



# Hyperspectral CoSMIR-H Aircraft Instrument and the WH<sup>2</sup>yMSIE Campaign

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GSICS Annual Meeting

19 March 2025



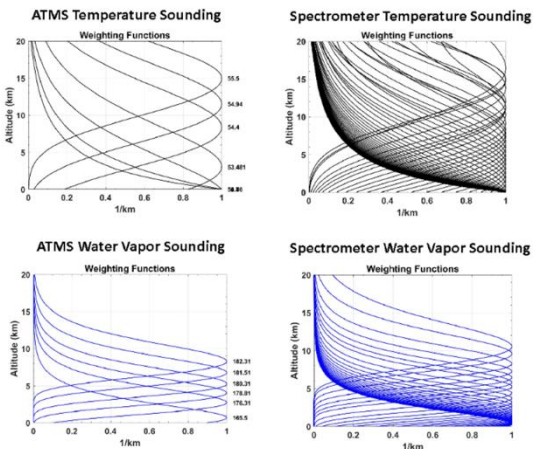
TOWARD A GLOBAL PLANETARY  
BOUNDARY LAYER OBSERVING SYSTEM

THE NASA PBL INCUBATION STUDY TEAM REPORT



João Teixeira <sup>(1)</sup>, Jeffrey R. Piepmeier <sup>(2)</sup>, Amin R. Nehrir <sup>(3)</sup>, Chi O. Ao <sup>(1)</sup>, Shuyi S. Chen <sup>(4)</sup>, Carol A. Clayson <sup>(5)</sup>, Ann M. Fridlind <sup>(6)</sup>, Matthew Lebsock <sup>(1)</sup>, Will McCarty <sup>(2)</sup>, Haydee Salmun <sup>(7)</sup>, Joseph A. Santanello <sup>(2)</sup>, David D. Turner <sup>(8)</sup>, Zhien Wang <sup>(9)</sup>, Xubin Zeng <sup>(10)</sup>

<https://science.nasa.gov/earth-science/decadal-pbl>



## Motivation:

Planetary Boundary Layer (PBL) Study Team Report lists hyperspectral microwave (HMW) sensors as one of the **“Essential Components”** of a future global PBL observing system, to provide **“accurate PBL and free tropospheric three-dimensional (3D) temperature and water vapor structure context”**

## Objective:

- Build an airborne HMW sensor and collect observations to demonstrate capability of HMW sounding, especially in the PBL.
- Enhance the airborne CoSMIR instrument with hyperspectral receivers (renamed CoSMIR-H) by utilizing ASIC spectrometers to give **full spectrum coverage at 50-58 GHz and 175-191 GHz** and conduct flights to collect data.

## Sampling of Papers: 2010-2015

IEEE TRANSACTIONS ON GEOSCIENCE AND REMOTE SENSING, VOL. 49, NO. 1, JANUARY 2011

### Hyperspectral Microwave Atmospheric Sounding

William J. Blackwell, *Senior Member, IEEE*, Laura J. Bickmeier, R. Vincent Leslie, Michael L. Pieper, Jenna E. Samra, Chinnawat Surussavadee, *Member, IEEE*, and Carolyn A. Upham

### Benefits of a Hyperspectral Microwave Sensor

*Applications in Environmental Monitoring and Weather Forecasting*

Sid-Ahmed Boukabara  
NOAA/NESDIS/STAR  
Camp Springs, MD, USA  
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Kevin Garrett  
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Camp Springs, MD, USA  
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### Journal of Geophysical Research: Atmospheres

**RESEARCH ARTICLE** Microwave hyperspectral measurements for temperature and humidity atmospheric profiling from satellite: The clear-sky case  
10.1002/2015JD023331

**Key Points:**  
• A hyperspectral MW instrument could improve temperature & humidity retrieval compared to MetOp-SG  
• The main impact from HYMS comes from higher resolution in the O<sub>2</sub> band

Filipe Aires<sup>1,2,3</sup>, Catherine Prigent<sup>1,2</sup>, Emiliano Orlandi<sup>4</sup>, Mathias Milz<sup>5</sup>, Patrick Eriksson<sup>6</sup>, Susanne Crewell<sup>4</sup>, Chung-Chi Lin<sup>7</sup>, and Ville Kangas<sup>7</sup>

### Information content on temperature and water vapour from a hyper-spectral microwave sensor

J.-F. Mahfouf,<sup>a\*</sup> C. Birman,<sup>a</sup> F. Aires,<sup>b,c</sup> C. Prigent,<sup>c</sup> E. Orlandi<sup>d</sup> and M. Milz<sup>e</sup>

## NOAA BAA 2022



## BROAD AGENCY ANNOUNCEMENT: Demonstrating the Hyperspectral Microwave Sensor (HyMS) and Assessing the Benefits for NOAA/NESDIS

**WE2.R18: Hyperspectral Microwave Sounder Science and Technology**

Wed, 19 Jul, 10:15 - 11:30 Pacific Time (UTC -7)

**Location:** Room 18  
**Session Type:** Oral  
**Session Co-Chairs:** Jeffrey Piepmeier, NASA and William Blackwell, MIT Lincoln Laboratory  
**Track:** Community-Contributed Sessions

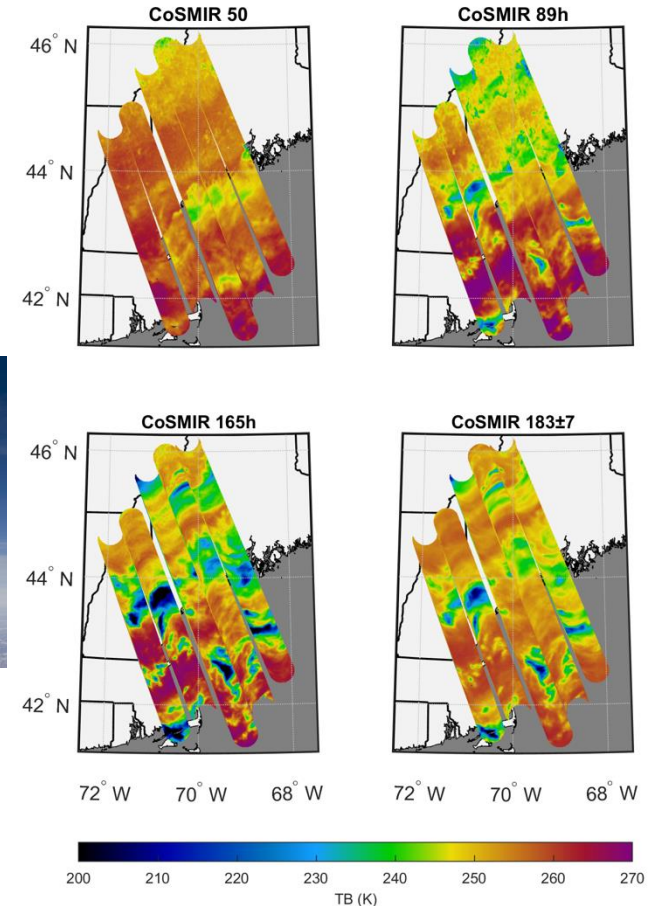
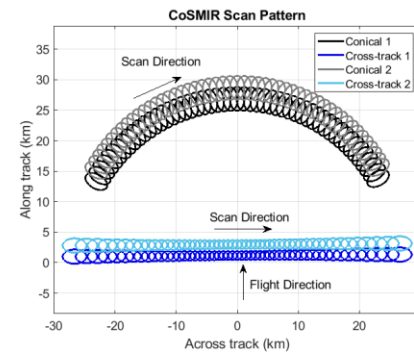
## IGARSS 2023 Special Session

How many channels are needed to be considered “hyperspectral”?

- Infrared Atmospheric Sounding Interferometer (IASI) = 8461 channels
- Microwave = >100 channels

- Programmable scanning airborne radiometer with frequencies from 50 to 183 GHz
- Originally built in the early 2000s for SSMIS cal/val (2004-2005)
- Modified in late 2000s to have frequencies similar to GMI for GPM GV
  - Flew in four GPM GV campaigns (2011-2015)
- Recently flew in IMPACTS (2020/2022)
- Significant upgrade to the receivers in 2023 to become CoSMIR-Hyperspectral

CoSMIR (1)	CoSMIR (2)	GMI
50.3h	50.3h	--
52.8h	52.8h	--
53.6h	--	--
91.665v/h	89.0v/h	89.0v/h
150h	165.5v/h	166v/h
183.31±1h	183.31±1h	--
183.31±3h	183.31±3h	183.31±3v
183.31±6.6h	183.31±7h	183.31±7v



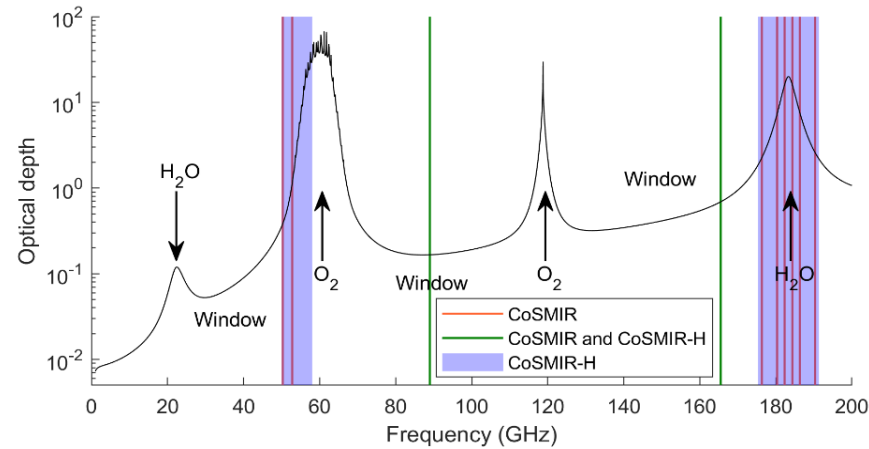
CoSMIR TB from 4 Feb 2022 IMPACTS flight Rain/snow system over the Northeast USA



CoSMIR has the unique ability to scan in various modes: fore/aft conical and cross/along-track (right).

Modified the CoSMIR 50-GHz and 183-GHz receivers to achieve full spectrum (4 MHz resolution) observations at 50-58 GHz ( $O_2$ ) and 175-191 GHz ( $H_2O$ ) using six PMCC ASIC spectrometers

ATMS	CoSMIR-H
23.8	--
31.4	--
50-58 (13 channels)	50-58h (~2k channels)
88.2	89v/h
165.5	165.3v/h
183±7-183±1 (5 channels)	175.3-191.3v (~4k channels)



- New receivers
  - 50-GHz and 183-GHz major updates with ASIC spectrometers
  - Used opportunity to update 89- and 165-GHz receivers
- Mechanical structure remained the same
  - Calibration targets and elevation/azimuth axes
  - Reduces cost associated with building a new instrument
- Data System
  - Major updates to handle large volume of data
  - CoSMIR-H is ~3 MB/s vs 1.8 kB/s for CoSMIR
- Software
  - New ground processing software required
  - Combine 4-MHz resolution into channels for science

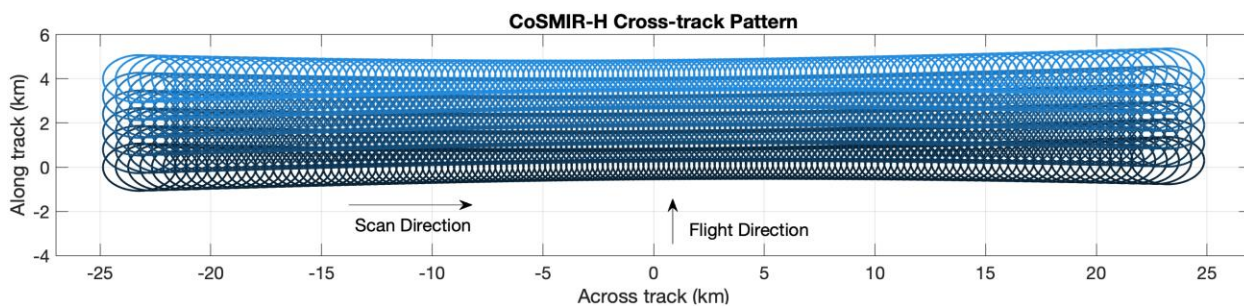
Channel	Center Frequency (GHz)	Bandwidth (MHz)	Spectral Band (GHz)	Approx. NEDT (K) (70 ms int time)
50-58h	50.002, 50.006,... 57.994, 57.998	3.906	50-50.004, 50.004-50.008,... 57.992-57.996, 57.996-58	1.5
175.3-191.3v	175.302, 175.306,... 191.294, 191.298	3.906	175.3-175.304, 175.304-175.308,... 191.292-191.296, 191.296-191.3	3.5
89v	89.0±1.25	1500 (x2)	87.0-88.5 and 89.5-91.0	0.1
89h	89.0±1.25	1500 (x2)	87.0-88.5 and 89.5-91.0	0.1
165v	165.3±1.25	1500 (x2)	163.3-164.8 and 165.8-167.3	0.5
165h	165.3±1.25	1500 (x2)	163.3-164.8 and 165.8-167.3	0.6

## 6148 total channels

### Channel set order in the L1 data product

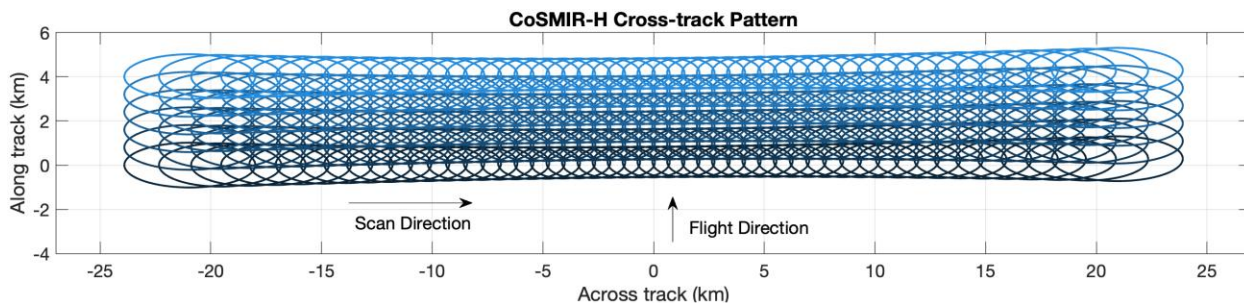
- 1-2048: 50-58 GHz
- 2049-6144: 175.3-191.3 GHz
- 6145-6148: 89v, 89h, 165v, 165h

- Cross-track scan: Scanning left to right (relative to forward direction of aircraft),  $-48.2^\circ$  to  $48.2^\circ$  elevation angle (NOT mixed polarization)
- Half-power beamwidth:  $\sim 4.5^\circ$  at 50-58 GHz,  $\sim 4.0^\circ$  at 89/165/183 GHz
- Nadir footprint: 1.6x1.6 km (50-58 GHz) and 1.4x1.4 km (89-183 GHz) for ER-2 (20 km altitude)
- Off-nadir footprint at  $48.2^\circ$ : 2.4x3.6 km (50-58 GHz) and 2.1x3.2 km (89-183 GHz) for ER-2 (20 km altitude)



(Top) Ground track pattern for 151 scan positions at 10 ms integration time.

(Bottom) Ground track pattern for 51 scan positions, averaging 7 pixels to give 70 ms integration time. This is the scan geometry included in the L1 files.

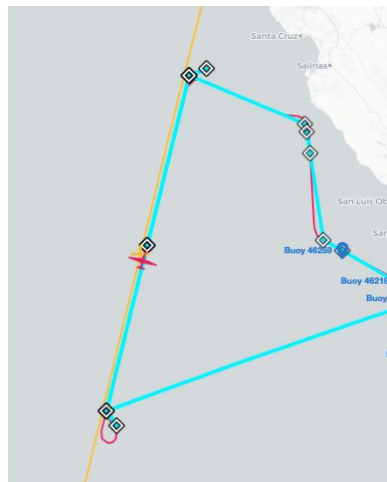


New footprint size with 70 ms averaging ( $4.0^\circ$  beamwidth)

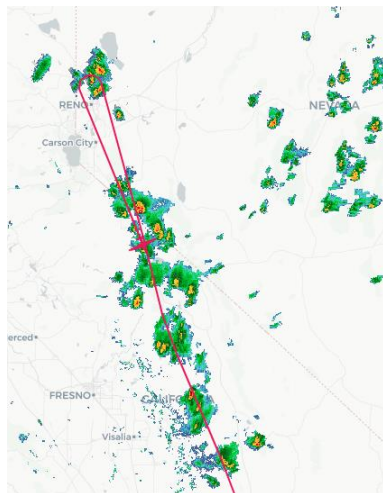
Nadir: 1.4x2.8 km

Off-nadir: 2.1x6.4 km

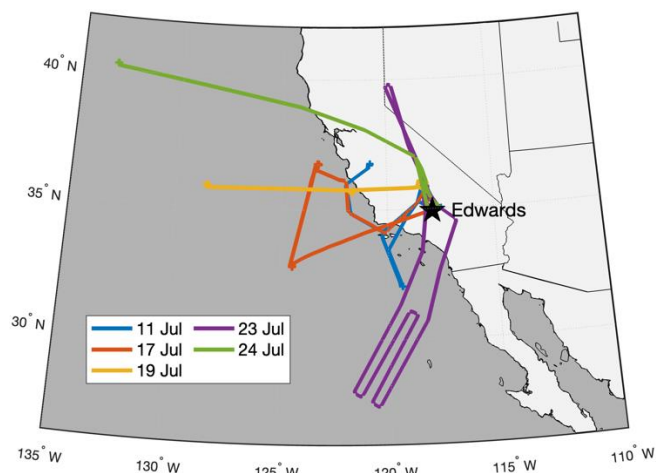
- Successful engineering check flights on the ER-2 in July 2024
  - CoSMIR-H and MBARS flew together
    - MBARS: Microwave Barometric Radar, PI: Matt Walker-McLinden
  - Five total flights, ~19 hours of observations
- Preparations for the WH<sup>2</sup>yMSIE campaign
  - Issues noticed during the check flights allowed us to further optimize CoSMIR-H prior to flying again in WH<sup>2</sup>yMSIE



MetOp-B Overpass



Convection Observations



Flight Tracks



(Top) CoSMIR-H and MBARS teams perform instrument check outs prior to first flight.  
(Bottom) ER-2 lands at Edwards AFB after a flight.  
(Credit: NASA AFRC)



## Westcoast and Heartland Hyperspectral Microwave Sensor Intensive Experiment

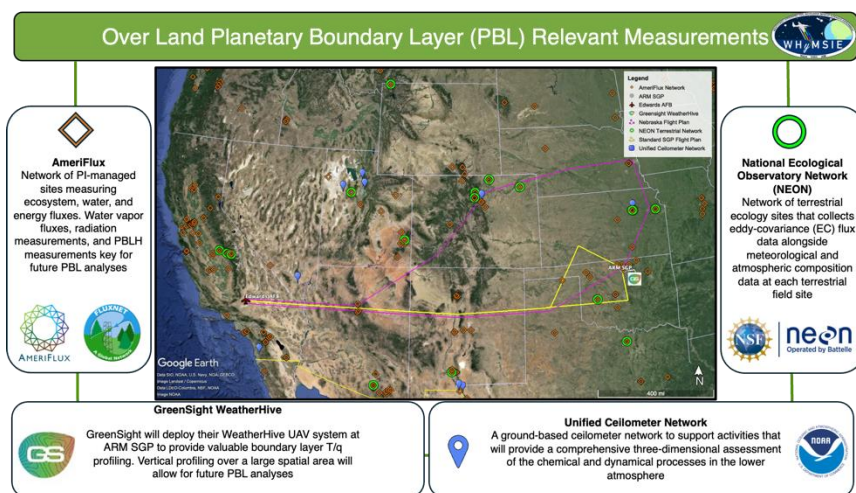
- NOAA Joint Ventures Partnership program to collect an extended CoSMIR-H dataset over a variety of atmospheric and surface conditions in Oct-Nov 2024
- WH<sup>2</sup>yMSIE also included several complementary sensors funded by NASA on the ER-2 and the G-III aircraft to demonstrate the benefits of a remote sensing instrument suite for measuring the PBL

### ER-2 payload

Hyperspectral microwave (CoSMIR-H)  
 Hyperspectral infrared (NAST-I and S-HIS)  
 Low frequency microwave radiometer (AMPR)  
 Visible/infrared imager (MASTER)  
 Backscatter lidar (CPL)  
 W-band radar (CRS) or pressure radar (MBARS)  
 GPS Radio Occultation (ARO)

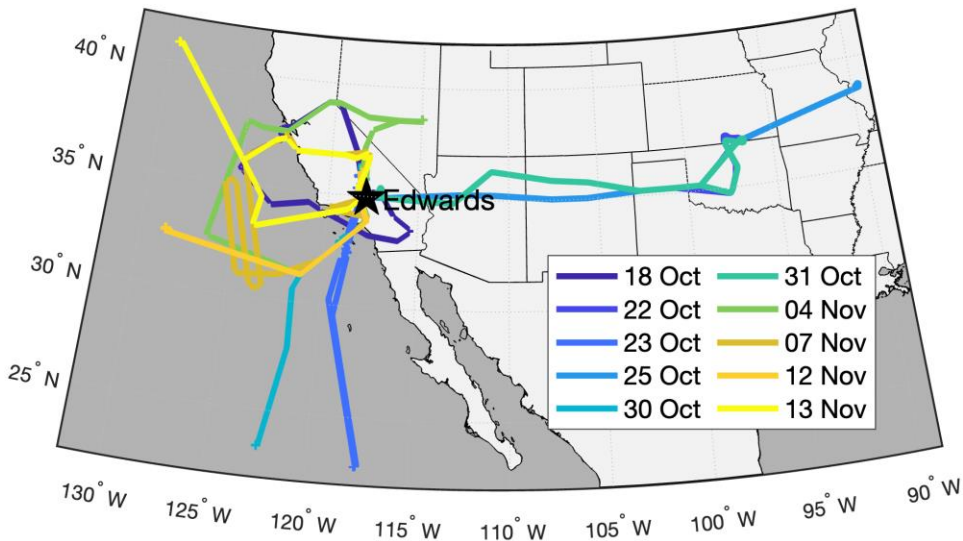
### G-III payload

Differential absorption lidar (HALO)  
 Doppler wind lidar (AWP)  
 Dropsondes



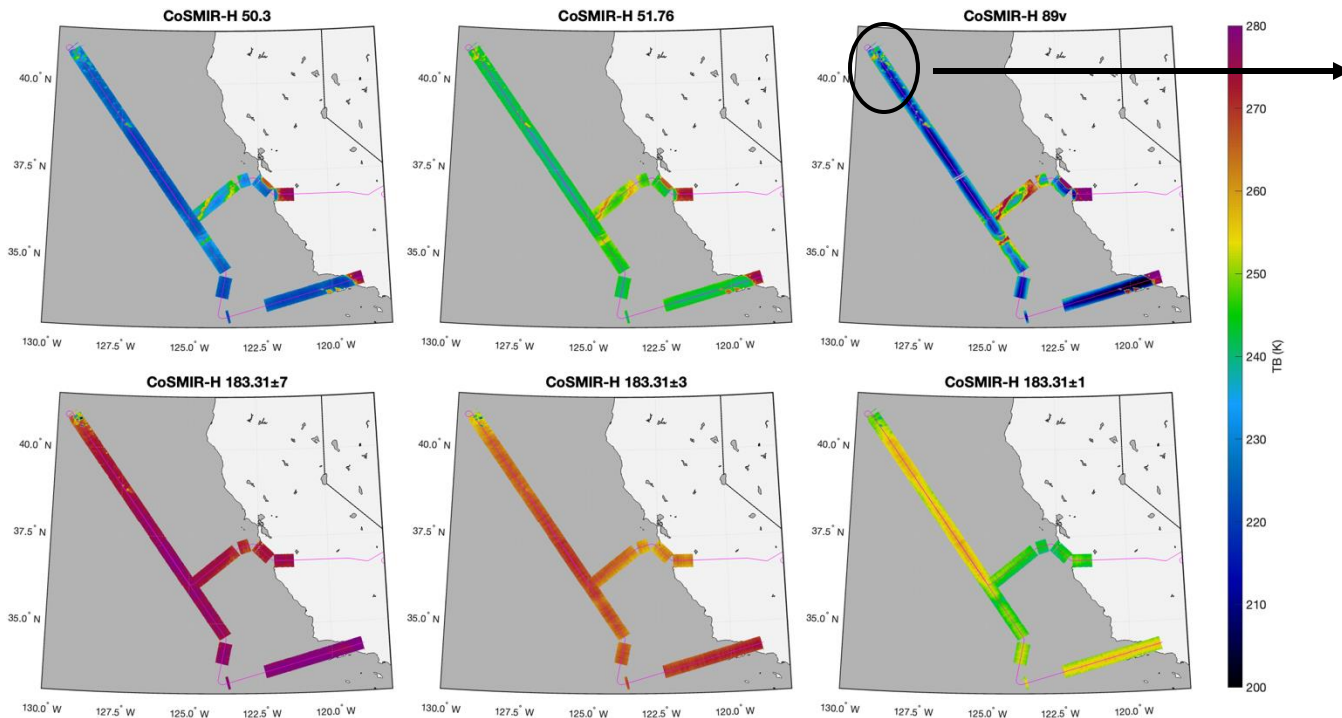
<https://earth.gsfc.nasa.gov/climate/campaigns/WHyMSIE>

- Ten total flights, ~43.5 hours of CoSMIR-H observations (~54 flight hours)
- Mix of over-land and over-ocean
- Consisted of mostly clear sky observations
- G-III joined the ER-2 for second half of the campaign

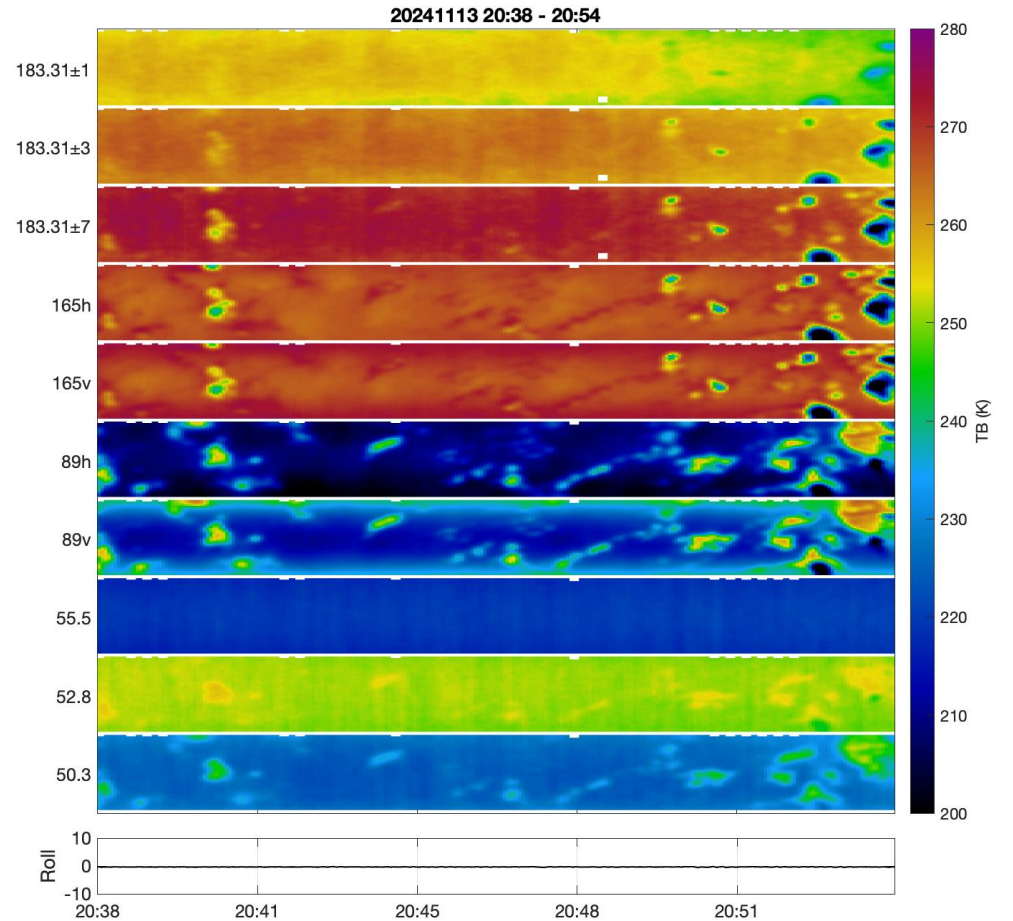


Flight Date	Approx. hours CoSMIR-H data	Observations/Location
10/18	2:50	Clear-sky over ocean and land. Flew over Lake Tahoe and Salton Sea.
10/22	5:13	Over-land out and back flight to ARM SGP. Flew over several NWS radiosonde sites and NSSL. Clear sky except some high clouds over New Mexico.
10/23	4:03	Clear-sky over-ocean with a NOAA21 overpass at 21:05 UTC. Also close to the AWS ground-track for an overpass at 18:52 UTC.
10/25	6:41	Over-land out and back flight to Scott AFB with overflights of several NWS radiosonde sites and ARM SGP. Clear sky except for very low clouds over Scott.
10/30	4:09	Over-ocean coordinated G-III flight with an EarthCARE overpass at 22:20 UTC. Clear-ish sky on the northern part of the flight path, very low puffy clouds (<2 km cloud top) on the southern EarthCARE overpass leg.
10/31	5:28	Over-land flight to ARM SGP. Flew over several NWS radiosonde sites and NSSL. G-III coordinated for the flight legs in Arizona.
11/04	3:54	Over-ocean targeting clear-sky AWS overpass at 19:20 UTC with over-land flight legs of Lake Tahoe and Railroad Valley. Fully coordinated with G-III.
11/07	4:44	Over-ocean targeting NOAA21 overpass at 21:25 UTC. Mostly clear sky except far southern parts of overpass legs. Fully coordinated with G-III.
11/12	3:35	Over-ocean out and back flight targeting slight pressure gradient and clear/cloudy transitions. No G-III coordination.
11/13	3:00	Over-ocean targeting an atmospheric river and pressure gradient with G-III coordination. ER-2 had to RTB early and did not go over Catalina Island like in the flight plan.
<b>ER-2 only ocean</b>		<b>ER-2 only land</b>
<b>ER2-/G-III coordinated</b>		

## Atmospheric River: 13 Nov

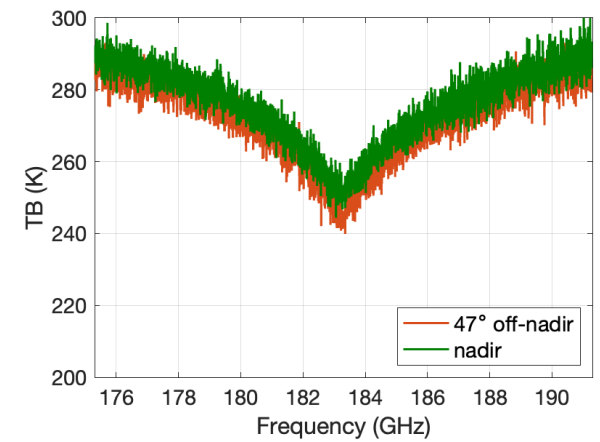
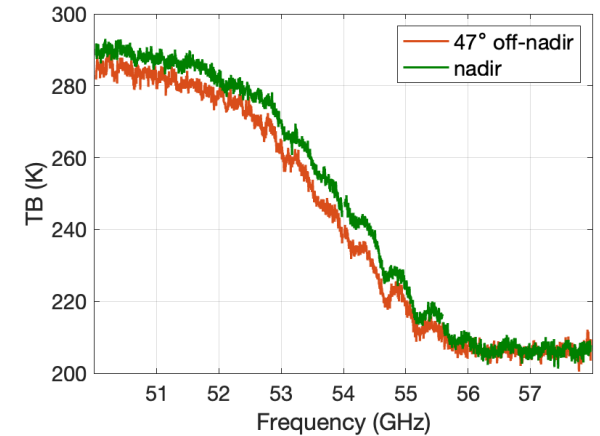
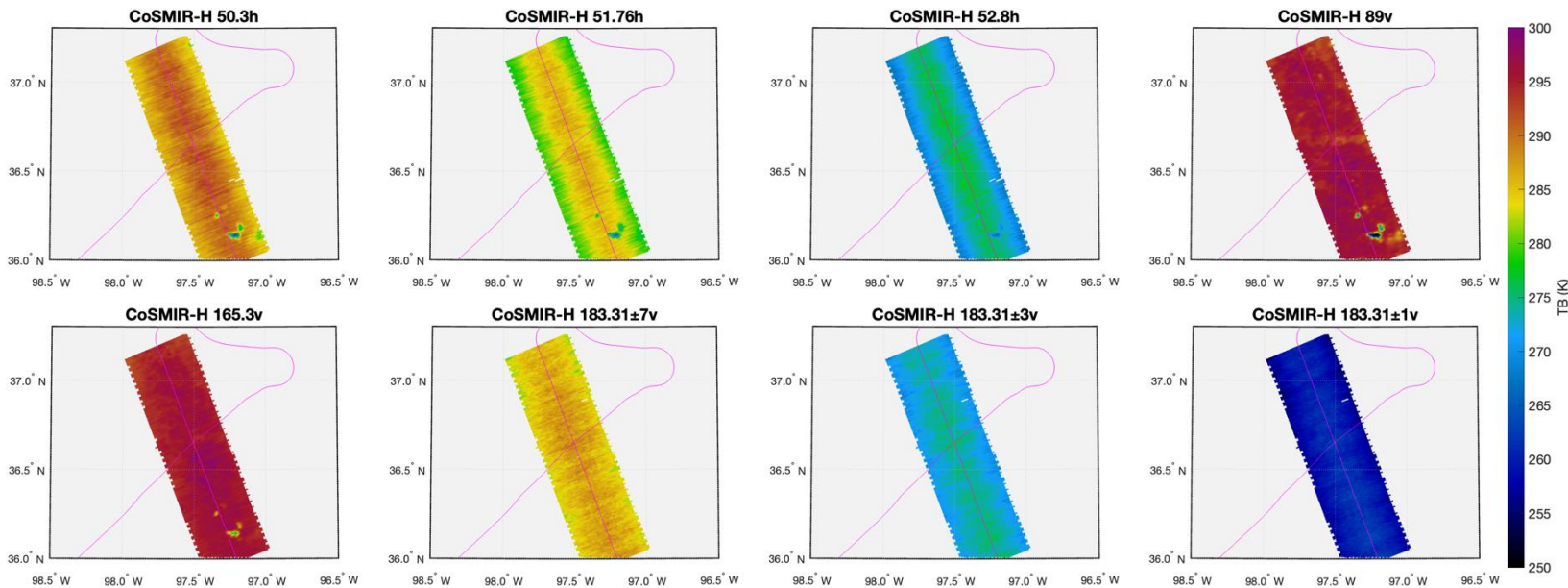


CoSMIR-H selected channels for the 11/13 atmospheric river flight. Precipitation can be seen in the 50/89 GHz channels and ice scattering in the 183 GHz channels on the north side of the leg.



50-GHz and 183-GHz channels are created by averaging the 4-MHz bins into ATMS-equivalent bandwidths

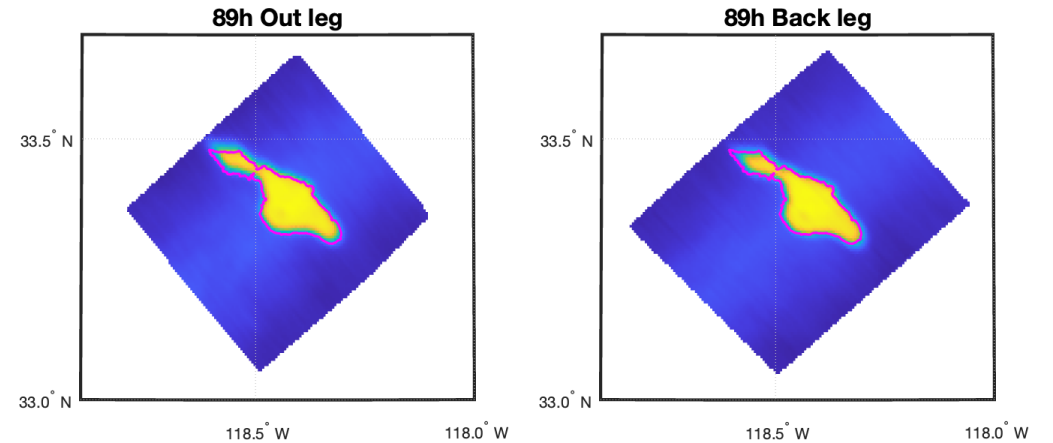
## Clear sky over ARM SGP site (Oklahoma, USA): 22 Oct



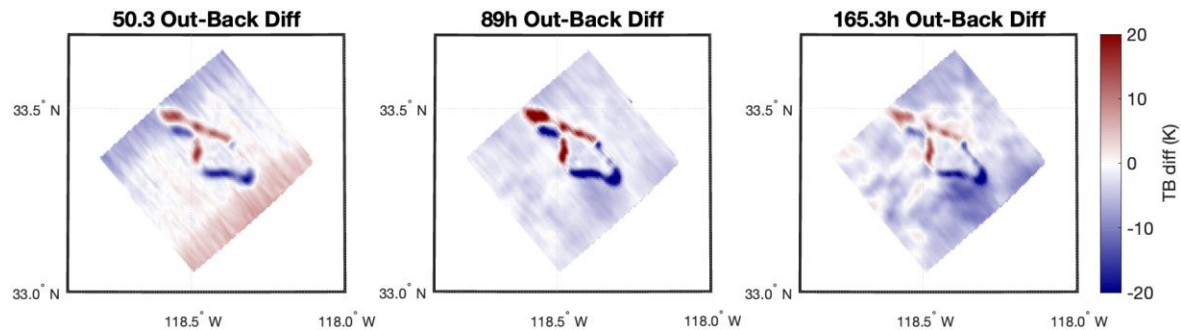
(Top) Selected band-averaged (ATMS-equivalent) channel observations. SGP central facility site where the pink lines intersect (aircraft ground track).

(Right) TB as a function of frequency at SGP overpass

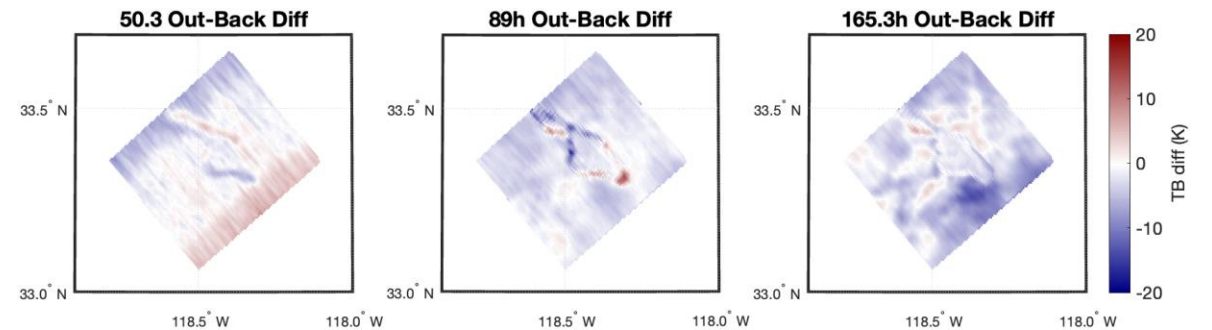
- 12 Nov out and back flight over Catalina Island provides excellent opportunity for geolocation checks
- Derive a pitch/roll offset using the coastlines from the out and back legs
  - Similar to doing ascending-descending differences for satellite geolocation



**Before correction**

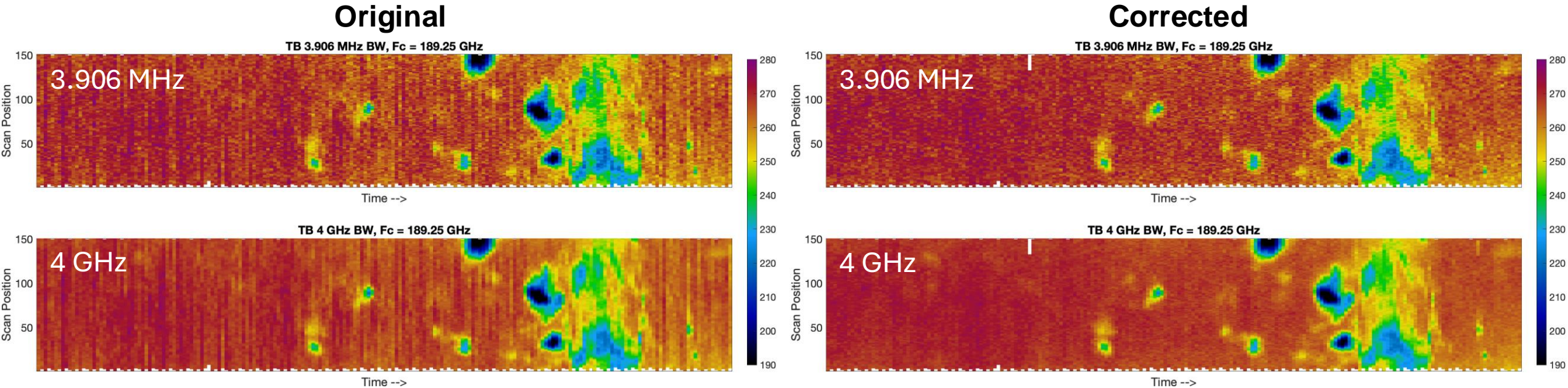


**After correction**



0° pitch, 1.6° roll offsets

- Significant striping (i.e. flicker or  $1/f$  noise) was noticed in the 183-GHz hyperspectral channels
- Correction was developed and applied for the first data release, but we are continuing to investigate other methods to remove the noise
- Conducting laboratory measurements to better characterize and mitigate the noise for future measurements

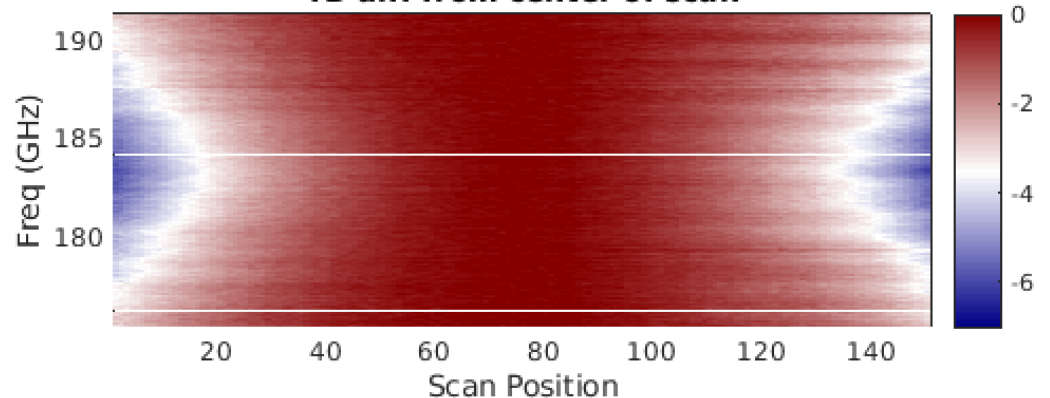


1/f noise correction significantly reduces the TB striping while maintaining the fine geophysical features.

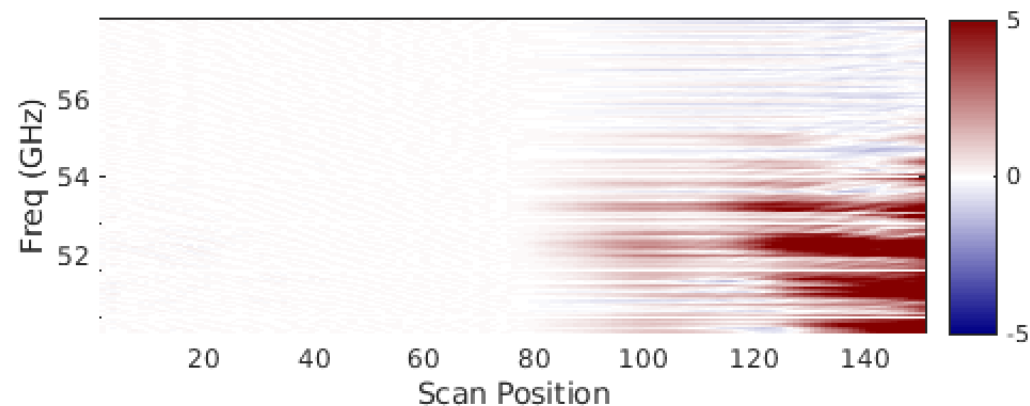
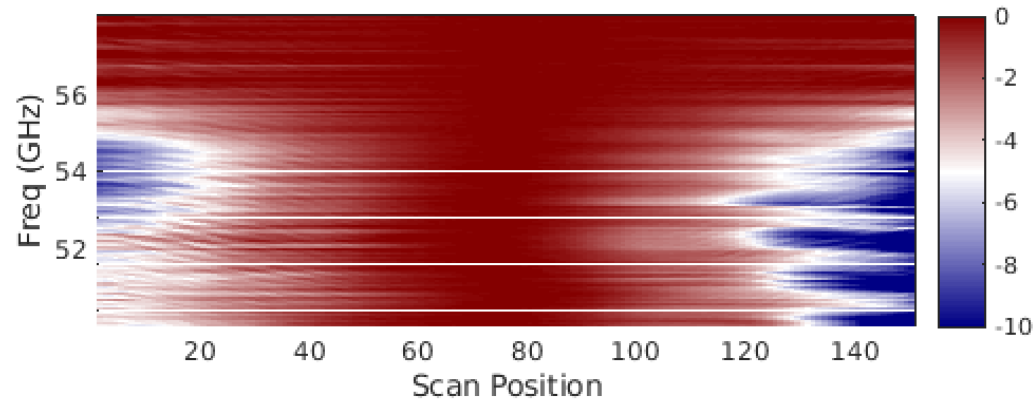
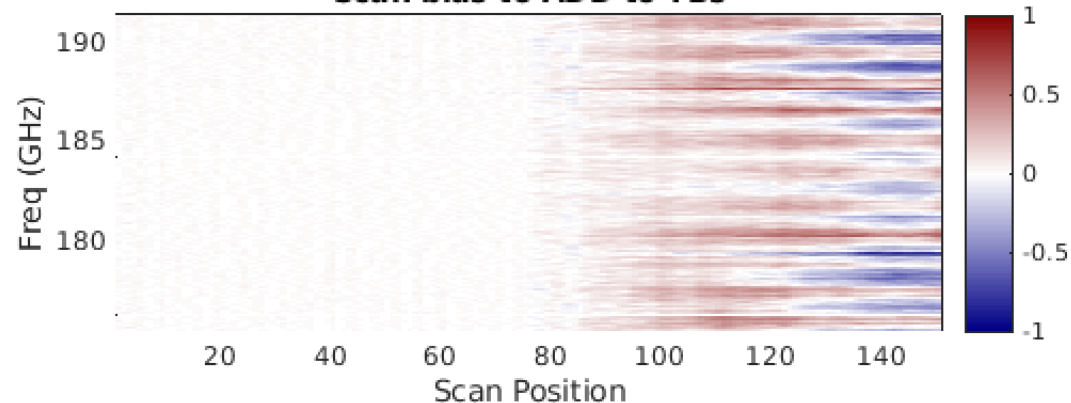
Some significant scan biases were noticed in the 50-GHz channels on the right side of the scan

Used difference with the left side of the scan to calculate a bias for the right side.

**TB diff from center of scan**

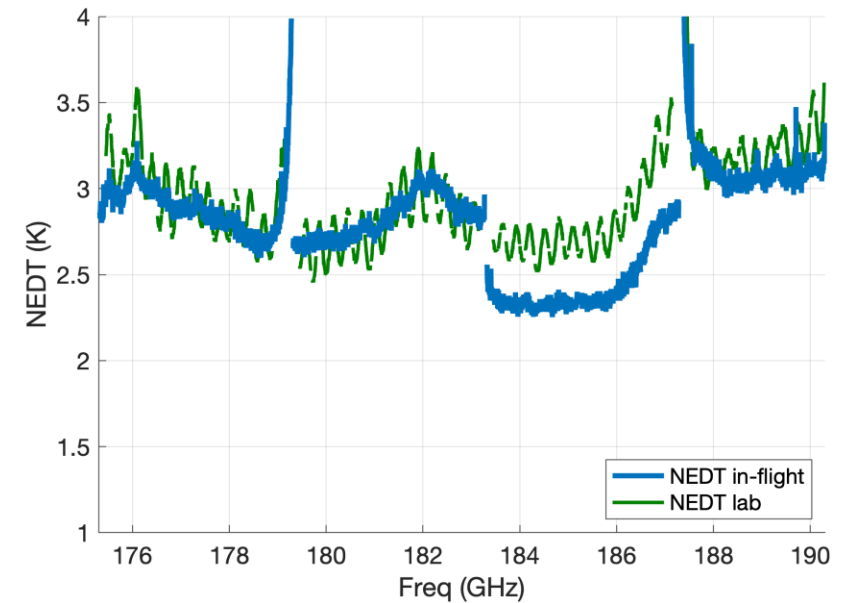
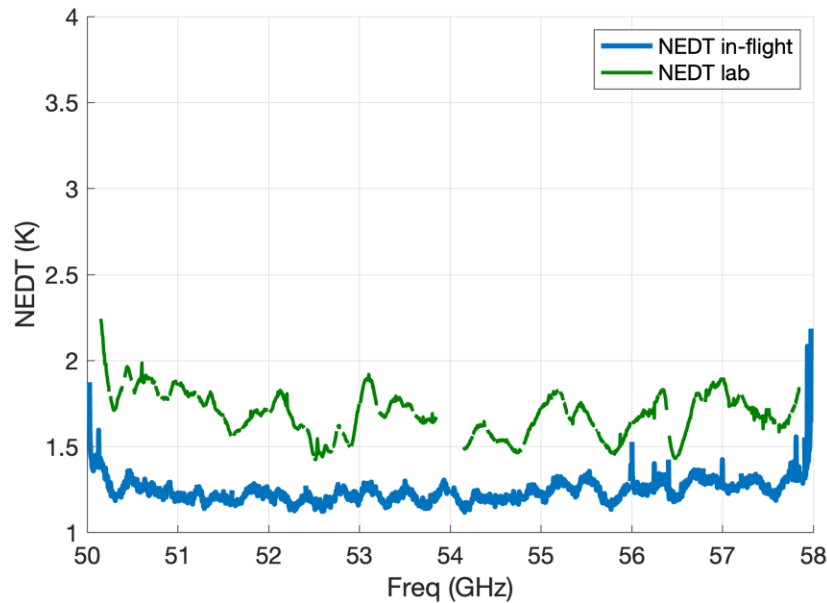


**Scan bias to ADD to TBs**



- Laboratory measurements of the receiver noise temperature ( $T_R$ ) as a function of frequency were made prior to WH<sup>2</sup>yMSIE using the onboard hot target and an LN<sub>2</sub> target
- NEDT is estimated from these measurements and is compared with in-flight NEDT calculations for 70 ms integration time ( $\tau$ ) and 4 MHz bandwidth (B)
- The in-flight NEDT shows very good results within what is expected. Work is currently being done to calculate the in-flight NEDT when the 4-MHz bins are averaged into wider bandwidths

$$NEDT = \frac{T_A + T_R}{\sqrt{B\tau}}$$

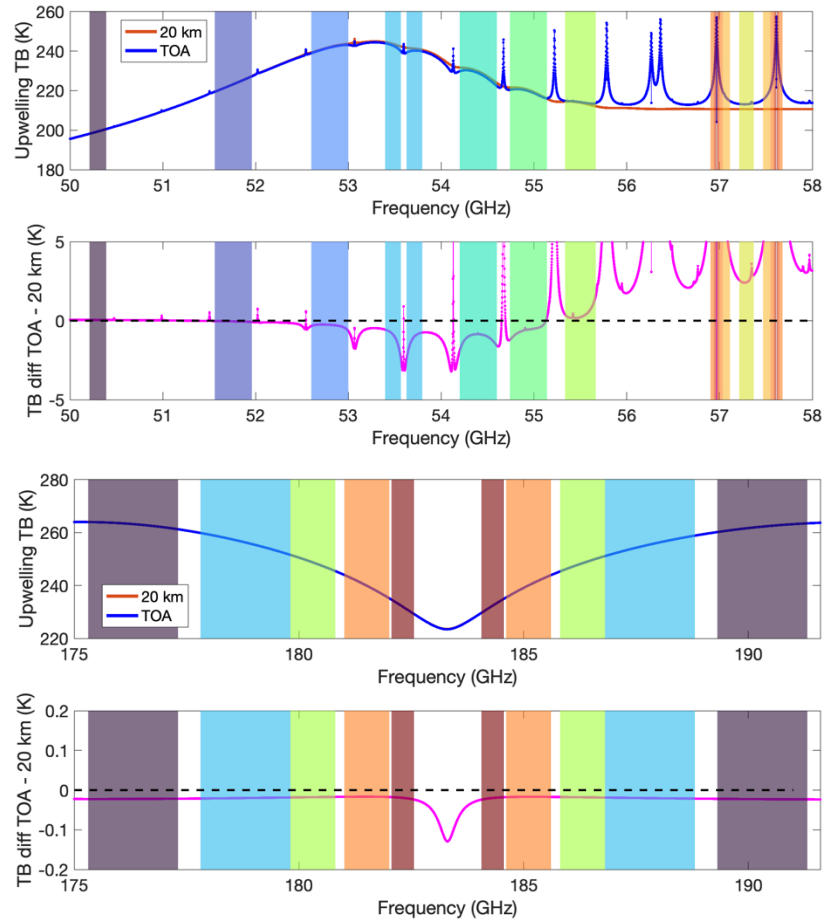




ATMS Freq (GHz)	ATMS Bandwidth (MHz)	S-NPP ATMS Measured NEDT (K)	CoSMIR-H Expected NEDT* (K)
50.3	180	0.37	TBD
51.76	400	0.28	TBD
52.8	400	0.28	TBD
53.596±0.115	170	0.29	TBD
54.4	400	0.27	TBD
54.94	400	0.27	TBD
55.5	330	0.29	TBD
57.290.344 [f <sub>0</sub> ]	155	0.43	TBD
f <sub>0</sub> ±0.217	78	0.56	TBD
f <sub>0</sub> ±0.322±0.048	36	0.59	TBD
f <sub>0</sub> ±0.322±0.022	16	0.86	TBD
f <sub>0</sub> ±0.322±0.010	8	1.23	TBD
f <sub>0</sub> ±0.322±0.0045	3	1.95	TBD
183.31±7.0	2000	0.38	TBD
183.31±4.5	2000	0.46	TBD
183.31±3.0	1000	0.54	TBD
183.31±1.8	1000	0.59	TBD
183.31±1.0	500	0.73	TBD

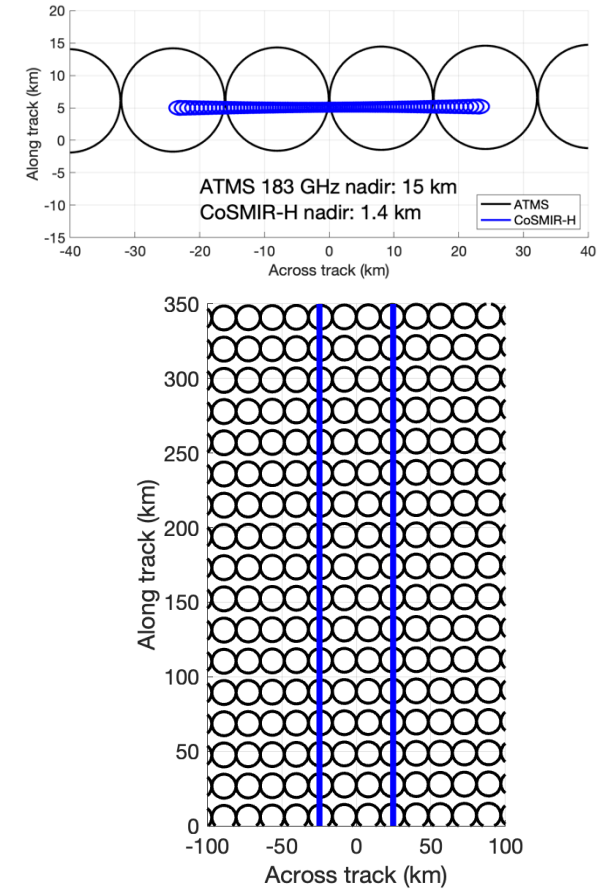
\*Currently working on calculations

## Upwelling TB comparison: TOA vs 20 km



Stratospheric O<sub>2</sub> sounding channels show significant differences from ATMS but the differences in the H<sub>2</sub>O sounding channels should be negligible.

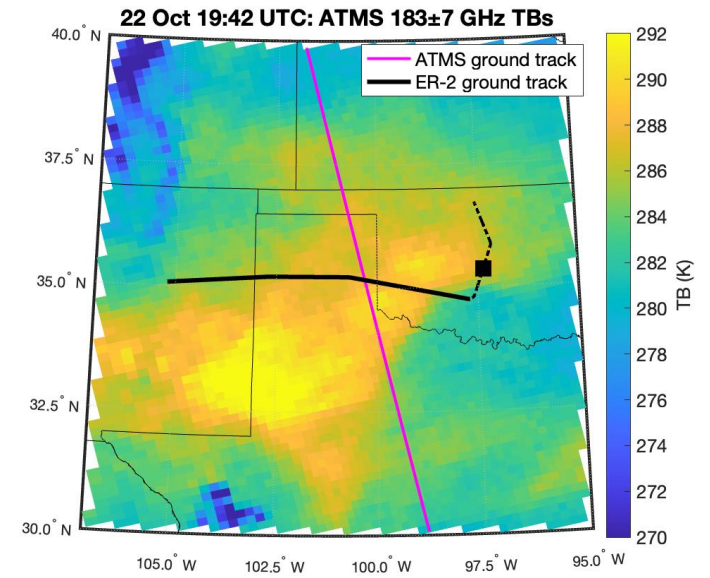
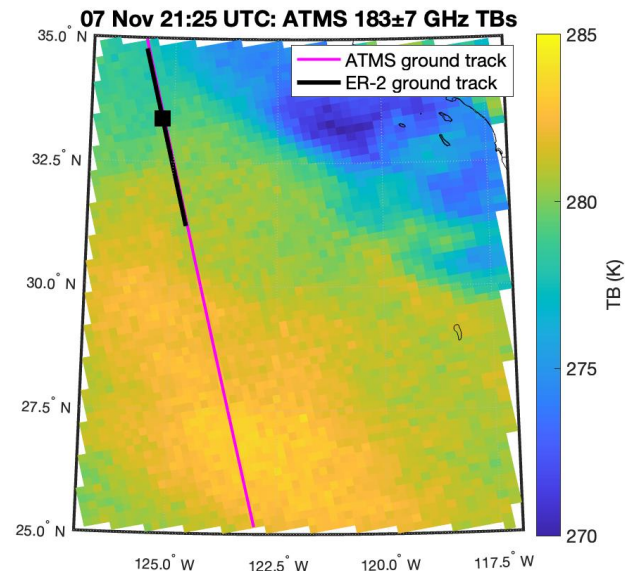
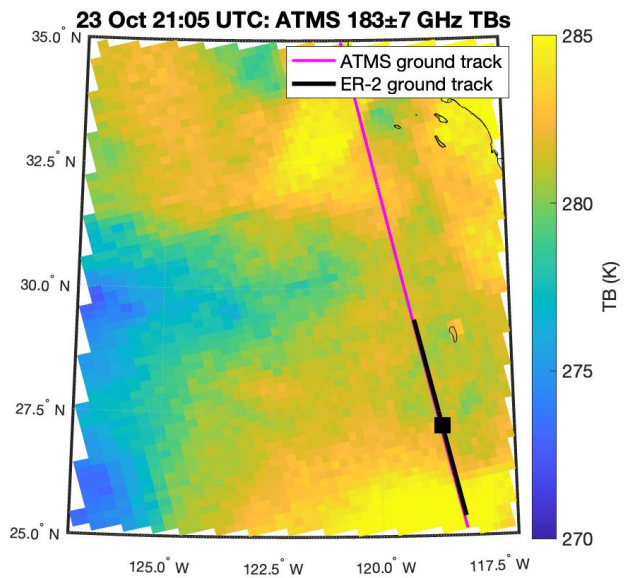
## Footprint/swath comparison

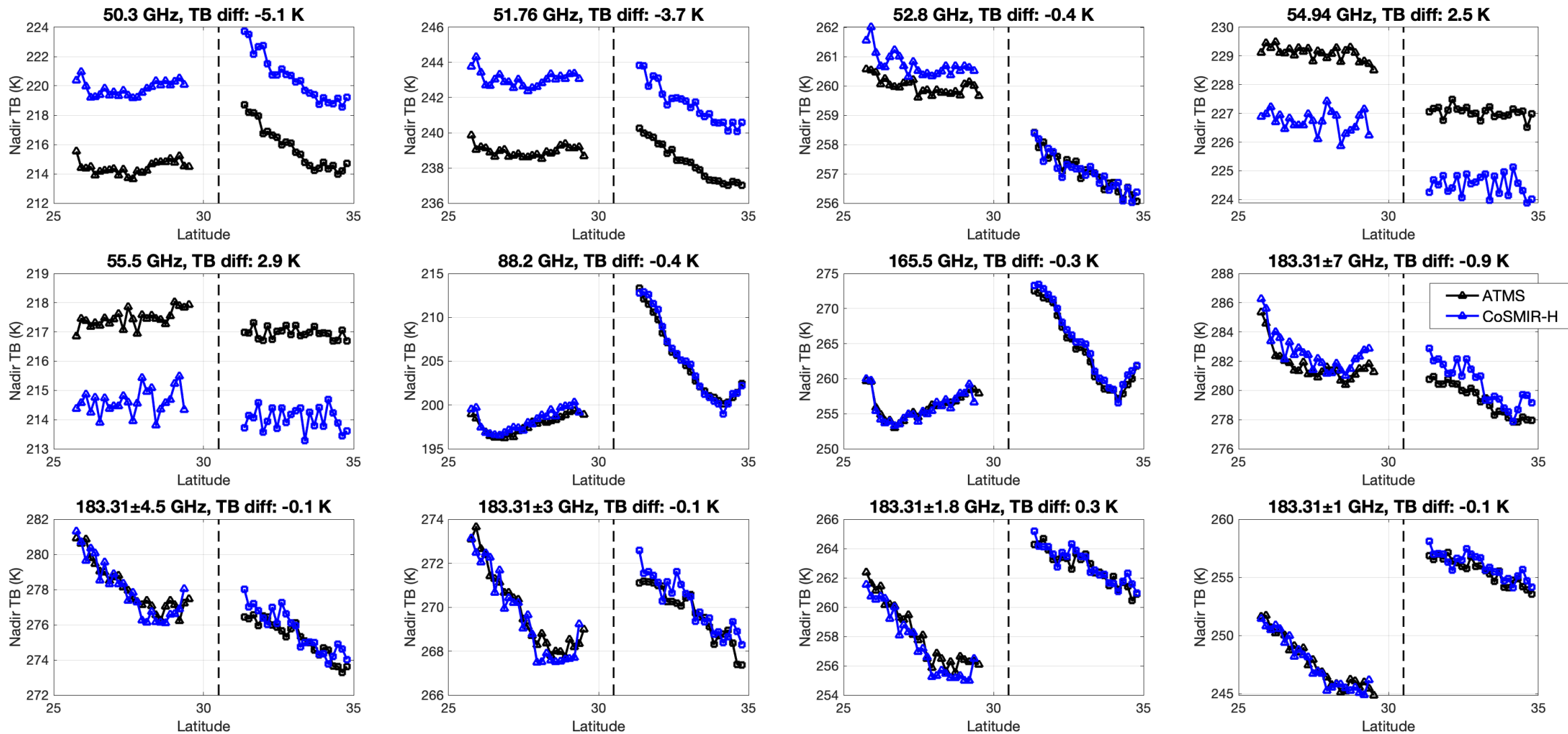


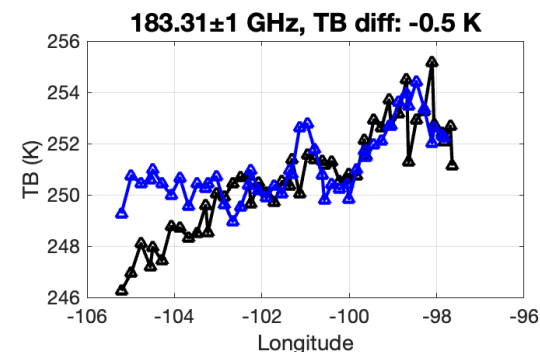
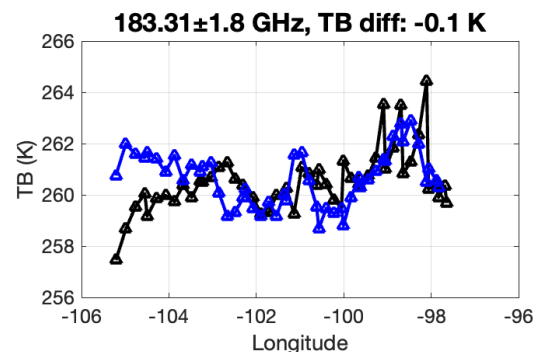
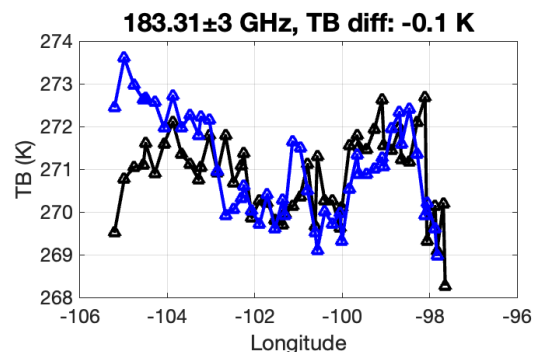
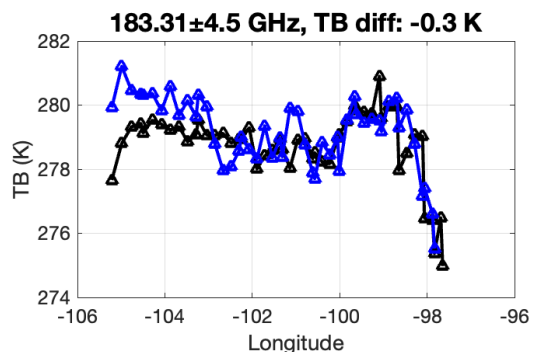
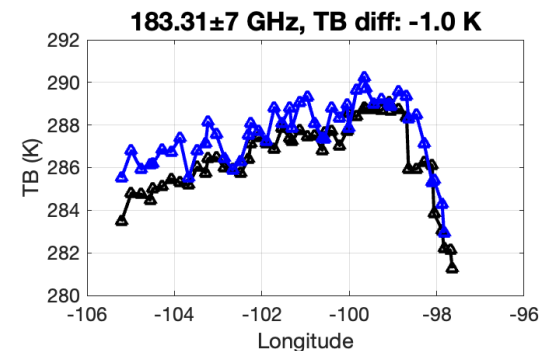
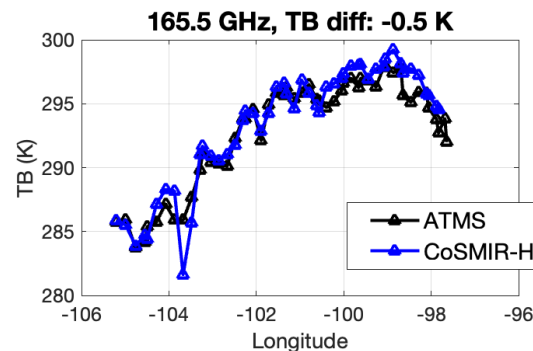
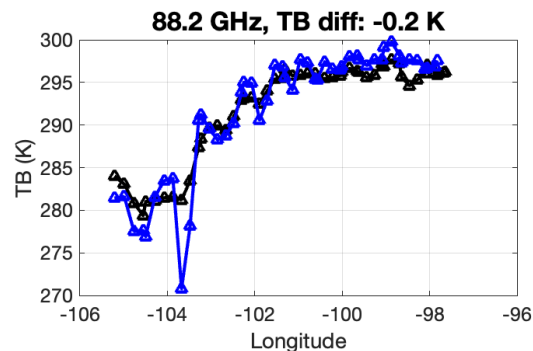
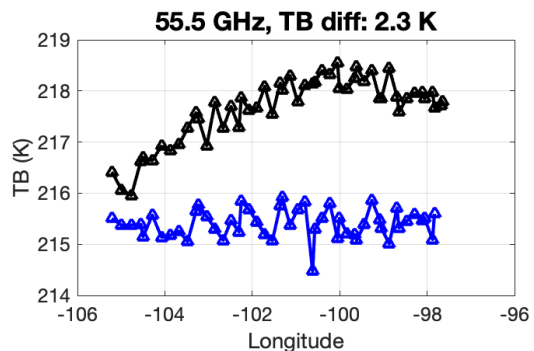
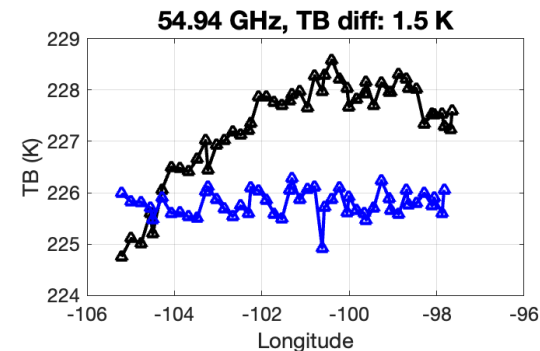
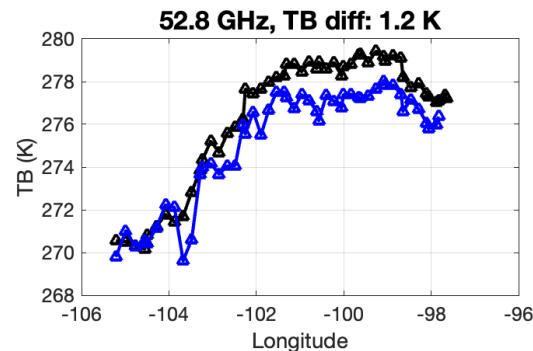
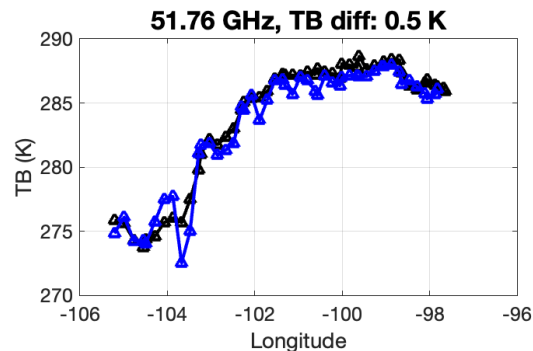
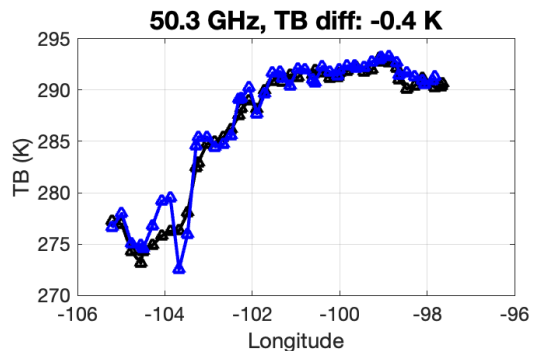
ATMS underflights over homogeneous scenes to compare observations

- Two primary underflights
  - Ocean clear-sky
  - ~30-minute aircraft flight track along NOAA21 satellite track
- One ‘underflight of opportunity’
  - NOAA21 intersected over-land flight track ~30 min after the ER-2
- CoSMIR-H 4-MHz frequency bins averaged to approximate ATMS channel bands using NOAA21 spectral response functions

Comparisons on the next two slides show very good agreement with ATMS







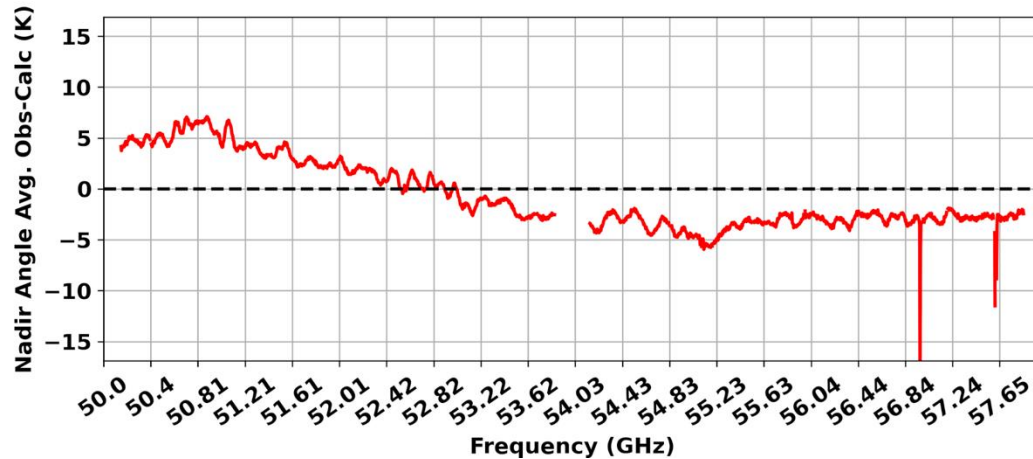
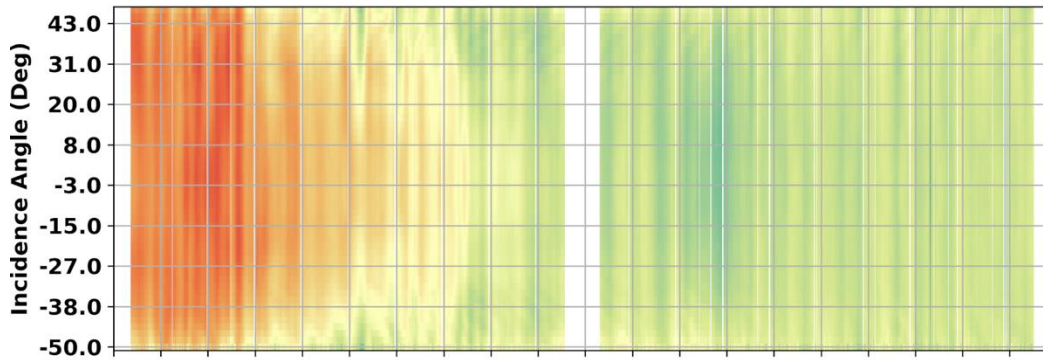
Over-ocean clear-sky CoSMIR-H observed TB spectrum (obs) compared with simulated TB from the Community Radiative Transfer Model CRTM (calc)

Obs (CoSMIR-H) - Calc (CRTM\_HRRR) from 1 segments

2048 Channels (50.0 GHz - 58.0 GHz)

GEOS-FP Cld Frac Max: 0% | Horiz. Avg: None | SDev Filter: 1.0 sigma

2024/10/23 19:36 - 2024/10/23 20:06 UTC

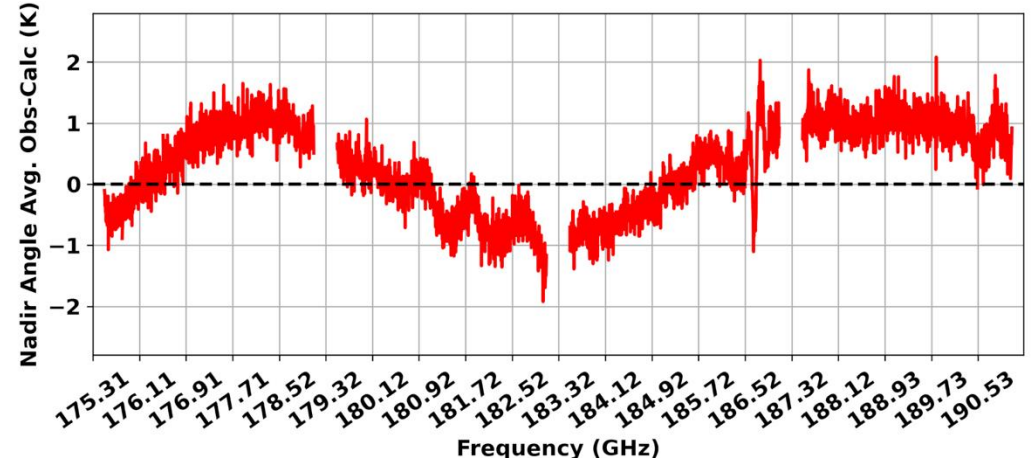
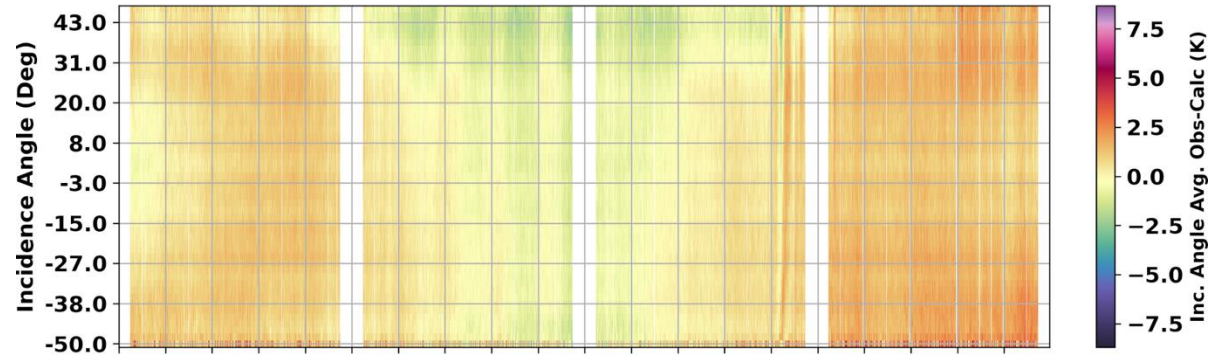


Obs (CoSMIR-H) - Calc (CRTM\_HRRR) from 1 segments

4096 Channels (175.31 GHz - 191.31 GHz)

GEOS-FP Cld Frac Max: 0% | Horiz. Avg: None | SDev Filter: 1.0 sigma

2024/10/23 19:36 - 2024/10/23 20:06 UTC



- L1 data files contain full spectrum: 6148 channels
- Example file name: whymsie-cosmirh\_ER2\_20241022\_S172056\_E175100\_R0\_crosstrack.nc
- Files are identified with the flight date and the start (S) and end (E) time in 30-min segments
- Each file ~900 MB

## Variable names included in the Level 1 data files

UTC Time	Aircraft Nav Data	Sensor Specific
Year	AC_Altitude	Tb
Month	AC_Pitch	Quality
DayOfMonth	AC_Roll	IncidenceAngle
Hour	AC_Heading	Latitude
Minute	AC_Latitude	Longitude
Second	AC_Longitude	Azimuth
MilliSecond	AC_Speed	Elevation
	AC_SolarZenith	SensorPitch
		SensorRoll

- The airborne hyperspectral microwave sounder CoSMIR-H performed well during flights and collected a quality dataset
- Level 1 calibrated brightness temperature dataset is available for download
- Thanks to NASA ESTO and NOAA's Joint Venture Partnerships program for funding this exciting project

Data download: <https://www-air.larc.nasa.gov/cgi-bin/ArcView/whymsie>

Contact: [rachael.a.kroodsma@nasa.gov](mailto:rachael.a.kroodsma@nasa.gov)