease, while GOES-18 exhibited a relatively slower decline. • Electron Flux Changes: Temporary rapid increase after magnetopause
- Magnetic Field Changes: Sudden and strong Bz reversal (w/ a shorter SSC contraction  $\rightarrow$  Possible temporary trapping of electrons in the tail? duration for GOES-16).
  - SSC Onset Time:
    - GOES-16: 2024-10-10 15:14
    - GOES-18: 2024-10-10 15:15

- Magnetic Field Changes: GK2A's Bz was already low but decreased slightly without reversal.
  - SSC Onset Time: ?



### October 2024 Case 2 (2024-10-10 19:00 - 20:00)



### **GOES-16 (Early Afternoon)**

- Exposure to Interplanetary Space: Briefly exposed or not exposed.
- Electron Flux Changes: Gradual decrease.
- Magnetic Field Changes: Slight Bz decrease, weak or no reversal.

#### **GOES-18 (Near Noon)**

- Exposure to Interplanetary Space: Exposed.
- Electron Flux Changes: Sharp decrease.
- Magnetic Field Changes: Bz decreased and reversed.
  - SSC Onset Time: 2024-10-10 19:37

### GK2A, Himawari-9 (Dawn)

- Exposure to Interplanetary Space: Not exposed.
- Electron Flux Changes: Gradual and minor decrease.
- Magnetic Field Changes: GK2A's Bz decreased slightly but did not reverse.
  - SSC Onset Time: ?





### October 2024 Case 3 (2024-10-10 22:00 - 23:00)



### **GOES-16 (Evening)**

- Exposure to Interplanetary Space: Briefly exposed.
- Electron Flux Changes: Sharp decrease.
- Magnetic Field Changes: Bz decreased and temporarily reversed.
  - SSC Onset Time: 2024-10-10 22:34

#### **GOES-18 (Near Noon)**

- Exposure to Interplanetary Space: Exposed.
- Electron Flux Changes: Sharp decrease.
- Magnetic Field Changes: Bz decreased and reversed.

SSC Onset Time: 2024-10-10 22:31

### **GK2A**, Himawari-9 (Morning)

- Exposure to Interplanetary Space: Exposed.
- Electron Flux Changes: Sharp decrease.
- Magnetic Field Changes: GK2A's Bz decreased sharply and reversed.
  - SSC Onset Time: 2024-10-10 22:33





### **Overview of Observed May Cases**



May Case 1



May Case 3

### May 2024

Date & Time	Satellite	Exposure to Interplanetary Space	<b>Electron Flux Changes</b>	Magnetic Field Changes	SSC Onset
2024-05-10 17:00 - 18:00	GOES-16 (Near Noon)	Exposed before GOES-18	Sharp decrease	Bz decreased and reversed	2024-05-10
	GOES-18 (Morning)	Exposed after GOES-16	Gradual decrease	Bz decreased and reversed	2024-05-10
	GK2A, Himawari-9 (Night)	Not exposed	Gradual decrease	Bz slightly decreased, no reversal	?
2024-05-11 00:00 - 01:00	GOES-16 (Night)	Not exposed	No significant change	No significant change	?
	GOES-18 (Early Afternoon)	Exposed	Sharp decrease	Bz reversed and unstable	2024-05-11
	GK2A, Himawari-9 (Morning)	Exposed	Gradual decrease	Bz decreased and reversed	?
2024-05-11 06:00 - 07:00	GOES-16 (Night)	Not exposed	No significant change	No significant change	?
	GOES-18 (Night)	Not exposed	No significant change	No significant change	?
	GK2A, Himawari-9 (Evening)	Exposed	Sharp decrease	Bz decreased and reversed	2024-05-11

May Case 1: GOES-16 and GOES-18 (day) were exposed to interplanetary space, showing sharp electron flux drops and Bz reversals, while GK2A and Himawari-9 (night) recorded gradual flux decreases without direct exposure.

May Case 2: GOES-18 (evening) showed unstable Bz, suggesting the magnetopause boundary remained near the satellite, while GK2A and Himawari-9 (day), moving toward noon, experienced moderate flux decreases and Bz reversals. May Case 3: GOES-16 and GOES-18 (night) showed minimal changes, while GK2A and Himawari-9 (evening) recorded strong Bz reversals and electron flux losses.









### **Overview of Observed OctoberCases**



**October Case 1** 



**October Case 3** 

#### October 2024

Date & Time	Satellite	Exposure to Interplanetary Space	<b>Electron Flux Changes</b>	Magnetic Field Changes	SSC Onset
2024-10-10 15:00 - 16:00	GOES-16 (Near Noon)	Exposed before GOES-18	Sharp decrease	Bz decreased and reversed	2024-10-10
	GOES-18 (Morning)	Exposed after GOES-16	Gradual decrease	Bz decreased and reversed	2024-10-10
	GK2A, Himawari-9 (Night)	Not exposed	Temporary increase	Bz decreased, no reversal	?
2024-10-10 19:00 - 20:00	GOES-16 (Early Afternoon)	Briefly exposed or not exposed	Gradual decrease	Bz decrease, weak or no reversal	?
	GOES-18 (Near Noon)	Exposed	Sharp decrease	Bz decreased and reversed	2024-10-10
	GK2A, Himawari-9 (Dawn)	Not exposed	Gradual and minor decrease	Bz decreased, no reversal	?
2024-10-10 22:00 - 23:00	GOES-16 (Evening)	Briefly exposed	Sharp decrease	Bz decreased and temporarily reversed	2024-10-10
	GOES-18 (Near Noon)	Exposed	Sharp decrease	Bz decreased and reversed	2024-10-10
	GK2A, Himawari-9 (Morning)	Exposed	Sharp decrease	Bz decreased and reversed	2024-10-10

October Case 1: GOES-16 and GOES-18 (noon) were exposed, showing sharp flux drops and Bz reversals, while GK2A and Himawari-9 (night) briefly saw flux increases, possibly from temporary particle trapping.

October Case 2: GOES-18 (noon) was exposed, showing flux drops and Bz reversals, while GOES-16 (afternoon) had weaker responses. GK2A and Himawari-9 (dawn) saw gradual flux decreases, indicating a weaker impact.

October Case 3: GOES-18 (noon) was exposed, showing strong flux drops and Bz reversals, while GOES-16 (evening) had brief exposure. GK2A and Himawari-9 (morning) saw electron flux losses and strong Bz changes, reflecting magnetospheric evolution.









### Summary

### **Multi-Point Observations in Geostationary Orbit for Space Weather Research**

- May and October 2024.
- space, showing sharp electron flux drops and Bz reversals.
- fluxes due to magnetospheric deformation.
- magnetospheric responses in space weather events.
- validation of space weather data are essential.

•Using GK2A, Himawari-9, GOES-16 and -18, we analyzed major space weather events from

•During magnetopause compression, satellites in the dayside were exposed to interplanetary

•Satellites in the nightside were not exposed but showed gradual changes in Bz and electron

•This shows the value of multi-point observations in providing a more comprehensive view of

•<u>To maximize the effectiveness of multi-point observations, continuous cross-calibration and</u>

•Well-calibrated data enhances the synergy of multi-point observations, improves space weather modeling, increases satellite operation stability, and strengthens forecasting capabilities.



# Thank you



